

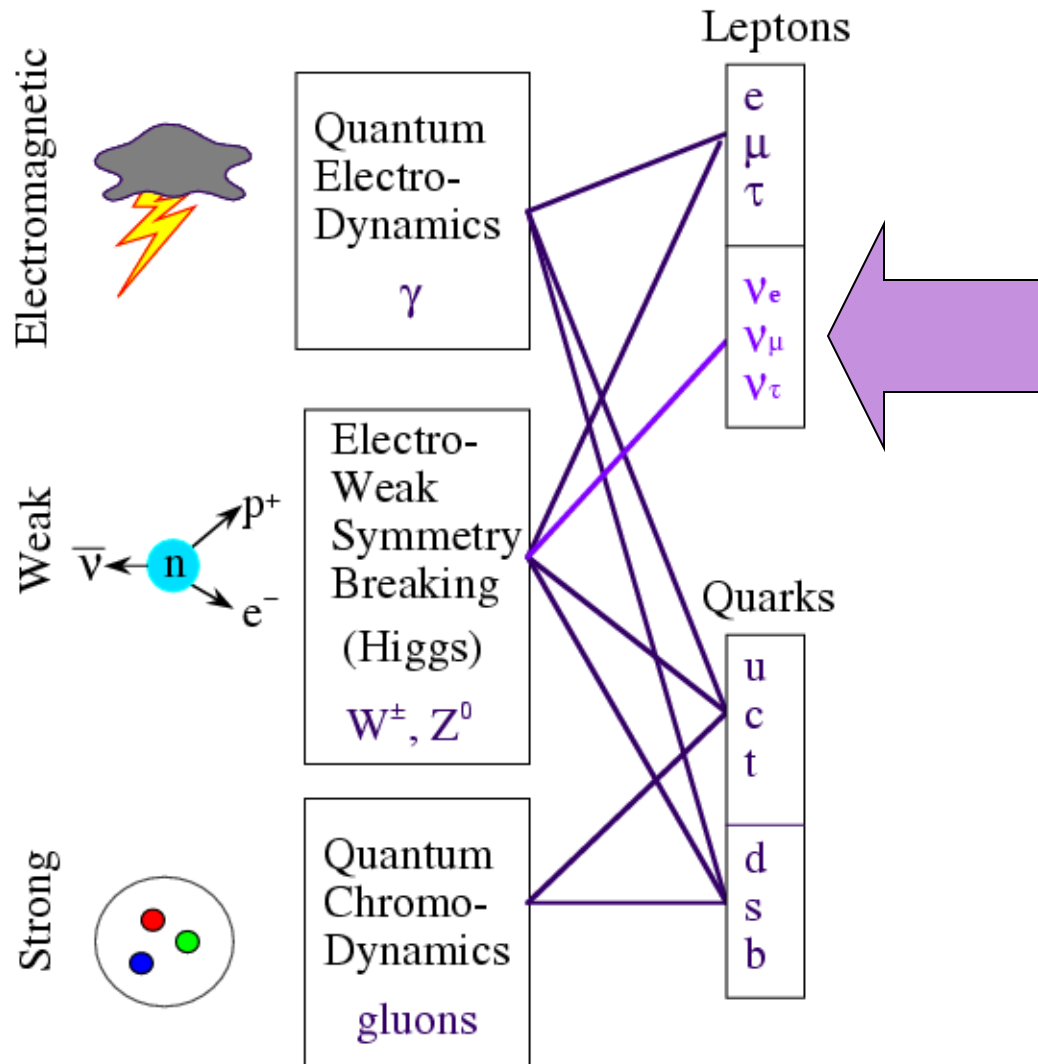
ν_e or not ν_e
...that is the Question!

Janet Conrad
MIT

*Three mysteries about observing
(or not observing)
electron neutrino interactions...*

- Chapter 1: Setting the stage...
- Chapter 2: Presenting the mysteries...
- Chapter 3: Speculating on how the plot may unfold
 (with a *new and improved*
 potential ending!)

The Standard Model



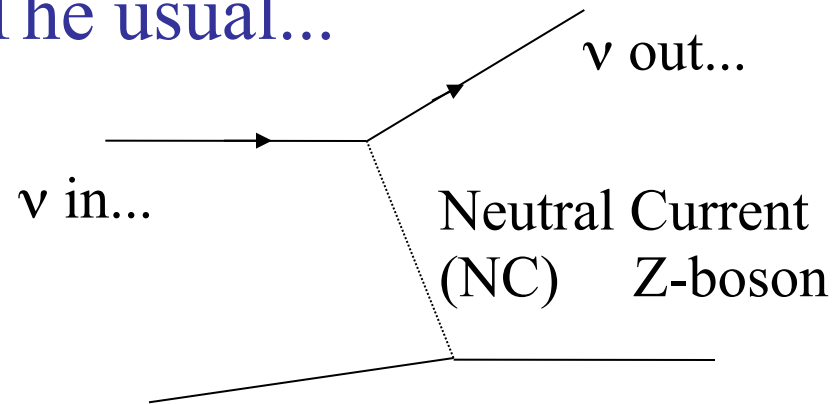
We've known about neutrinos since the 1930's

The electron flavor was the first predicted and first observed!

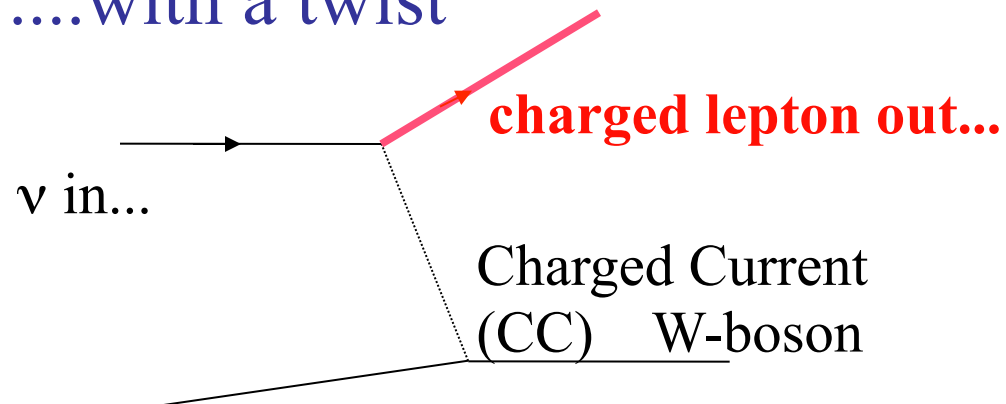
Neutrinos
interact via

*Only the
Weak Interaction*

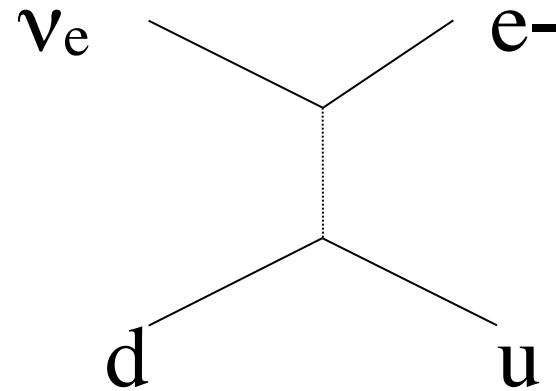
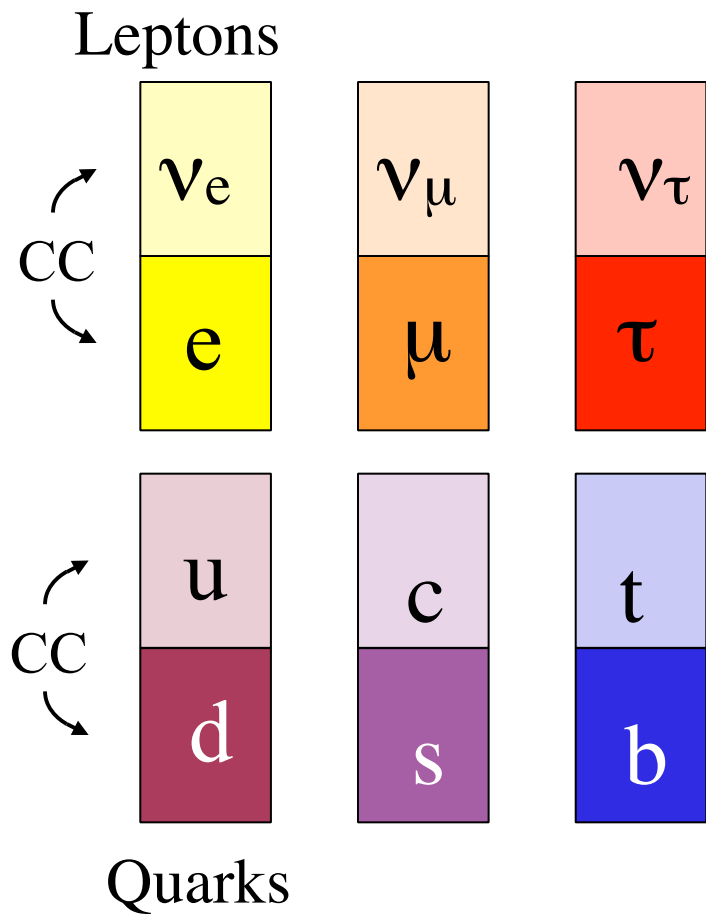
The usual...



...with a twist



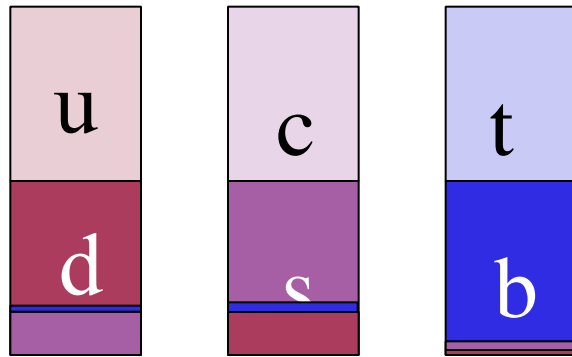
There are 3 types,
which form “weak doublets” with the charged leptons



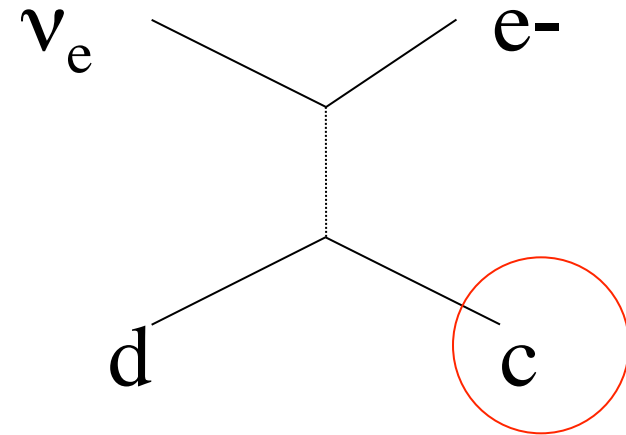
We identify the
neutrino flavor
via the CC interaction

Actually, with a modification in the quark sector:

MIXING: quark mass eigenstates \neq quark weak eigenstates



Small effect,
but clearly
seen in weak
interactions...



$$\begin{bmatrix} d' \\ s' \\ b' \end{bmatrix} = \begin{bmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{bmatrix} \begin{bmatrix} d \\ s \\ b \end{bmatrix}$$

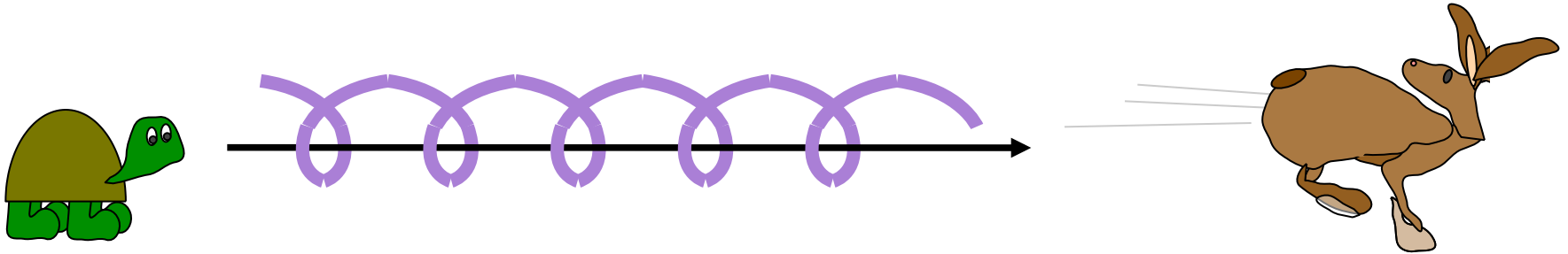
The CKM
Matrix

... and
kaon decays,
D meson decays,
etc.

Another feature: “handedness” ... but first, helicity

All spin 1/2 particles have “helicity”

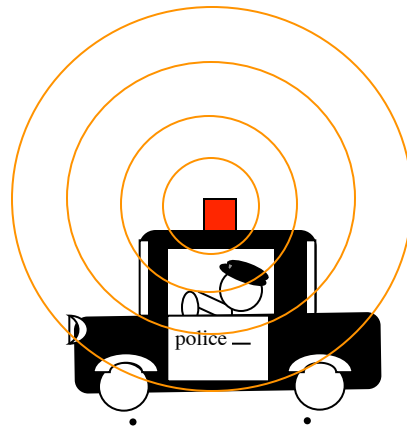
The operator: $\sigma \cdot \mathbf{p}$



Handedness (or chirality) is the Lorentz-invariant counterpart
Identical to helicity for massless particles (standard model ν 's)

Experimental observation of Parity Violation

Neutrinos are LH (and antineutrinos RH) ... always



How do you enforce the law of left-handedness?

Well... what couples left-handed particles to right?

A Dirac mass term
in the SM Lagrangian:

$$m(\bar{\nu}_L \nu_R + \bar{\nu}_R \nu_L)$$

If you want to build parity violation into “the law”
you want keep this term out of the Lagrangian...
a simple solution is: $m=0$

in the Standard Model, neutrinos are massless

The problem is...

apparently that's wrong!

It has long been known that neutrinos can,
in principle, oscillate...

If we postulate:

- Neutrinos have (different) masses
- The **Weak Eigenstate** is a mixture of **Mass Eigenstates**:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

Then a pure ν_μ beam at $t = 0$,
may evolve a ν_e component with time!

The Probability for Oscillations...

$$P_{osc} = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L/E)$$

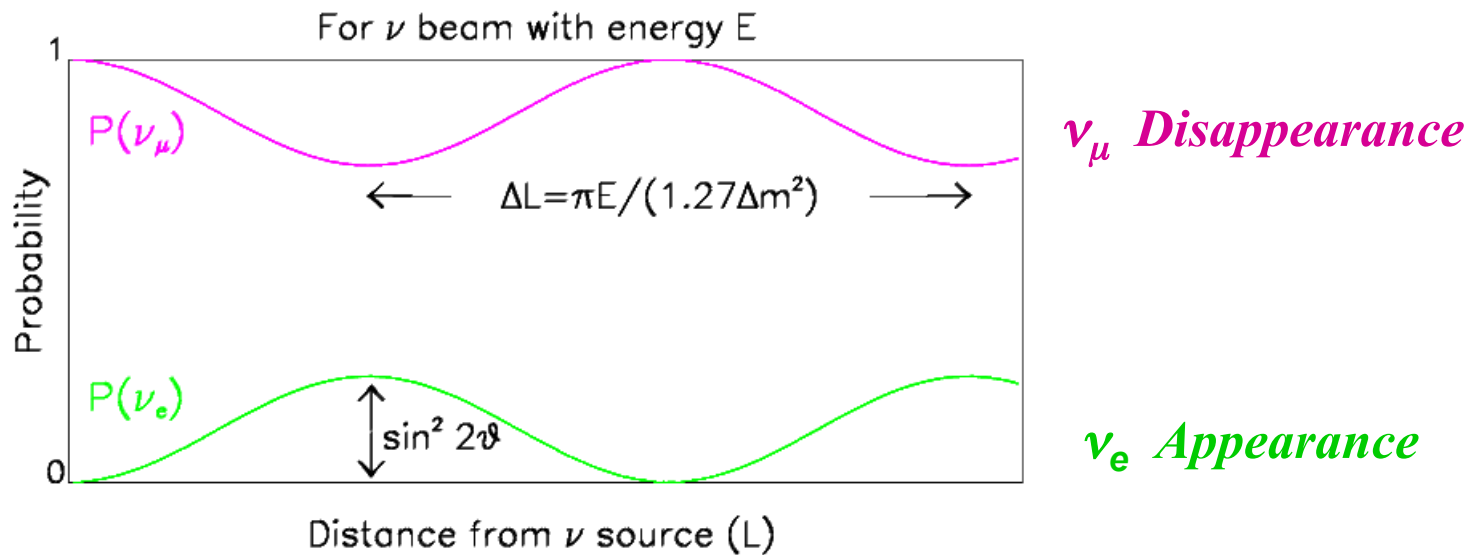
$$P_{Osc} = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 L / E \right)$$

...Depends Upon Two Experimental Parameters:

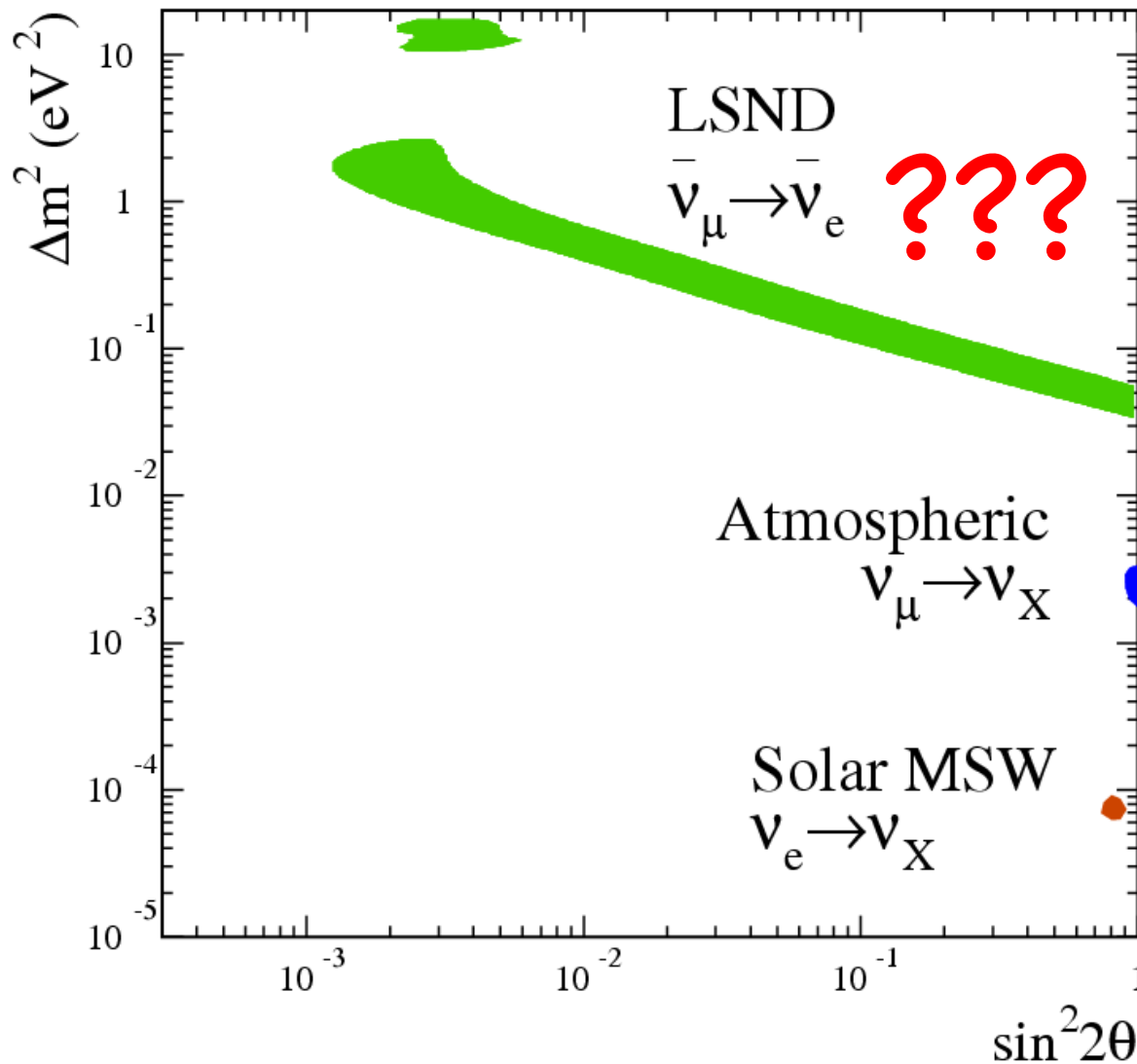
- L – The distance from the ν source to detector (km)
- E – The energy of the neutrinos (GeV)

...And Two Fundamental Parameters:

- $\Delta m^2 = m_1^2 - m_2^2$ (eV^2)
- $\sin^2 2\theta$



A mystery that I will only touch on... High Δm^2

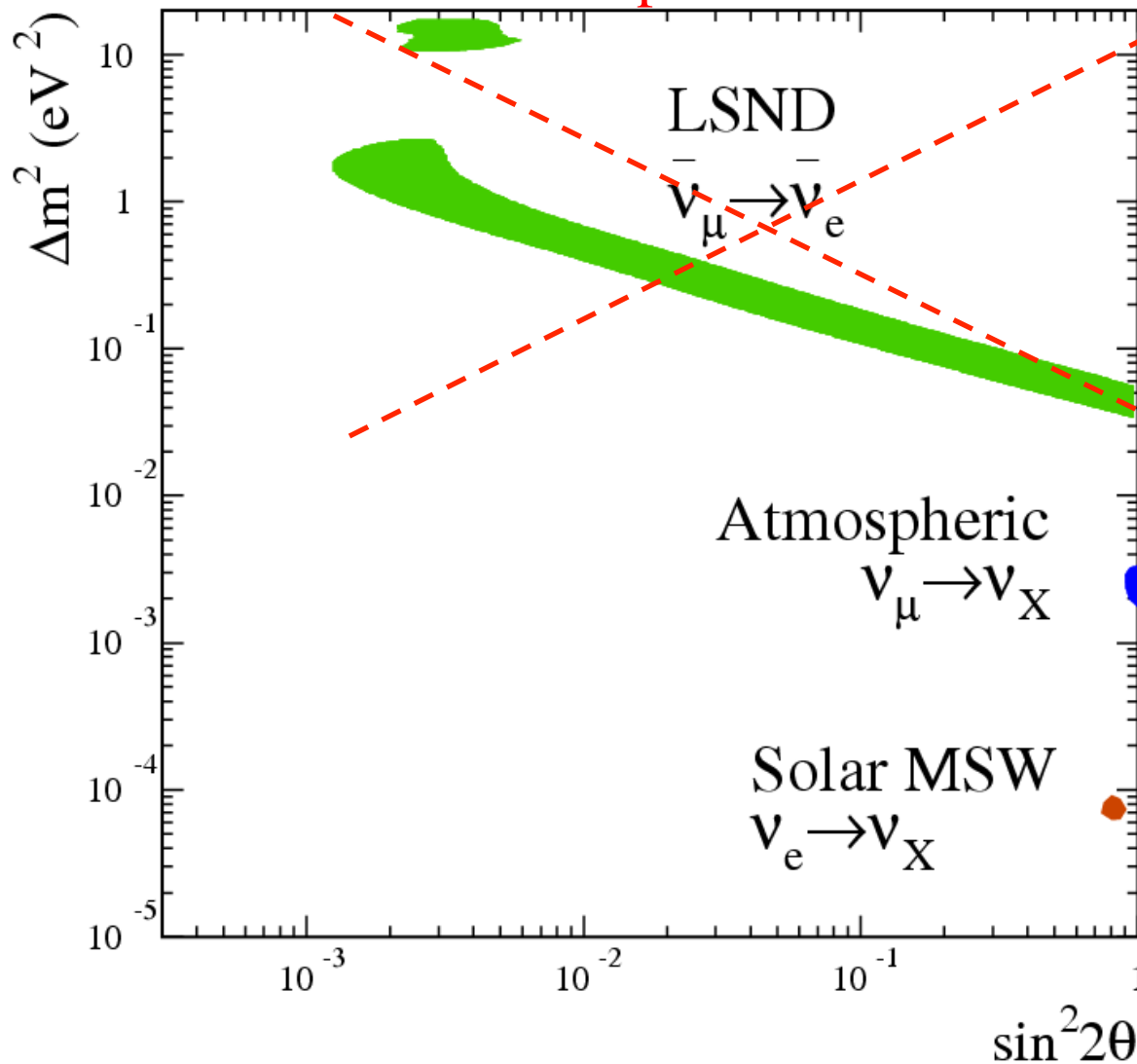


Ruled out by
MiniBooNE in ν -mode
But in $\bar{\nu}$ mode there may be
a signal??

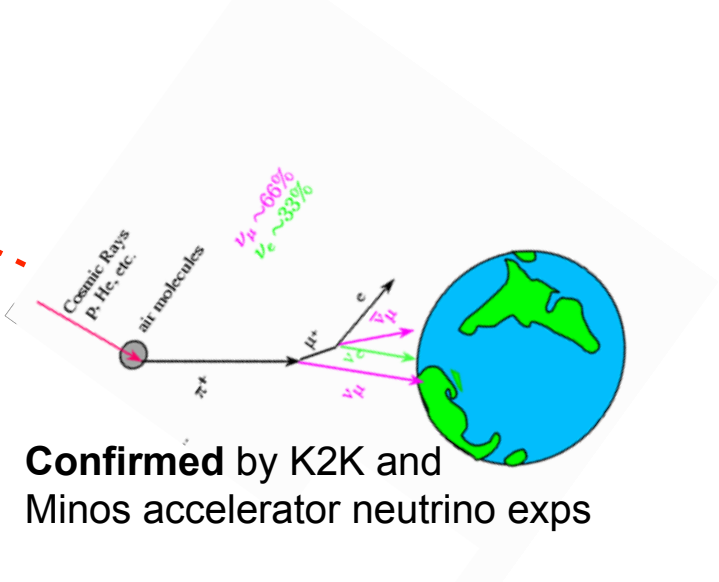
And a reanalysis of
 $\bar{\nu}$ production from reactors
may also indicate a
high Δm^2 signal at $>2\sigma$!

The community is
developing the strategy
for the next round of attack!

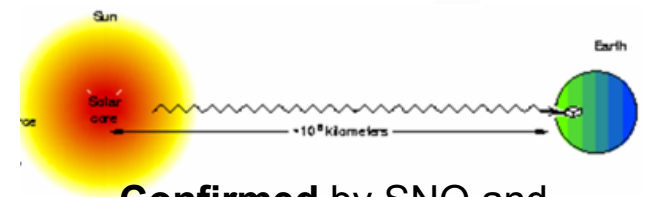
lets skip this for now...



Other oscillation signals are well-confirmed!

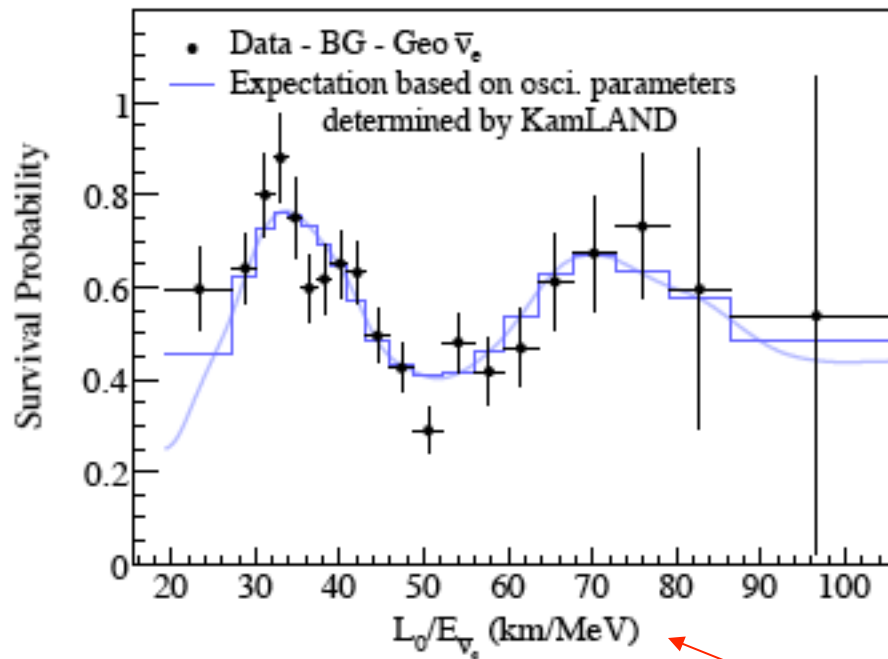


Confirmed by K2K and Minos accelerator neutrino exps



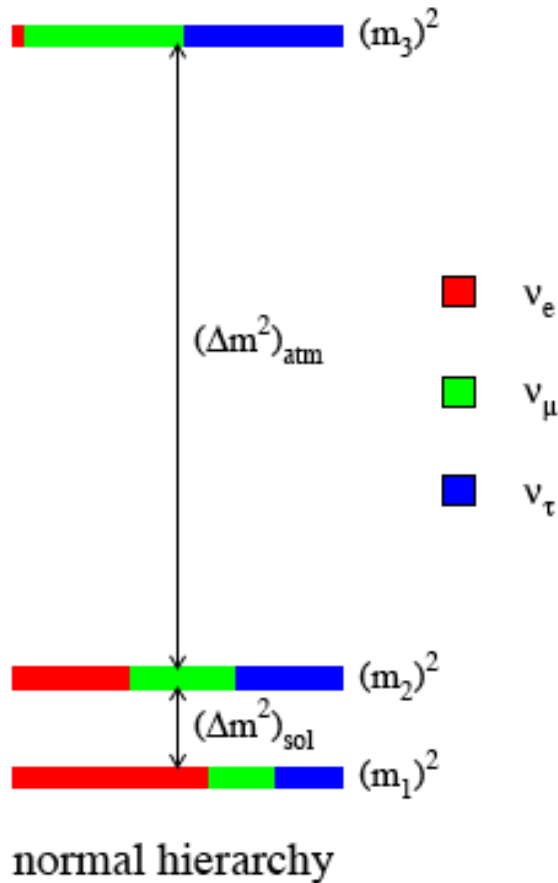
Confirmed by SNO and by Kamland reactor neutrino exp

The result from the Kamland reactor experiment also shows the L/E dependence one expects from oscillations!



$$P_{Osc} = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 L / E \right)$$

We have a fully self-consistent model
for how neutrinos behave...

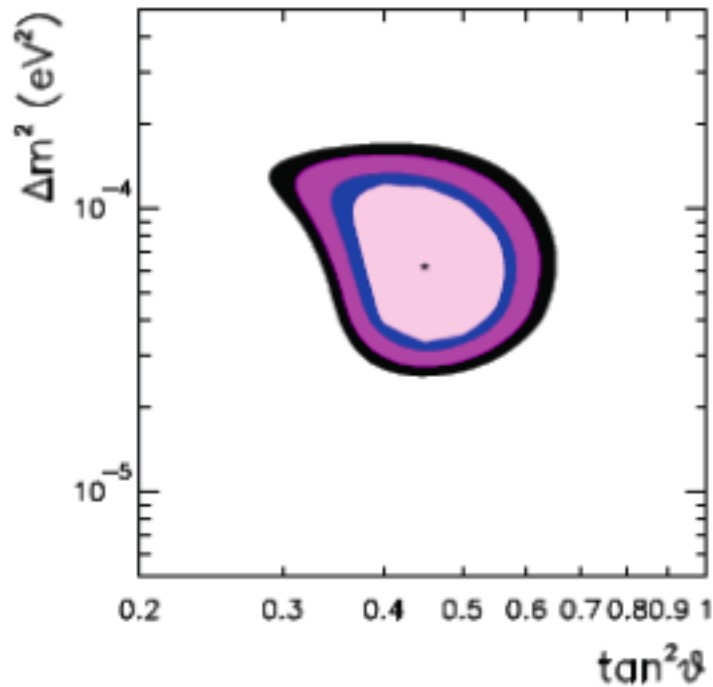


$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

“mixing” between neutrinos
is parameterized by
three “mixing angles”
 $\theta_{12}, \theta_{13}, \theta_{23}$

This model is predictive!

Allowed region for
solar neutrino oscillation
measurements,

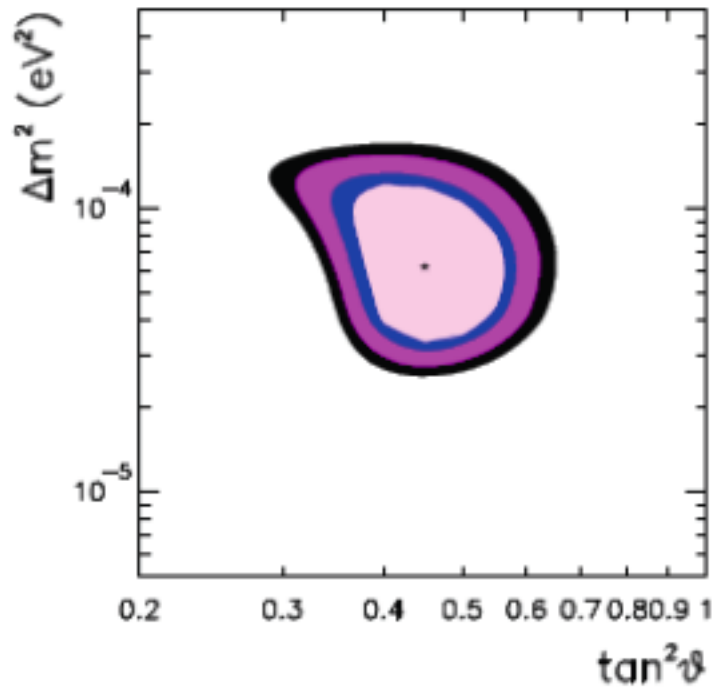


if this is due to $\nu_e \rightarrow \nu_{\text{other}}$

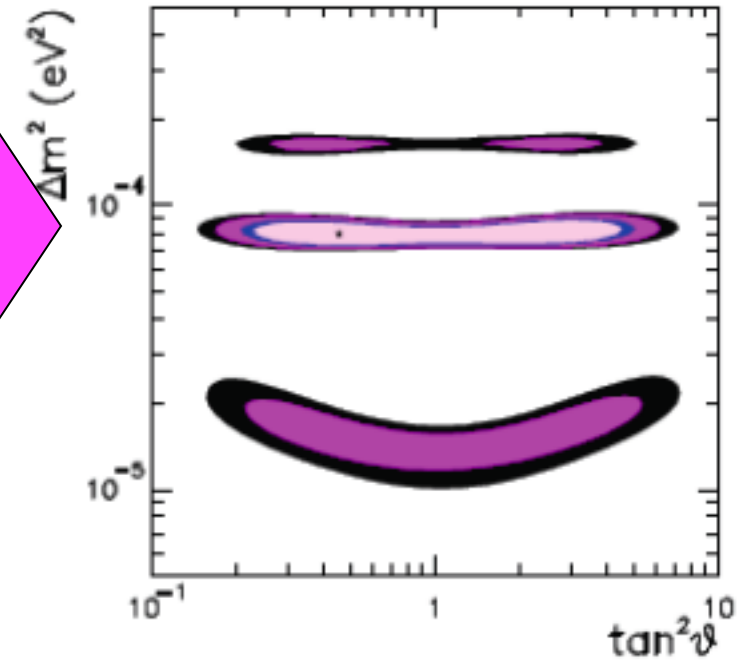
Then $\bar{\nu}_e \rightarrow \bar{\nu}_{\text{other}}$
should be observable
with the same wavelength

This model is predictive!

Allowed region for
solar neutrino oscillation
measurements,



Allowed region for the
Kamland reactor
 $\bar{\nu}_e \rightarrow \bar{\nu}_{\text{other}}$ Experiment!



fits by Gonzalez-Garcia, an old plot, but illustrative!

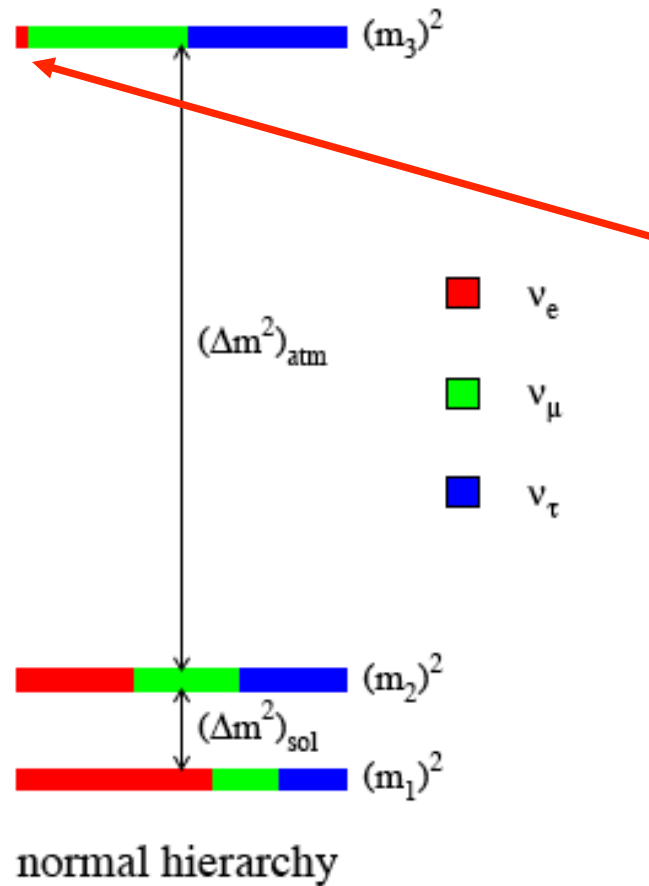


Three mysteries about the

ν_e

Our Model

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



Mystery 1:

What's happening here?

Is there any ν_e content
at all?

Writing that mixing matrix more explicitly...

$$c_{ij} = \cos\theta_{ij}$$

$$s_{ij} = \sin\theta_{ij}$$

This element is tiny or even zero!

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

$$= \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

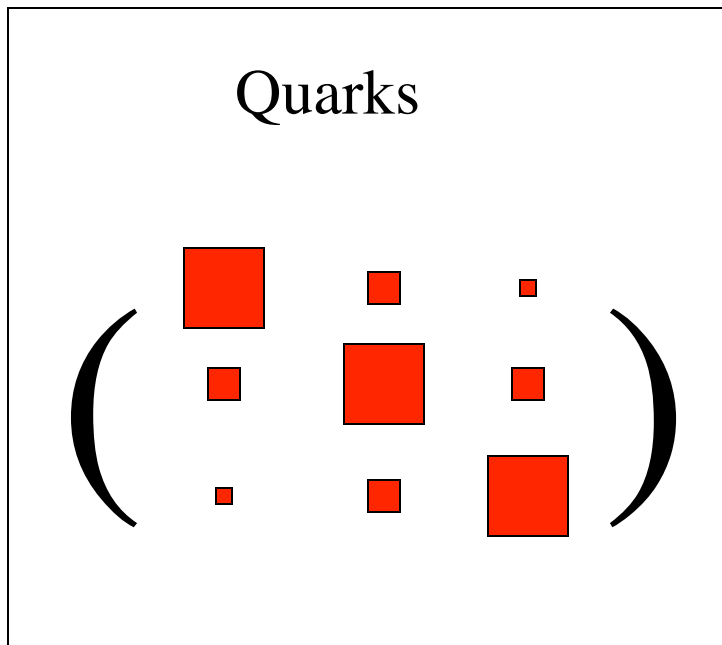
From Atmospheric and Long Baseline Disappearance Measurements

From Reactor Disappearance Measurements

From Appearance Measurements

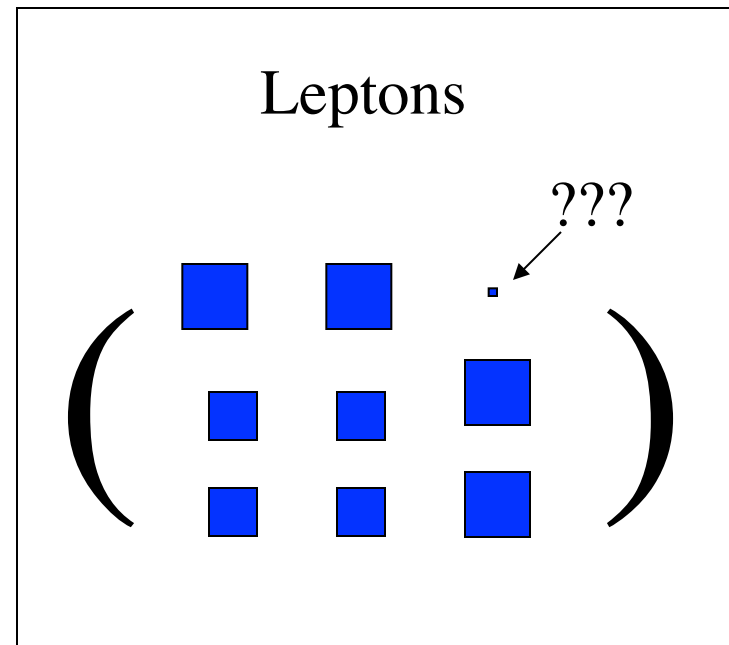
From Solar Neutrino Measurements

The lepton mixing matrix is NOT like the quark matrix!
(*WHY???*)



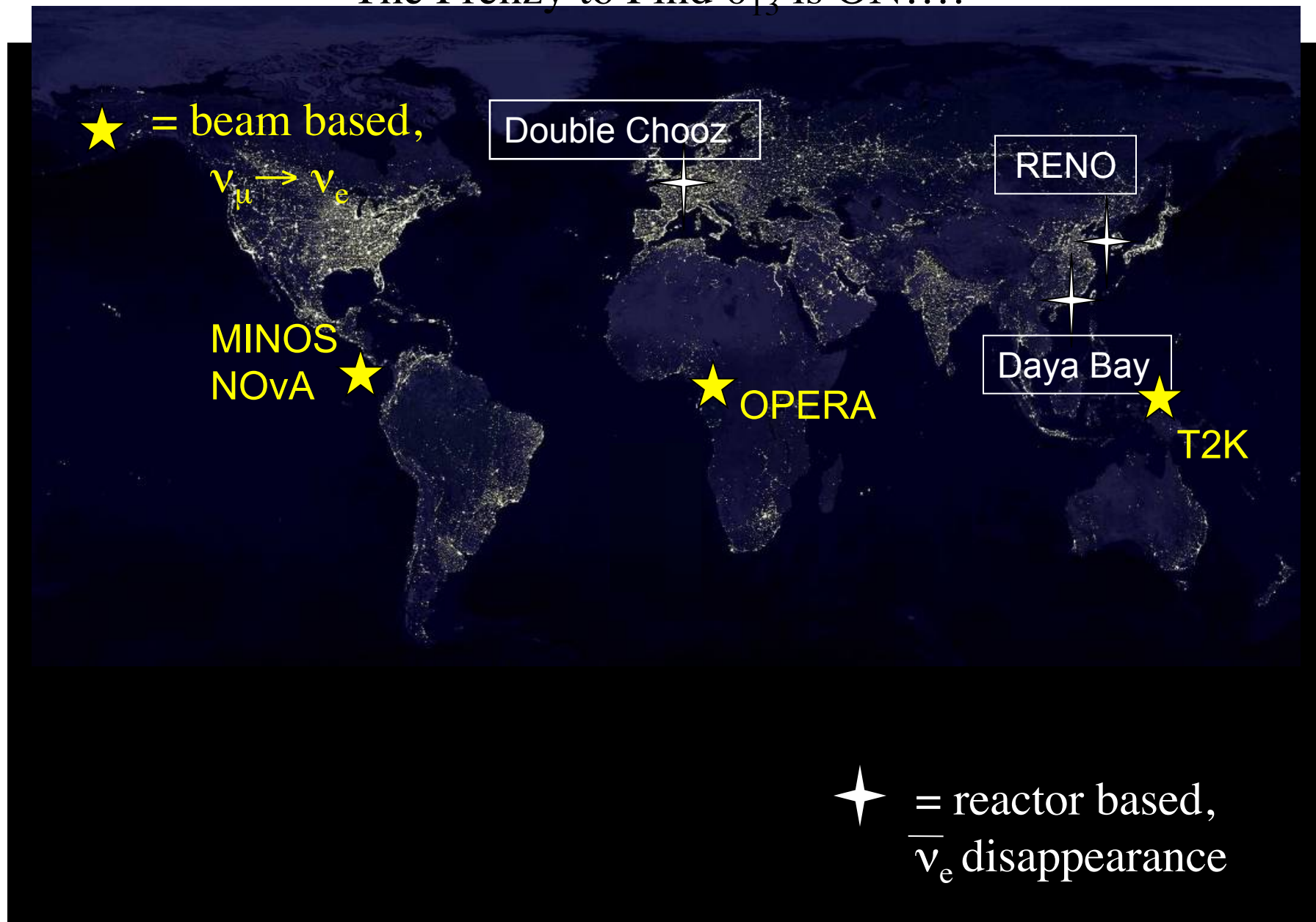
Large entries on diagonal
small off diagonal

vs.

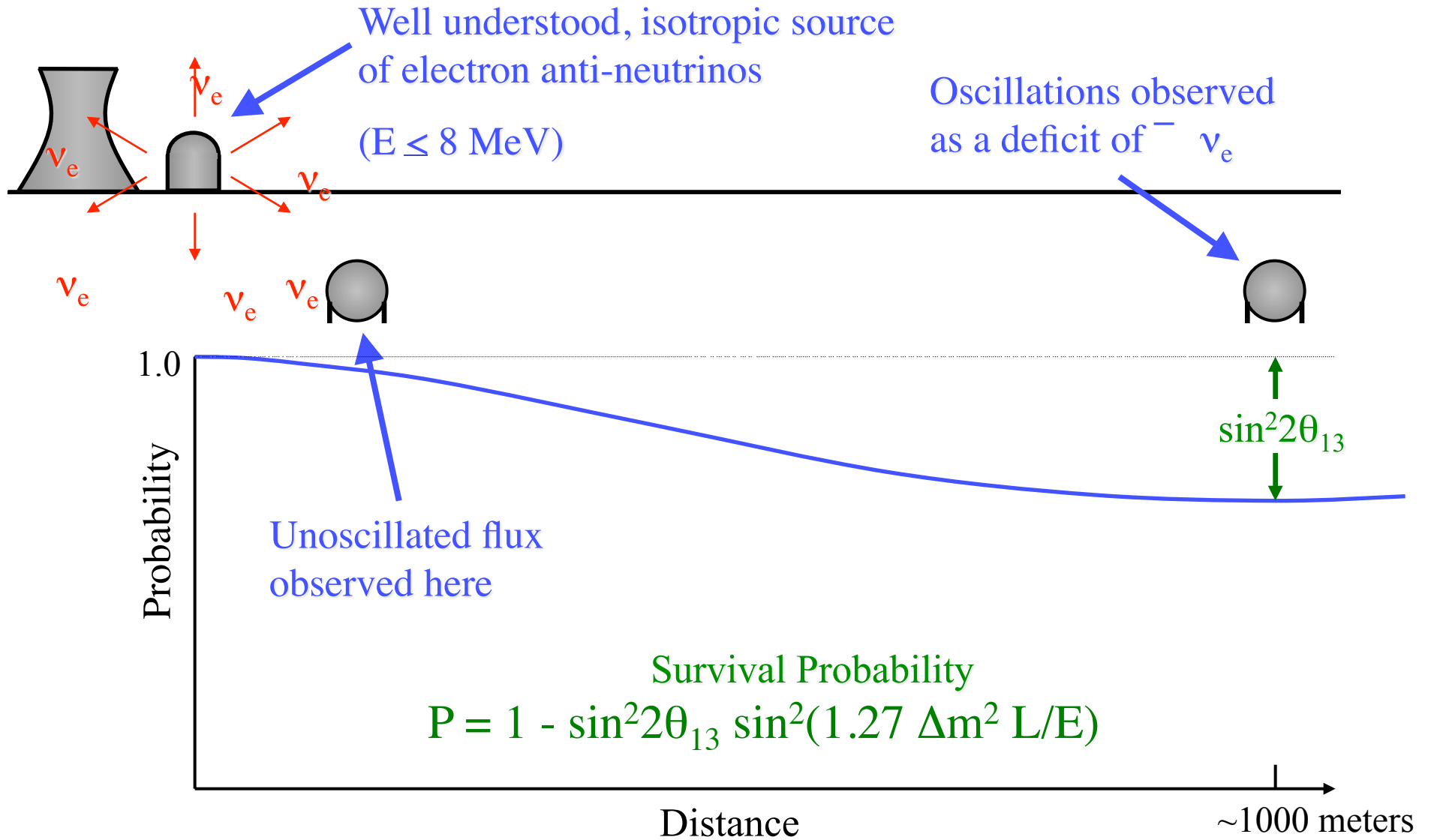


Moderately large entries
except for one,
which might be zero!

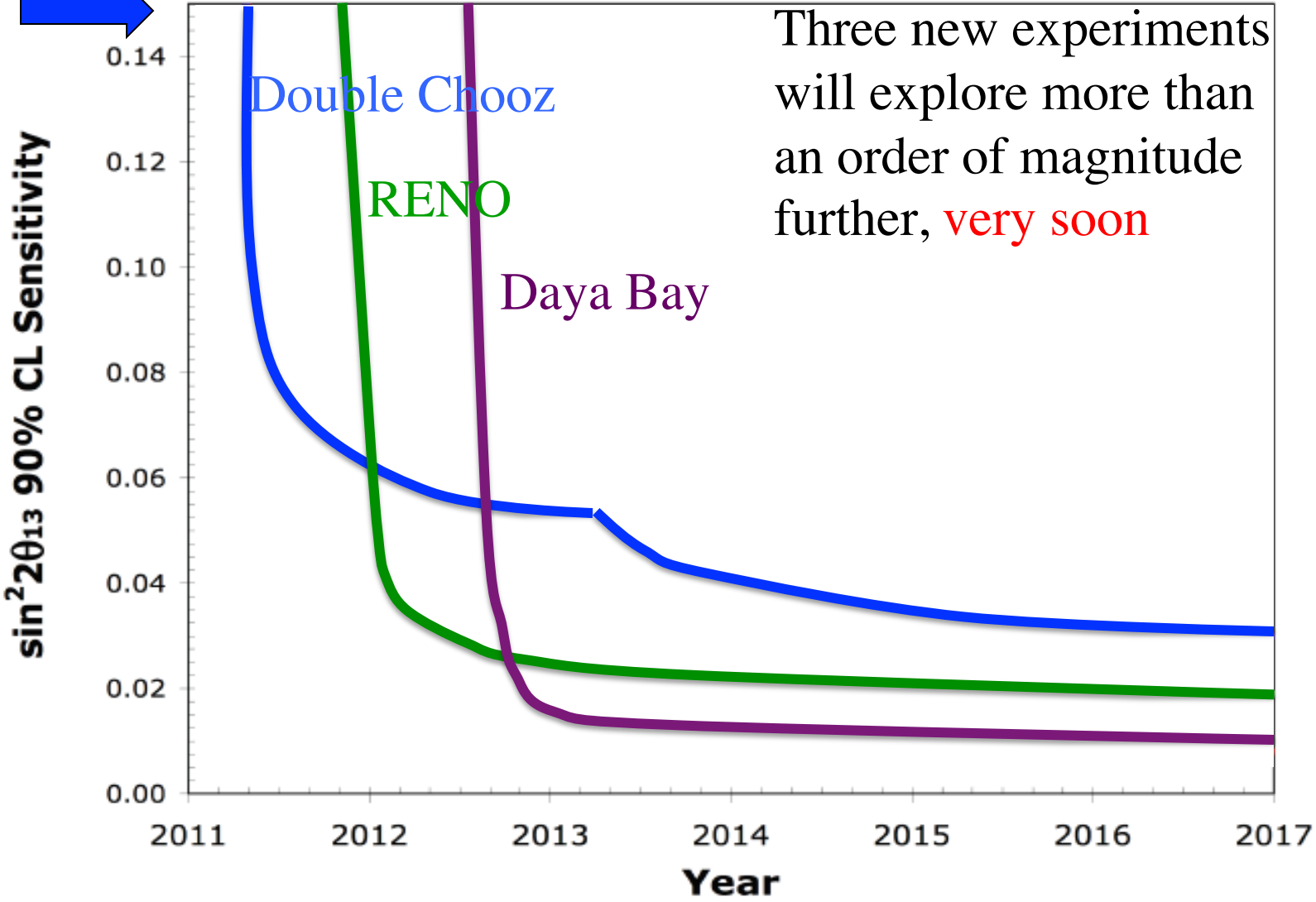
The Frenzy to Find θ_{13} is ON!!!!



How θ_{13} reactor experiments are designed:

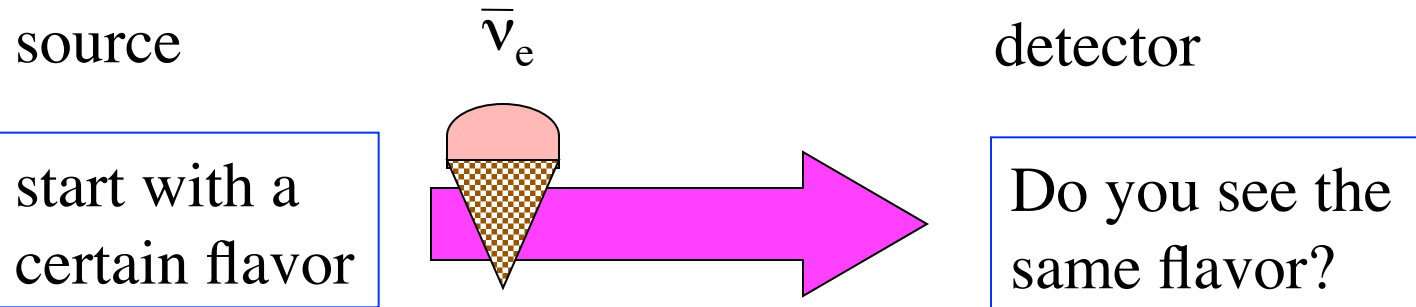


Today, the
Limit from
Reactors is here

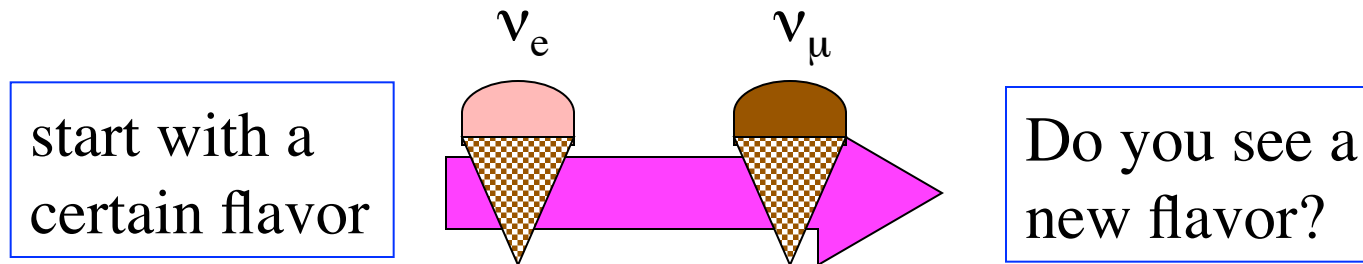


Three new experiments
will explore more than
an order of magnitude
further, **very soon**

Reactor Experiments are **disappearance** experiments



You can also search with **appearance** experiments



But appearance is more complicated because of
mysteries 2 and 3!

Mystery 2: Is there CP Violation in the Mixing Matrix Too?

$$c_{ij} = \cos\theta_{ij}$$

$$s_{ij} = \sin\theta_{ij}$$

The CP Violation Parameter

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

$$= \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

From Atmospheric
and Long Baseline
Disappearance
Measurements

From Reactor
Disappearance
Measurements

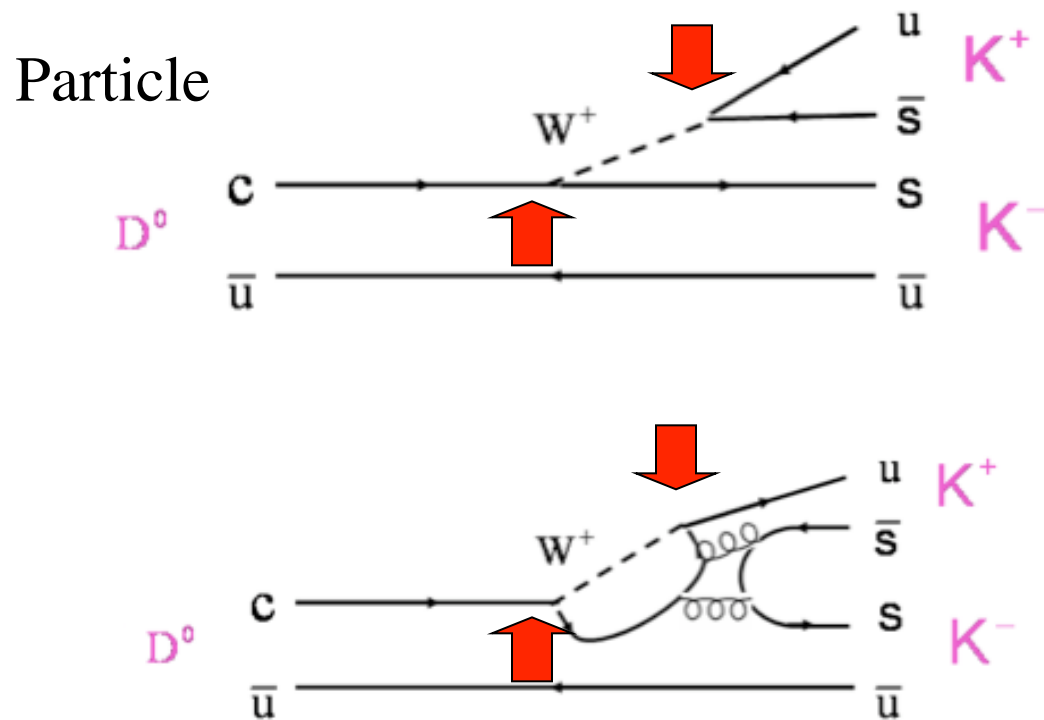
From $\nu_\mu \rightarrow \nu_e$
Appearance
Measurements

From Solar Neutrino
Measurements

If so, is it like the quark sector? or not? (Why???)

In the quark sector, CP violation can show up as a difference in rates of decay for particles vs. antiparticles.

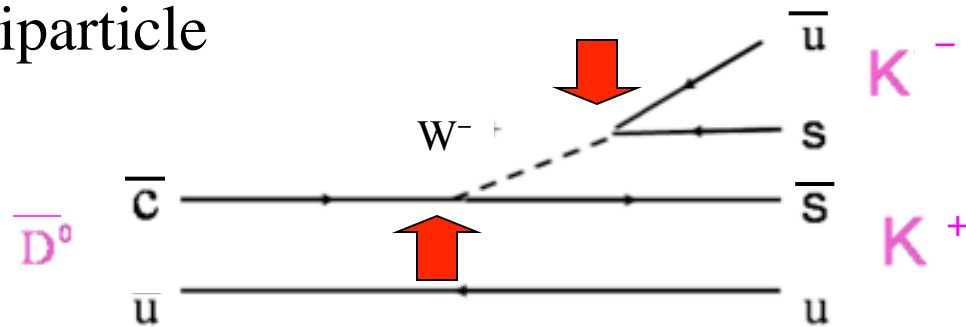
The effect shows up
when you have 2 paths to the same outcome...



You will get an
interference term
in the decay probability...

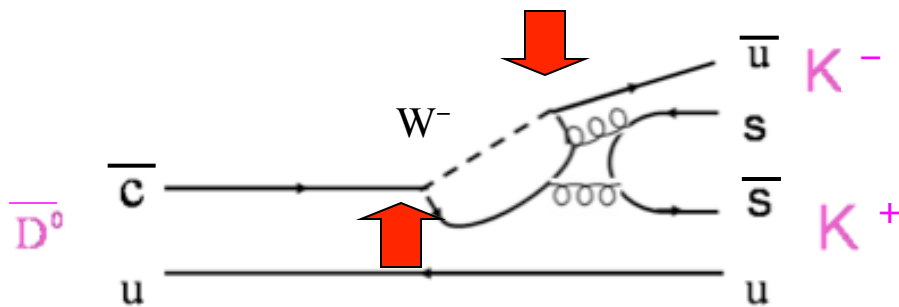
Now consider the $\overline{D^0}$

antiparticle



There are still 2 paths to the outcome.

Compared to the D^0
the interference
term changes sign!



e.g. D^0 and $\overline{D^0}$ decays can have different decay rates
if δ is nonzero!

Nonzero δ has been seen in the quark mixing matrix,
but it is a relatively small effect

But what about the lepton sector???

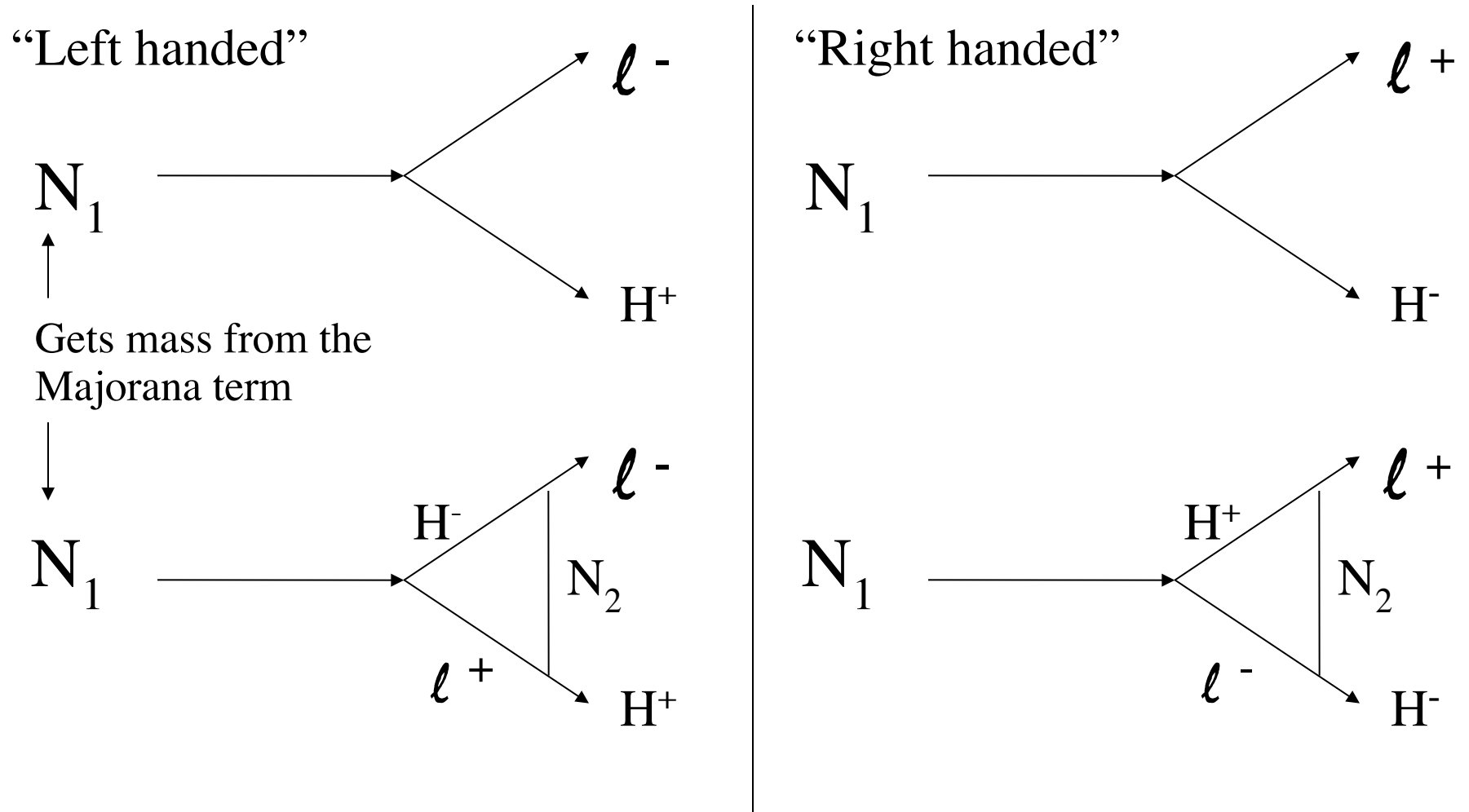
In a model where...

1. Neutrinos are Majorana particles
2. With GUT scale partners
3. And there is CP violation...

Then...

*CP violation in the neutrino sector may explain
the matter-antimatter asymmetry in the universe!*

Before the electroweak phase transition...



The interference terms will have opposite sign!

*Are neutrinos the
reason we exist?*



It's a big question and
it turns out to be
very hard to answer!

A first step would be observation
of CP violation in the light neutrinos

How would CP violation manifest itself?

In oscillation of muon-flavor to electron-flavor
at the atmospheric Δm^2

... it's all about the ν_e events again!

Recall the

$$2\nu \text{ osc formula: } P_{Osc} = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 L / E \right)$$

For 3 ν oscillations, in a vacuum, with CP Violation...

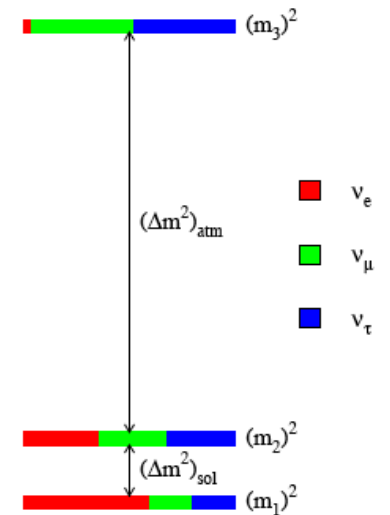
$$\begin{aligned}
 P = & \quad (\sin^2 \theta_{23} \sin^2 2\theta_{13}) (\sin^2 \Delta_{31}) \\
 & \mp \sin \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin^2 \Delta_{31} \sin \Delta_{21}) \\
 & + \cos \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin \Delta_{31} \cos \Delta_{31} \sin \Delta_{21}) \\
 & + (\cos^2 \theta_{23} \sin^2 2\theta_{12}) (\sin^2 \Delta_{21}).
 \end{aligned}$$

We want to see
if δ is nonzero

terms depending on
mixing angles

terms depending on
mass splittings

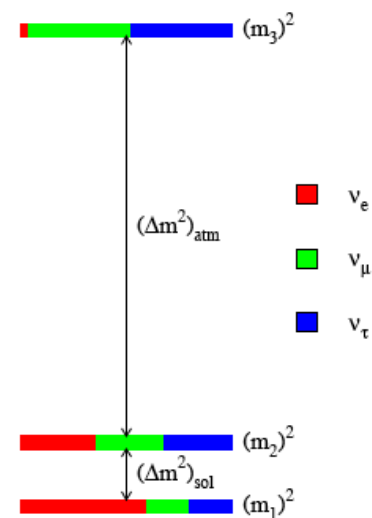
$$\Delta_{ij} = \Delta m_{ij}^2 L / 4E_\nu$$



Seeing CP violation is all about **interference**.

in a vacuum...

$$\begin{aligned}
 P = & \quad (\sin^2 \theta_{23} \sin^2 2\theta_{13}) (\sin^2 \Delta_{31}) \\
 & \mp \sin \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin^2 \Delta_{31} \sin \Delta_{21}) \\
 & + \cos \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin \Delta_{31} \cos \Delta_{31} \sin \Delta_{21}) \\
 & + (\cos^2 \theta_{23} \sin^2 2\theta_{12}) (\sin^2 \Delta_{21}).
 \end{aligned}$$



We want to see
if δ is nonzero

The δ -dependent terms
arise from **interference** between the
 Δm_{13}^2 and Δm_{12}^2 oscillations

Our equation flips sign between

$$\nu_{\mu} \rightarrow \nu_e \text{ \& \ } \bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$$

in a vacuum...

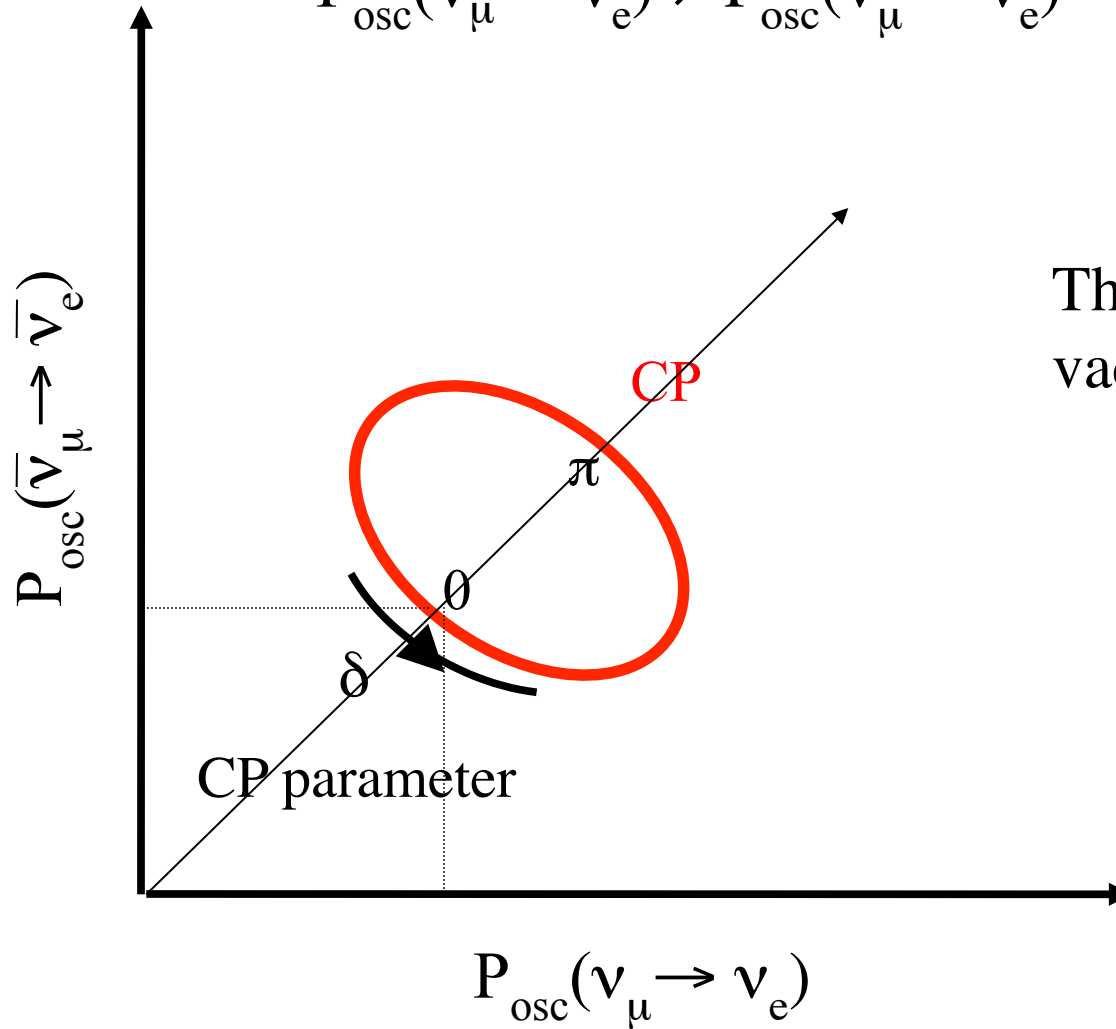
$$P = \begin{aligned} & (\sin^2 \theta_{23} \sin^2 2\theta_{13}) (\sin^2 \Delta_{31}) \\ & \mp \sin \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin^2 \Delta_{31} \sin \Delta_{21}) \\ & + \cos \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin \Delta_{31} \cos \Delta_{31} \sin \Delta_{21}) \\ & + (\cos^2 \theta_{23} \sin^2 2\theta_{12}) (\sin^2 \Delta_{21}). \end{aligned}$$

what we want
to measure

The matter and antimatter oscillation
probabilities will be different!

The classic idea for how to see CP violation:

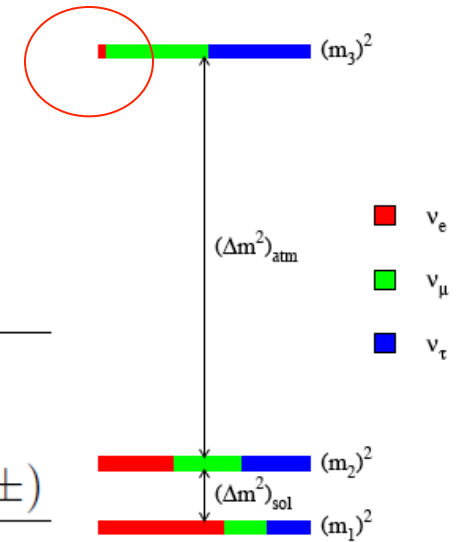
$$P_{\text{osc}}(\nu_{\mu} \rightarrow \nu_e) \neq P_{\text{osc}}(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)$$



This is in a vacuum (or air).

Most parameters are well known...

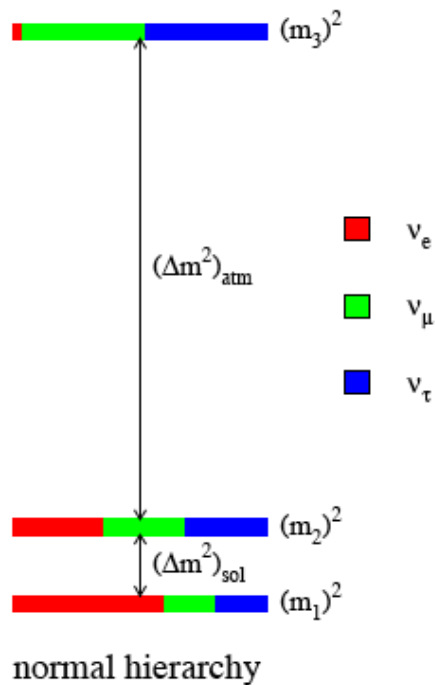
| Parameter | Present: | | Assumed Future: | |
|--|----------|----------------------|-----------------|----------------------|
| | Value | Uncert. (\pm) | Value | Uncert. (\pm) |
| $\Delta m_{21}^2 \times 10^{-5} \text{eV}^2$ | 7.65 | 0.23 | 7.65 | — |
| $\Delta m_{31}^2 \times 10^{-3} \text{eV}^2$ | 2.40 | 0.12 | 2.40 | 0.02 |
| $\sin^2(2\theta_{12})$ | 0.846 | 0.033 | 0.846 | — |
| $\sin^2(2\theta_{23})$ | 1.00 | 0.02 | 1.00 | 0.005 |
| $\sin^2(2\theta_{13})$ | 0.11 | 0.06 | 0.05 | 0.005 |



Except for that pesky θ_{13} !

We will end up having to quote our sensitivity
as allowed regions in both θ_{13} and δ

This design requires a long baseline!



The Probability for Oscillations...

$$P_{osc} = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L/E)$$



P is maximized when $\Delta m^2(L/E) \sim 1$

The atmospheric $\Delta m^2 \sim 0.003 \text{ eV}^2$

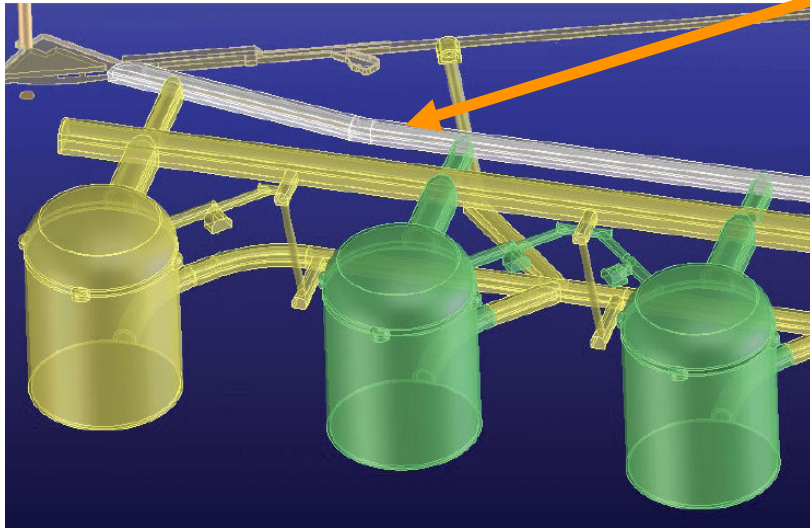
If $E \sim 3 \text{ GeV}$

Then $L = 1000 \text{ km} !!!$

E.g., LBNE -- starting in 2021

Beam from Fermilab

Shoots to detectors in South Dakota
1300 km

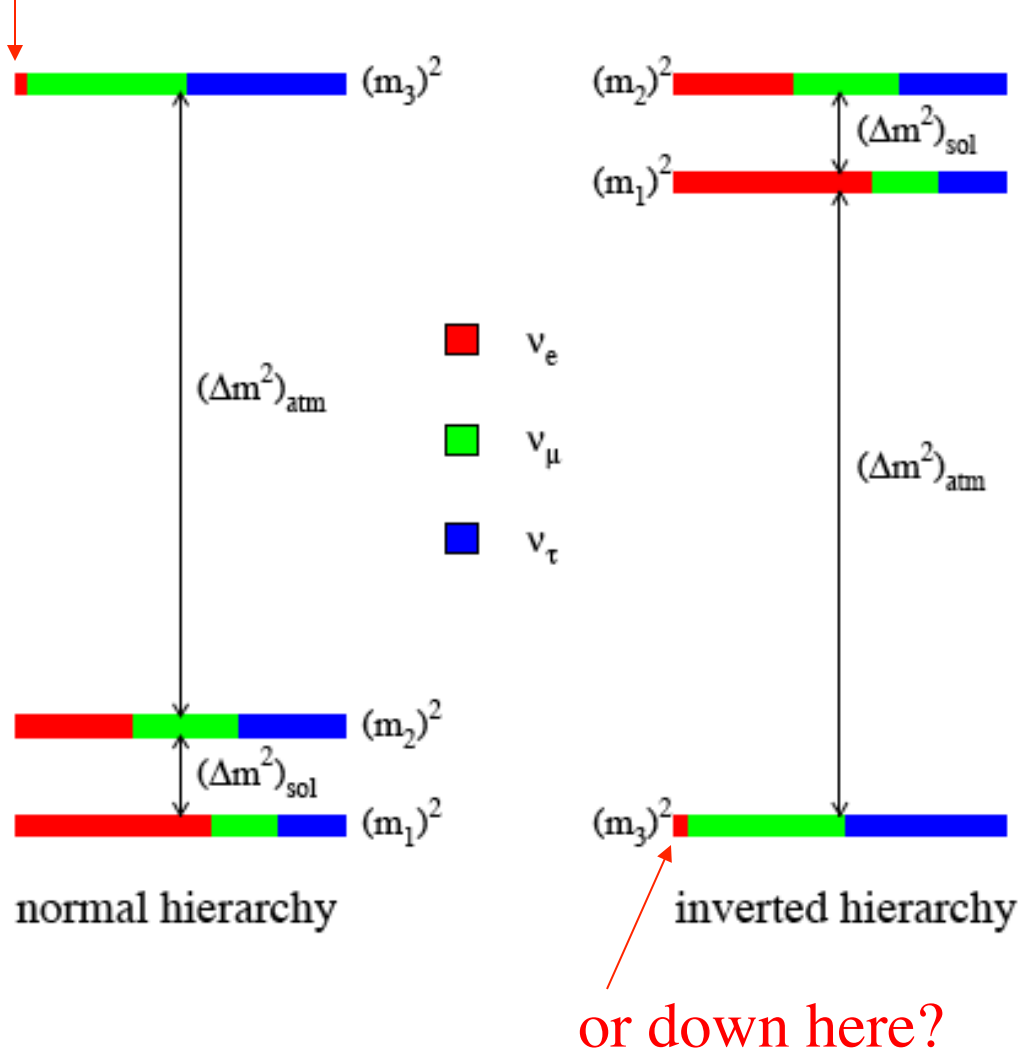


**And there is lots and lots
of matter along a 1300 km path!**

also true for LAGUNA and HyperK designs

Mystery 3:

Is the small ν_e content up here?



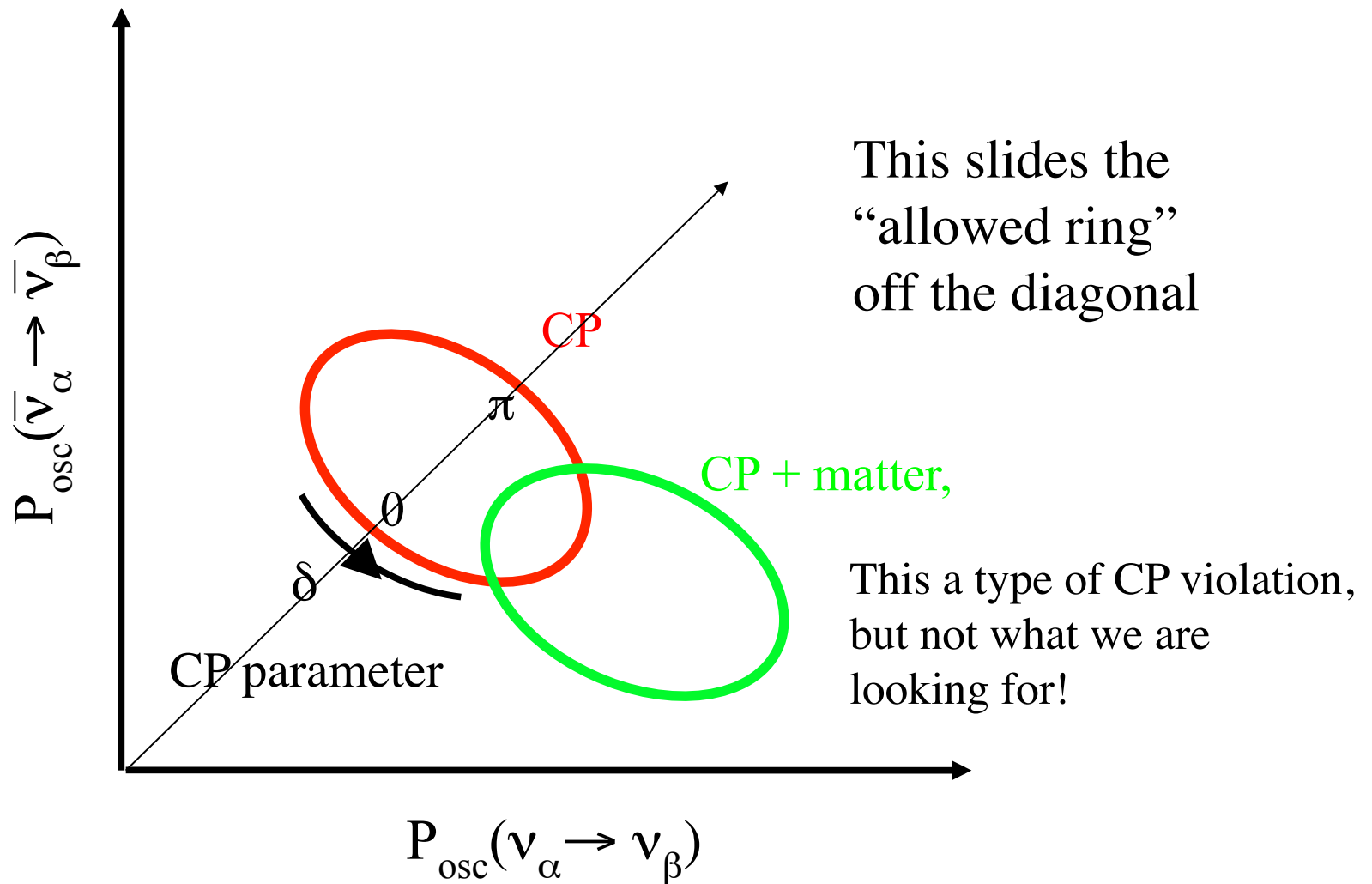
i.e. What is the “mass hierarchy”?

Is it “opposite” to the quark sector?
(WHY???)

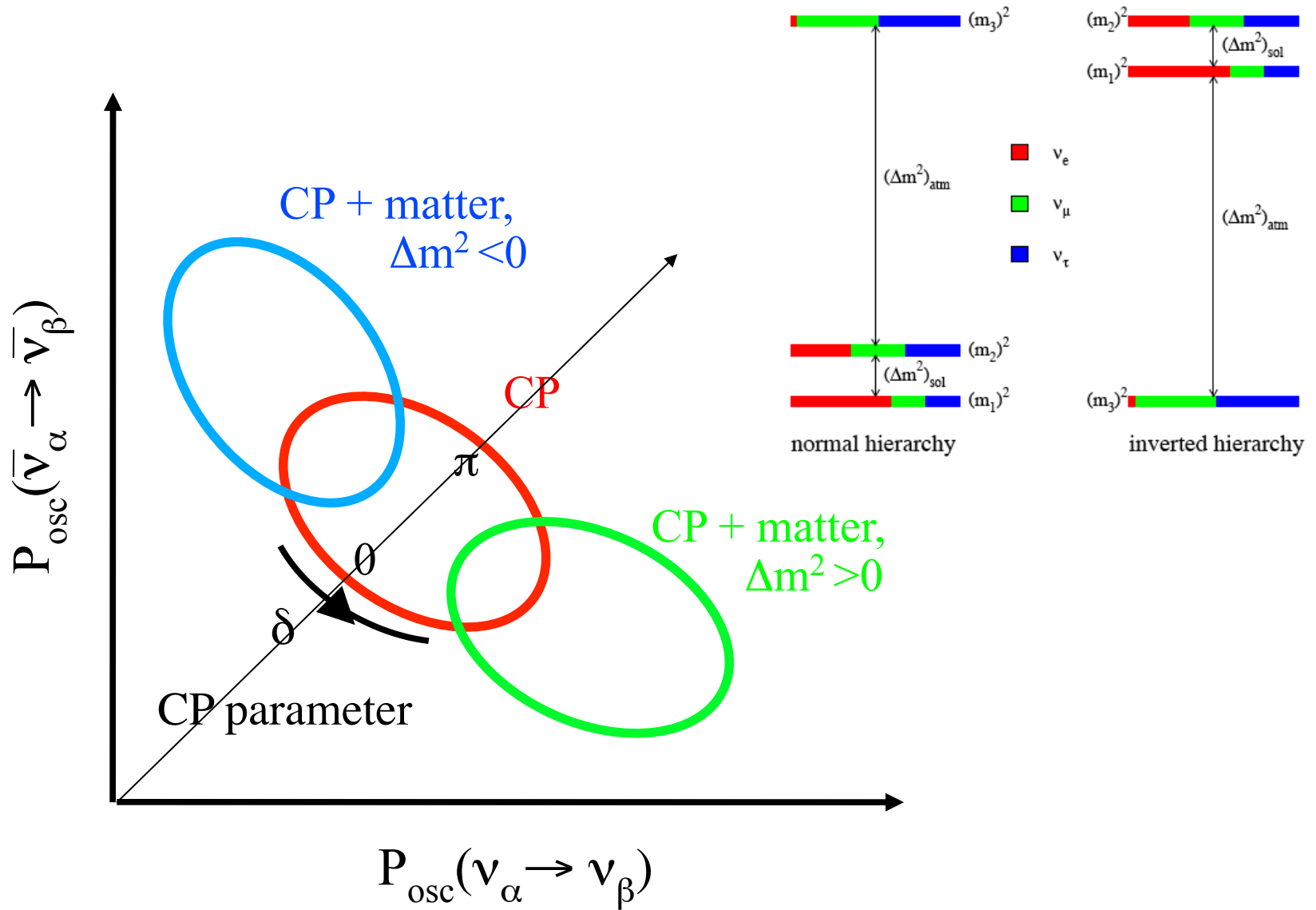
This affects the rates of
 $\nu_\mu \rightarrow \nu_e$
Versus
 $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$

The ground is made of matter (electrons)
not antimatter (positrons)

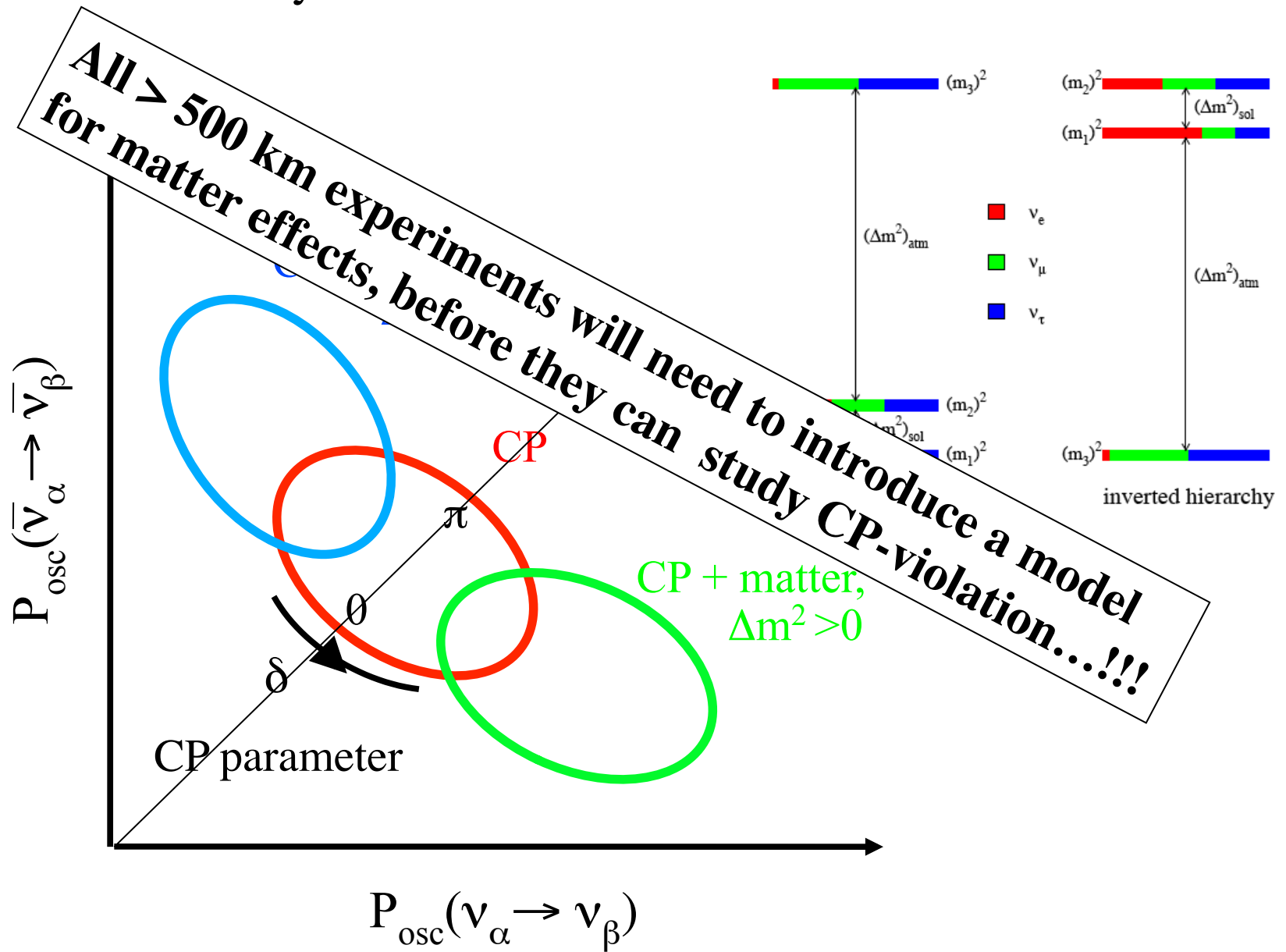
Forward scattering affects neutrinos differently than antineutrinos.

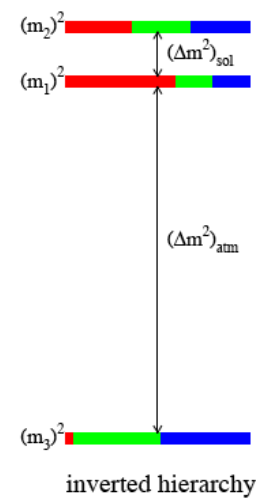
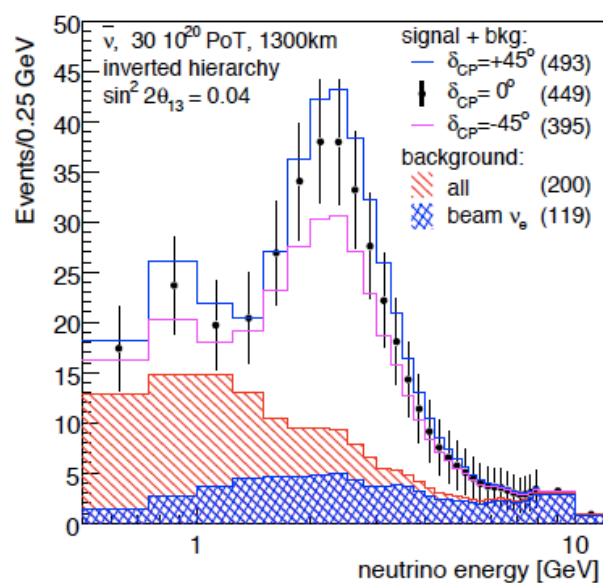
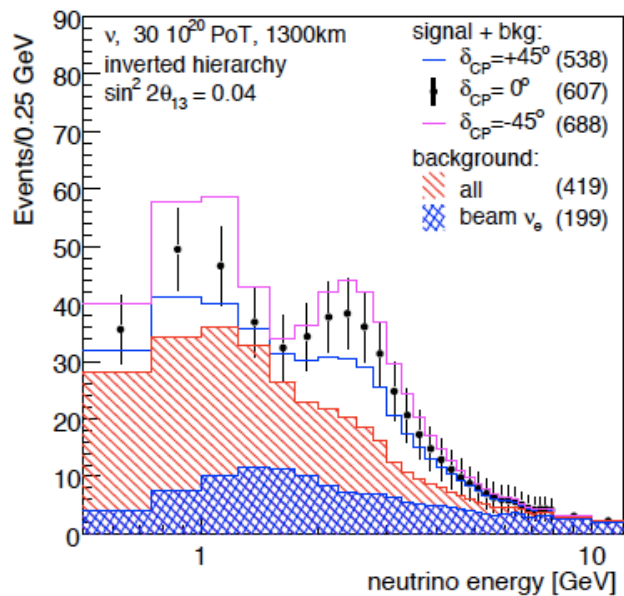
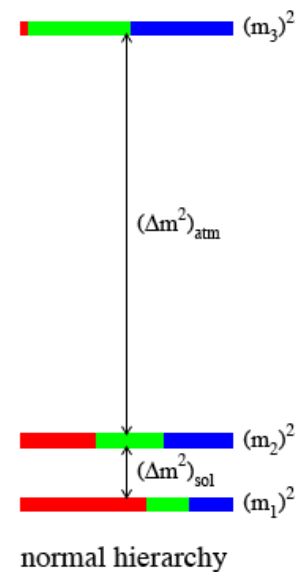
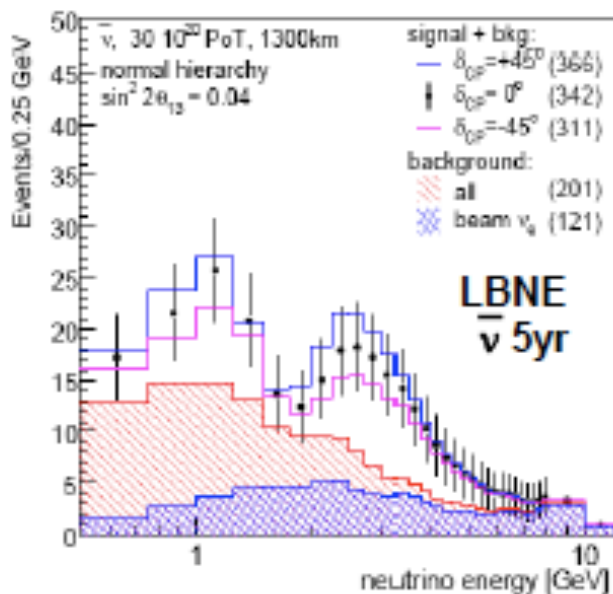
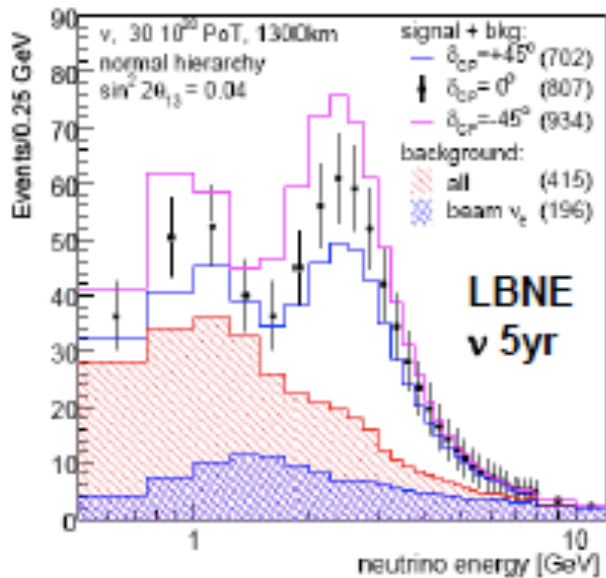


We actually don't know which direction...



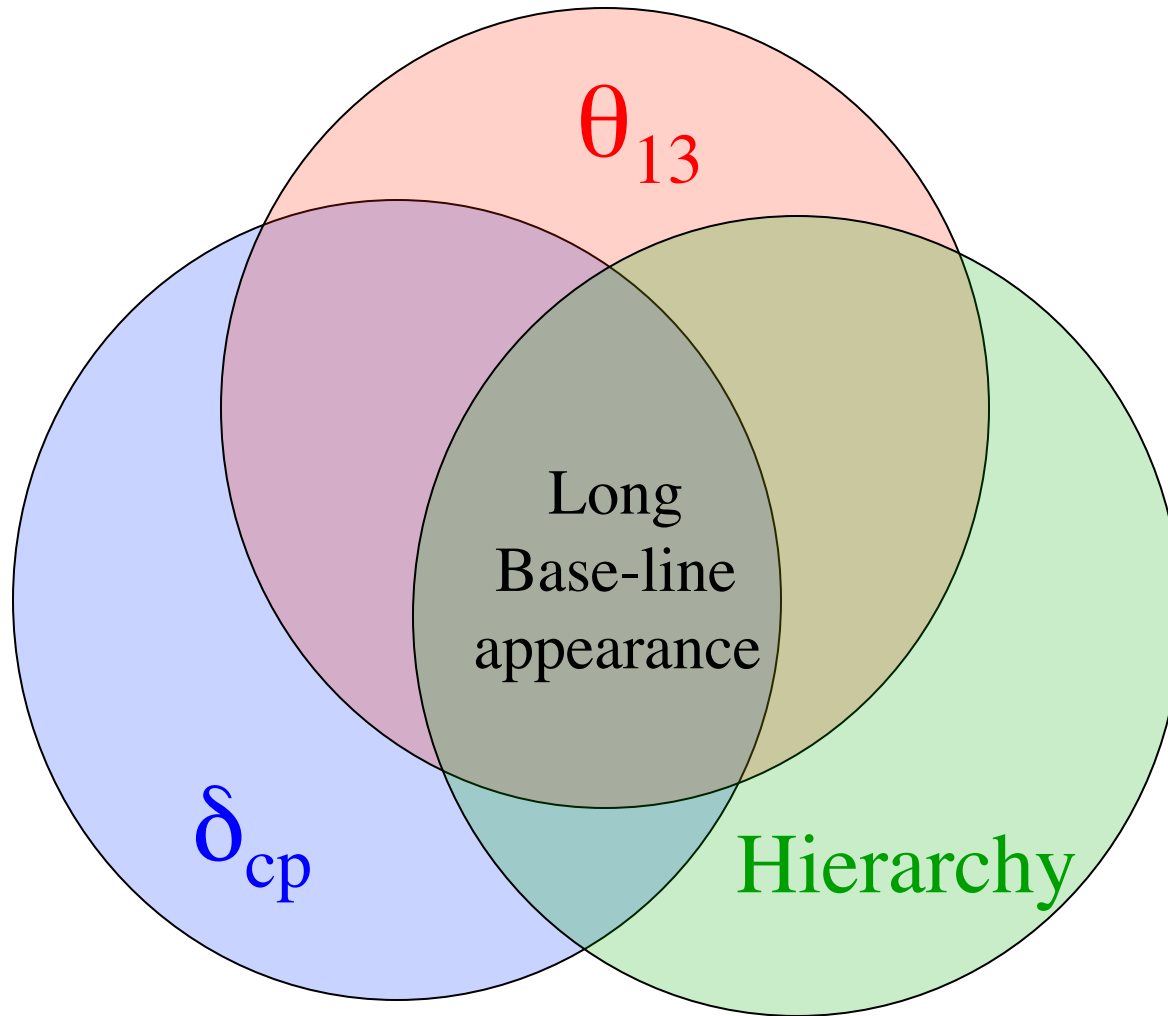
We actually don't know which direction...





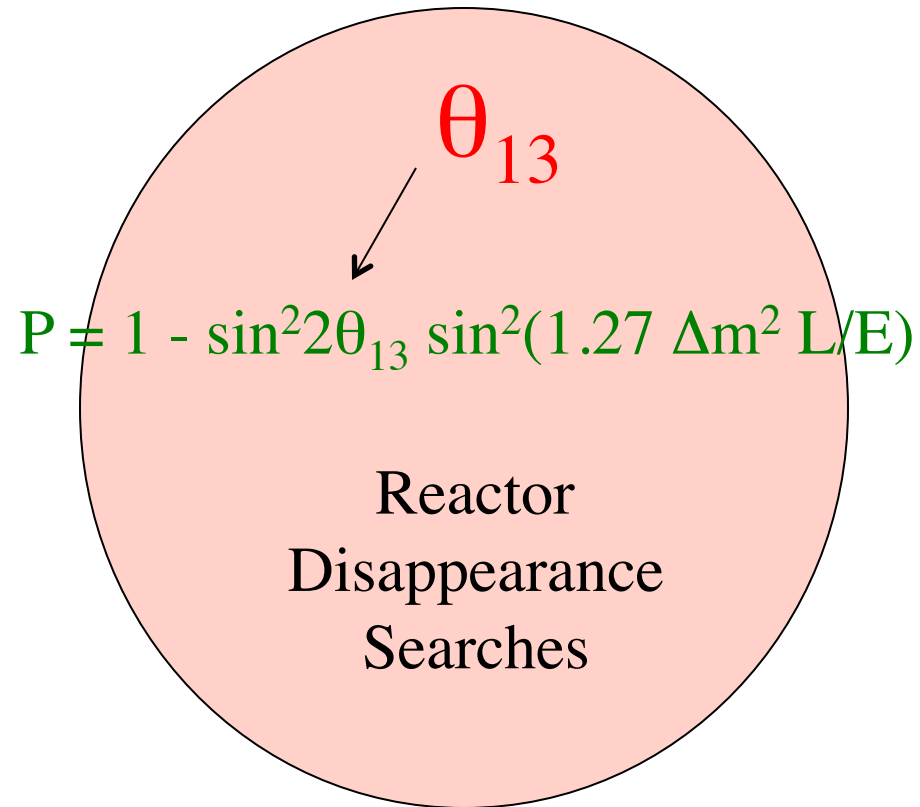
So now matters become very convoluted.

If I ask: How well can I unravel these mysteries?

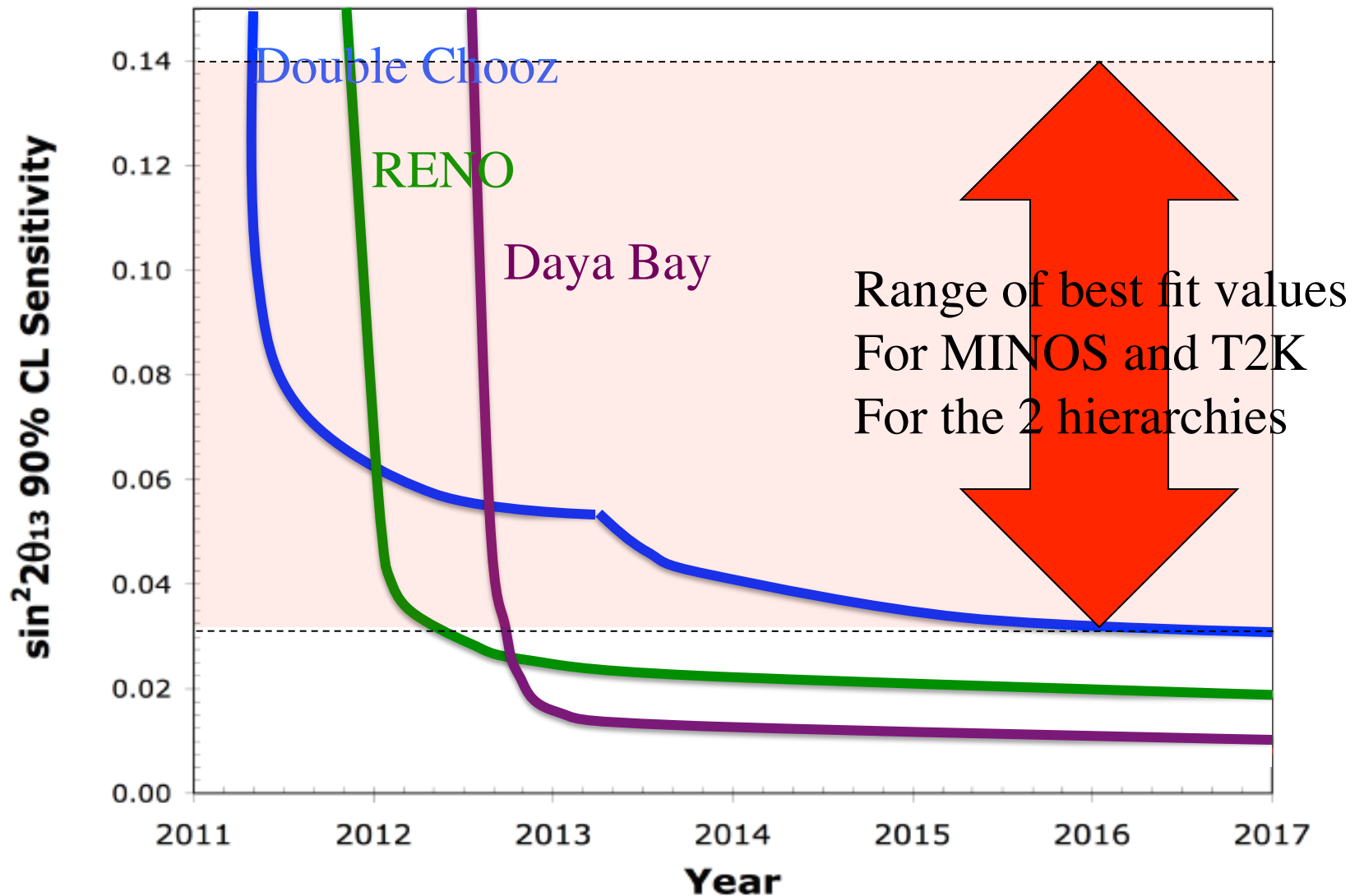


But don't forget the reactors!

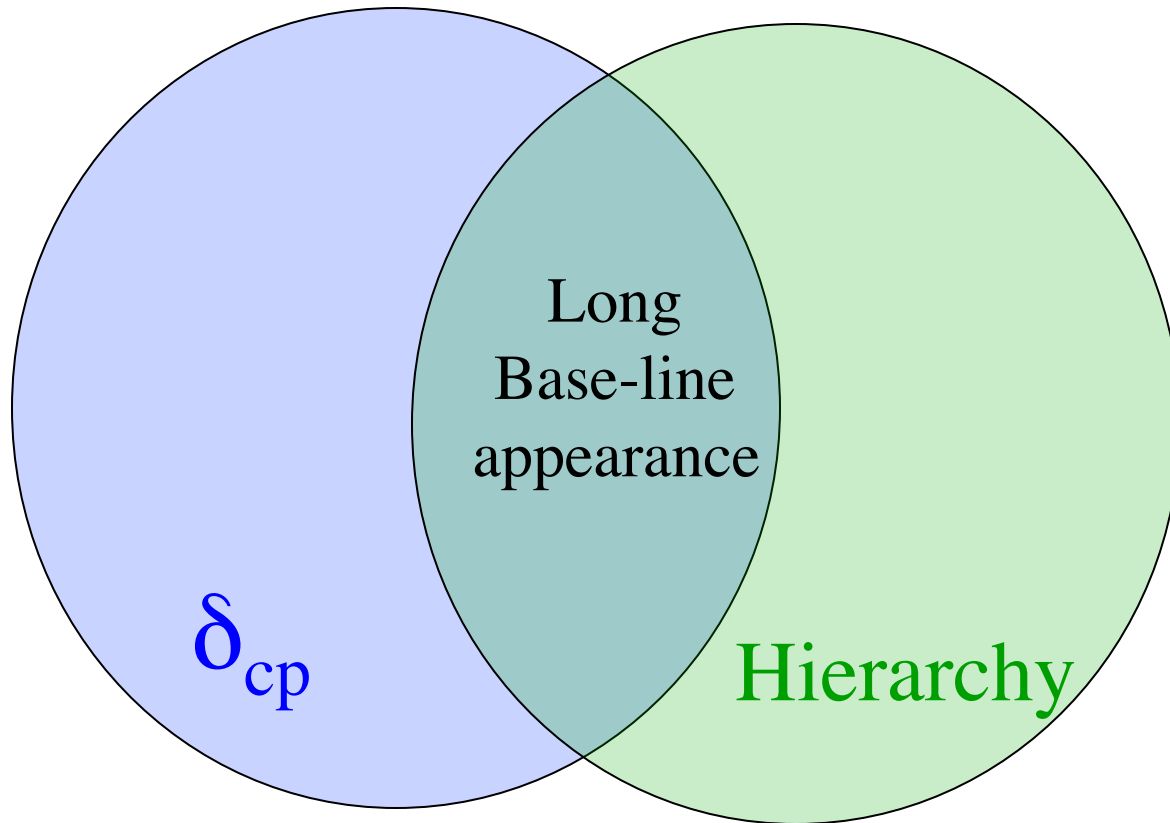
They can work on one and only one mystery



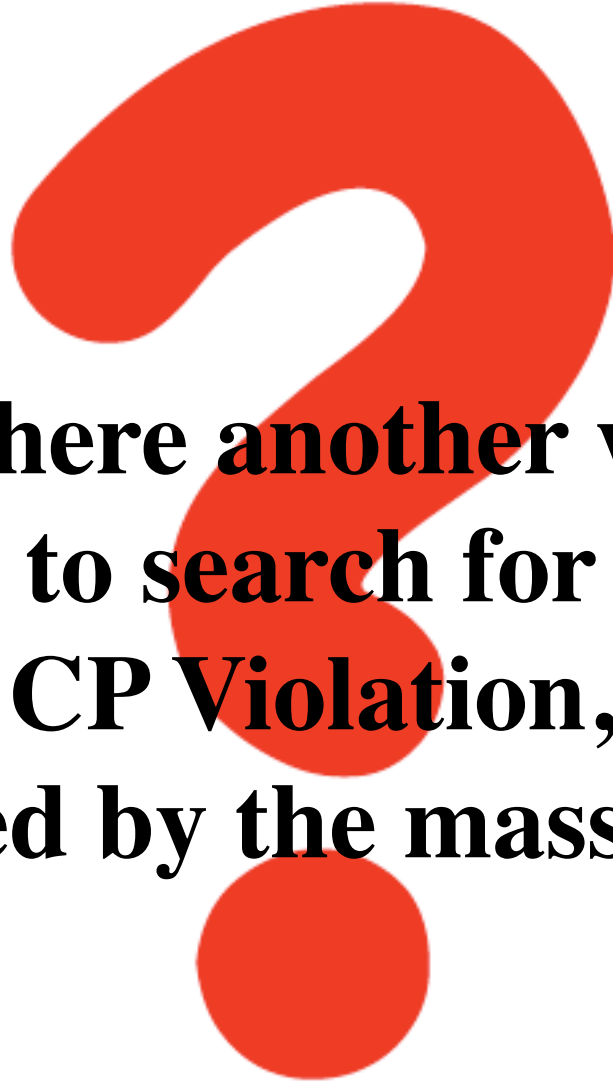
And it seems that the reactors are looking at just the right place!
New from appearance experiments this summer!



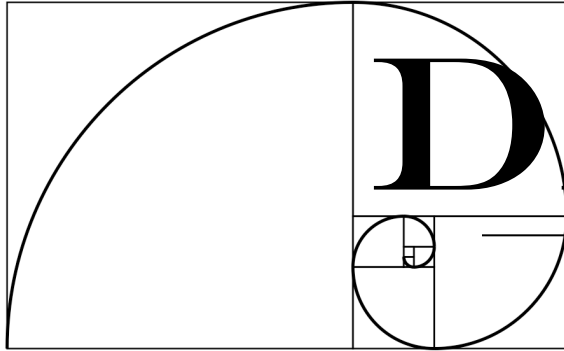
This still leaves us with:



And this is still messy



**Is there another way
to search for
CP Violation,
Not confused by the mass hierarchy?**



DAE δ ALUS

**Decay
At rest
Experiment
for δ_{cp} studies
At the
Laboratory for
Underground
Science**

- New ν source for an ultralarge detector
 - Enhanced neutrino oscillation program
 - New experiments possible
- Complementary to the long baseline proposals
 - Comparable measurements for osc parameters
 - Much improved measurements by combining DAE δ ALUS and long-baseline!

The plan: Use $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
 and exploit the L/E dependence in absolute rates

in a vacuum...

$$\begin{aligned}
 P = & (\sin^2 \theta_{23} \sin^2 2\theta_{13}) (\sin^2 \Delta_{31}) \\
 & \mp \sin \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin^2 \Delta_{31} \sin \Delta_{21}) \\
 & + \cos \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin \Delta_{31} \cos \Delta_{31} \sin \Delta_{21}) \\
 & + (\cos^2 \theta_{23} \sin^2 2\theta_{12}) (\sin^2 \Delta_{21}).
 \end{aligned}$$

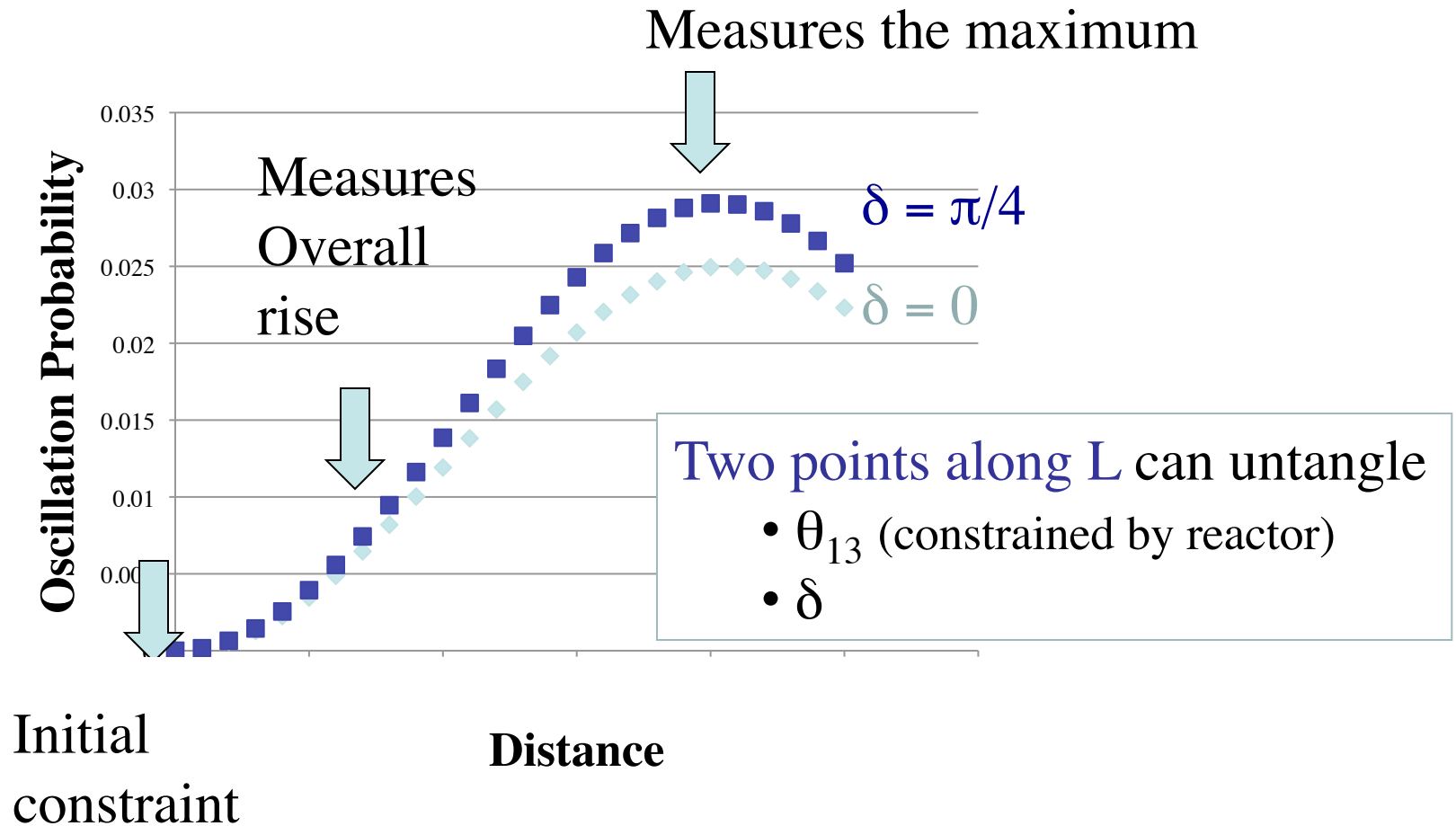
We want to see
 if δ is nonzero

terms depending on
 mixing angles

terms depending on
 mass splittings

$$\Delta_{ij} = \Delta m_{ij}^2 L / 4E_\nu$$

The terms that depend on δ change the oscillation wave L dependence



At low energy and short distances,
 Modifications to this vacuum equation
 By the mass hierarchy are very small

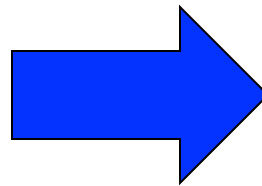
$$\begin{aligned}
 P = & \quad (\sin^2 \theta_{23} \sin^2 2\theta_{13}) (\sin^2 \Delta_{31}) \\
 & \mp \sin \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin^2 \Delta_{31} \sin \Delta_{21}) \\
 & + \cos \delta (\sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12}) (\sin \Delta_{31} \cos \Delta_{31} \sin \Delta_{21}) \\
 & \quad + (\cos^2 \theta_{23} \sin^2 2\theta_{12}) (\sin^2 \Delta_{21}).
 \end{aligned}$$

We want to see
 if δ is nonzero

terms depending on
 mixing angles

terms depending on
 mass splittings

I can change L and E,
 But maintain L/E



$$\Delta_{ij} = \Delta m_{ij}^2 L / 4E_\nu$$

And access the same oscillations as long baseline!

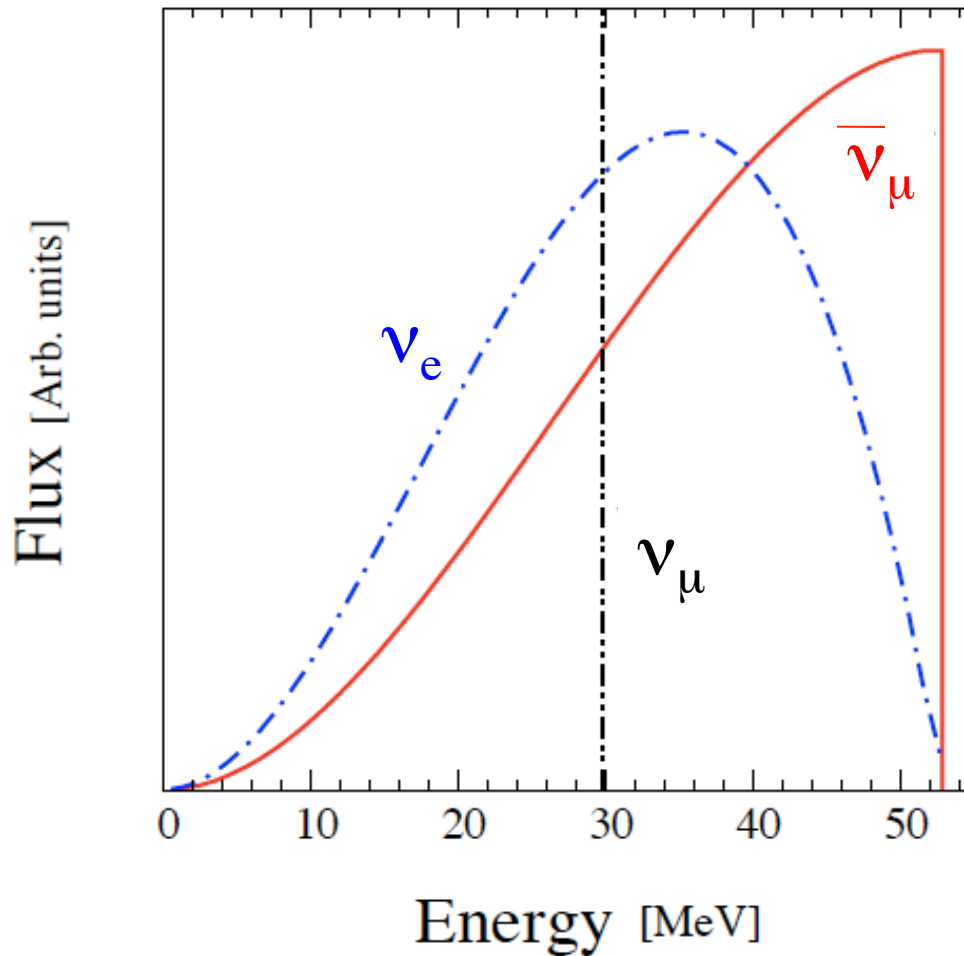
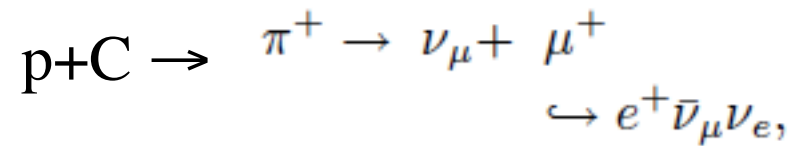
So we need:

A multiple baseline
(at least 3 points),
Short baseline,
Low energy,
Experiment
To isolate

$$\delta_{cp}$$

A really nice
low-energy beam

A π^+ decay at rest beam:



Shape driven by nature!

Only the normalization
varies from beam to beam

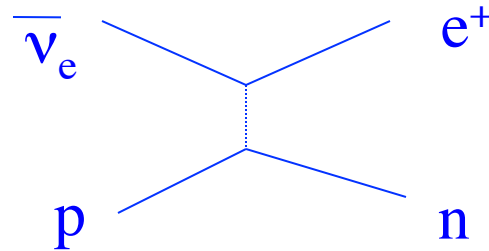
No intrinsic $\bar{\nu}_e$

Perfect for a

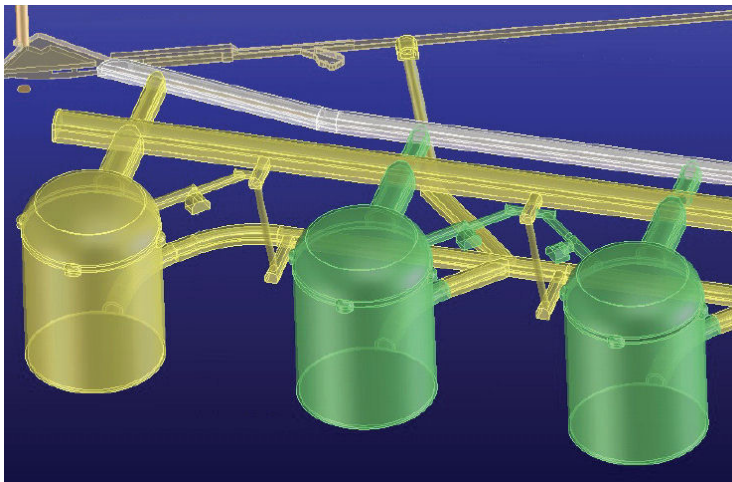
$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
search

How do you observe $\sim 50 \text{ MeV } \bar{\nu}_e$ events?

The signal:
inverse beta decay, IBD

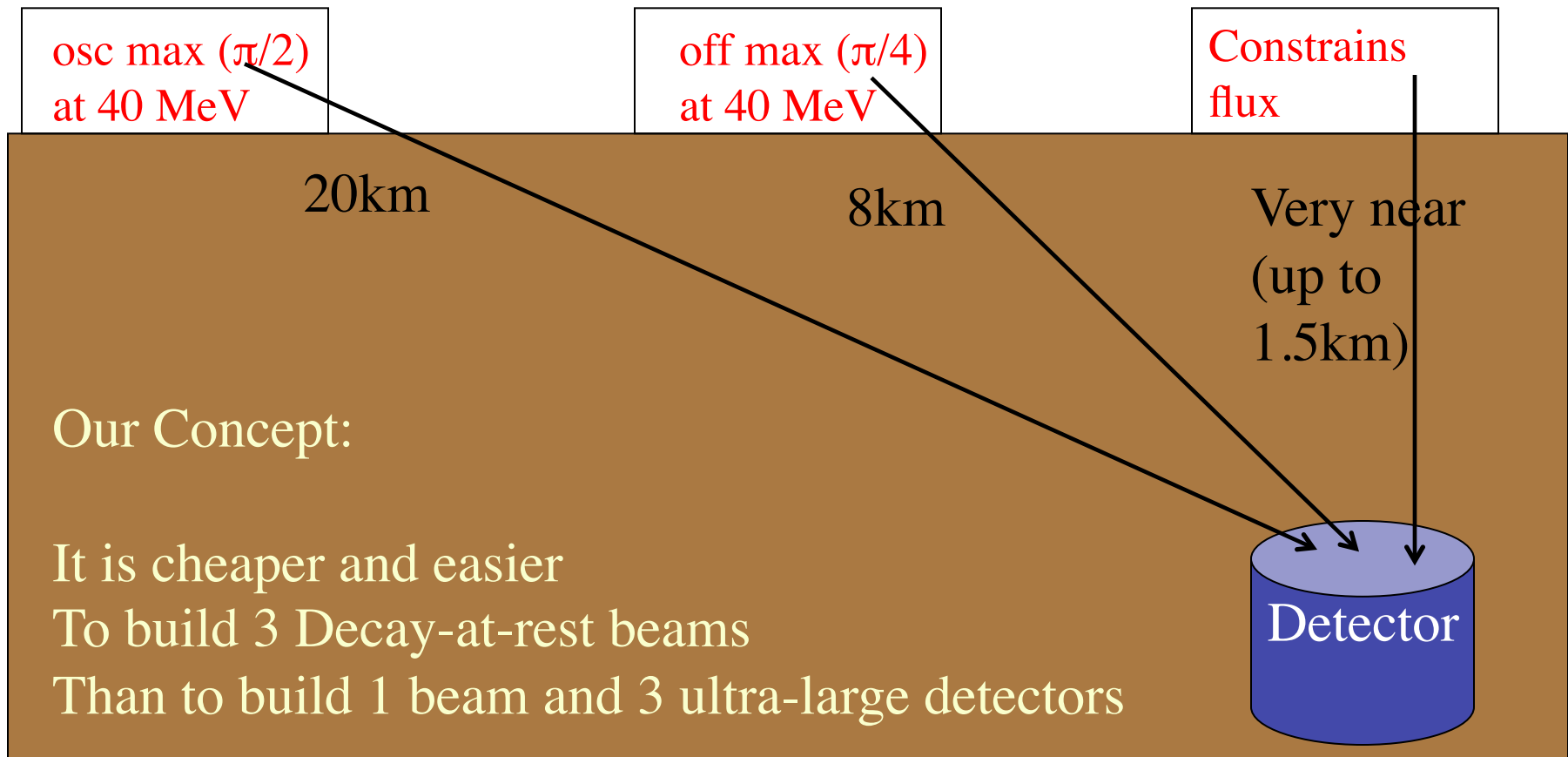


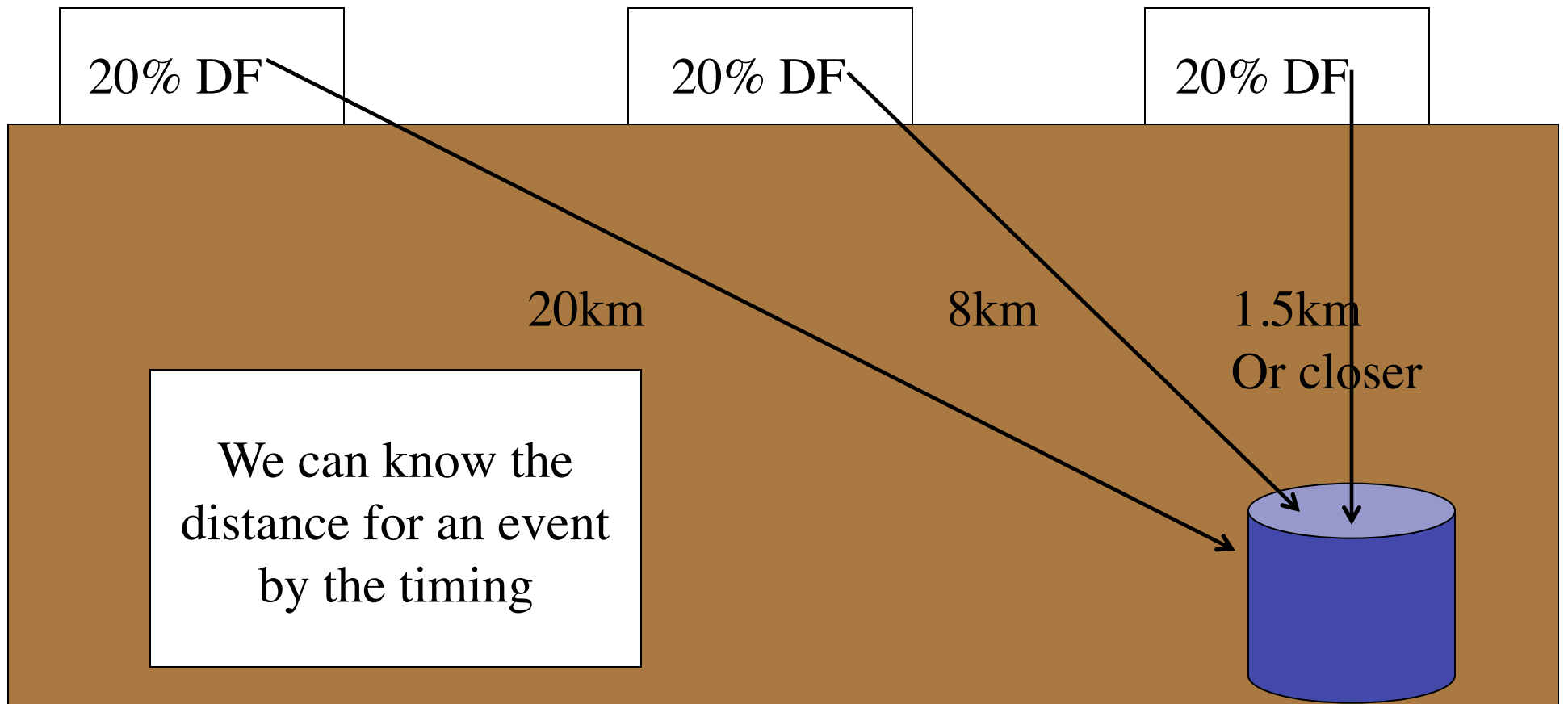
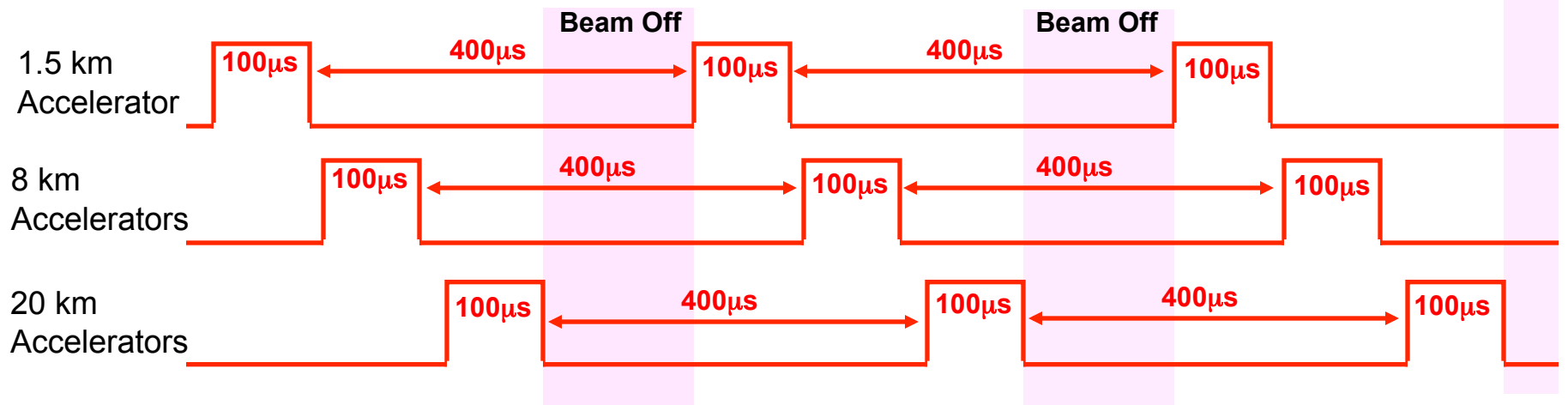
You need a lot of **free protons!**



Use the same ultra-large
detector system as
the long baseline

Water – MEMPHYS, LBNE
Oil -- LENA





Luckily, there are other people who want low-energy,
High-power proton sources!

Accelerator driven systems for thorium reactors



Columbia University,
10-12 October 2011

Like the accelerators that we need
(and unlike Project X, SNS, etc)

- ~800 MeV protons only
 Single energy, no upgrade path
- No fancy beam structure
- No stringent emittance requirements

Among all of the types of accelerators out there...

Cyclotrons
Synchrotrons
Linacs
FFAGs
etc.

Why cyclotrons?

Inexpensive,
Only practical below ~ 1 GeV
(ok for us!)
Only good if you don't need
timing structure (ok!)
Typically single-energy (ok!)
Taps into existing industry

Very interesting
R&D ongoing,
but these
machines
are not yet
proven

Can do what
we need
right now,
but are expensive.

Use linacs if
you want a nice
beam for transfer
to another line
and flexibility
on energy (We don't)

*We do not rule out other
options, but cyclotrons
seem like a good fit.*

Approaches using cyclotrons:

The compact cyclotron with self-extraction

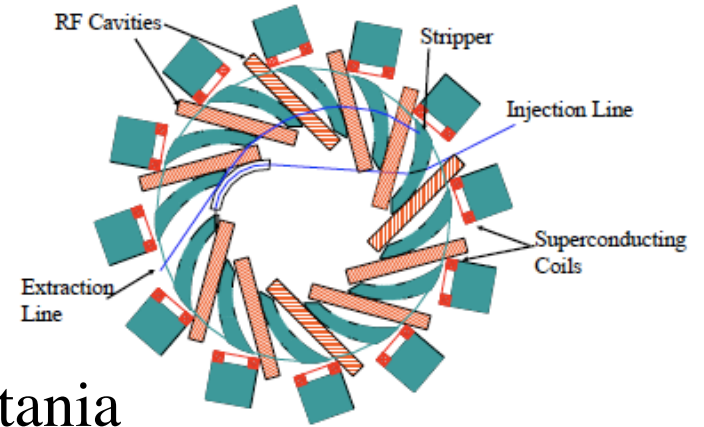


under development for DTRA at MIT

An H₂⁺ accelerator

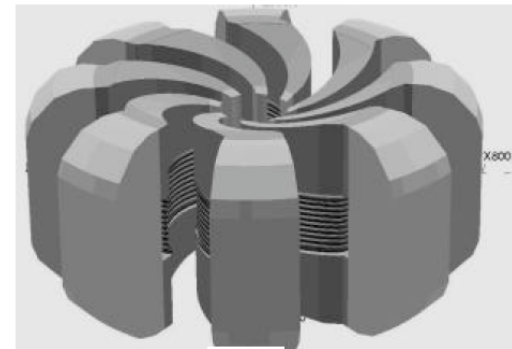
for ADS applications

Under dev. by INFN, Catania



The stacked cyclotron:

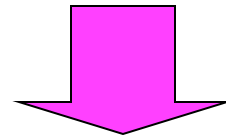
7 cyclotrons in one flux return



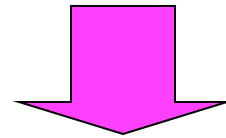
Under dev. for ADS at TAMU

Measurement strategy:

Using **near accelerator**
measure **absolute flux normalization** with ν -e events to $\sim 1\%$,
Also, measure the $\nu_e O$ event rate.



At far and mid accelerator,
Compare predicted to measured $\nu_e O$ event rates
to get the **relative flux normalizations between 3 accelerators**



In all three accelerators,
given the known flux, **fit for the $\overline{\nu}_\mu \rightarrow \overline{\nu}_e$ signal**
with free parameters: θ_{13} and δ

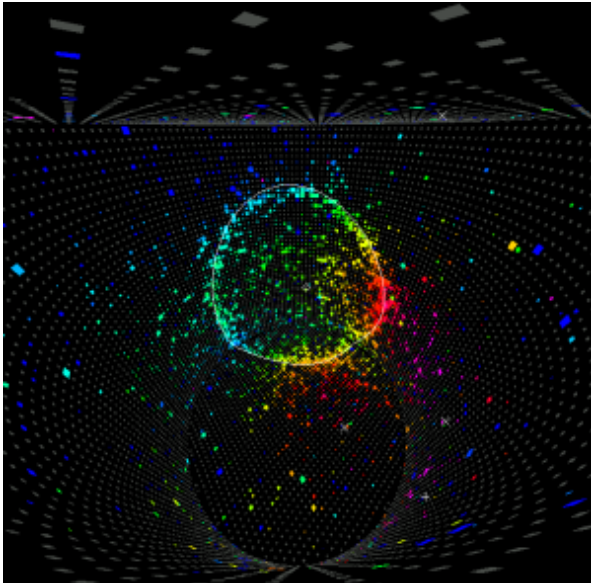
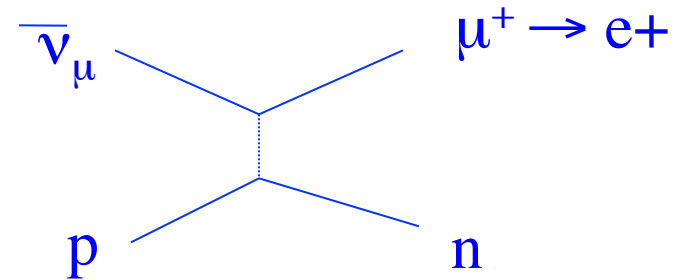
Non-beam backgrounds

Depends on the detector

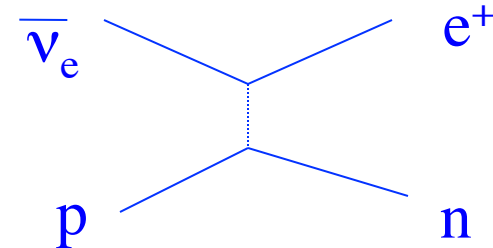
In water Cerenkov Detectors:

Atmospheric $\bar{\nu}_\mu$ “Invisible muons”:

$\bar{\nu}_\mu + p \rightarrow \mu^+ + n$ where
 μ^+ is below Cherenkov threshold,
stops and decays.



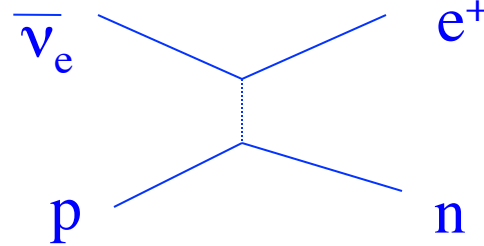
Looks just like



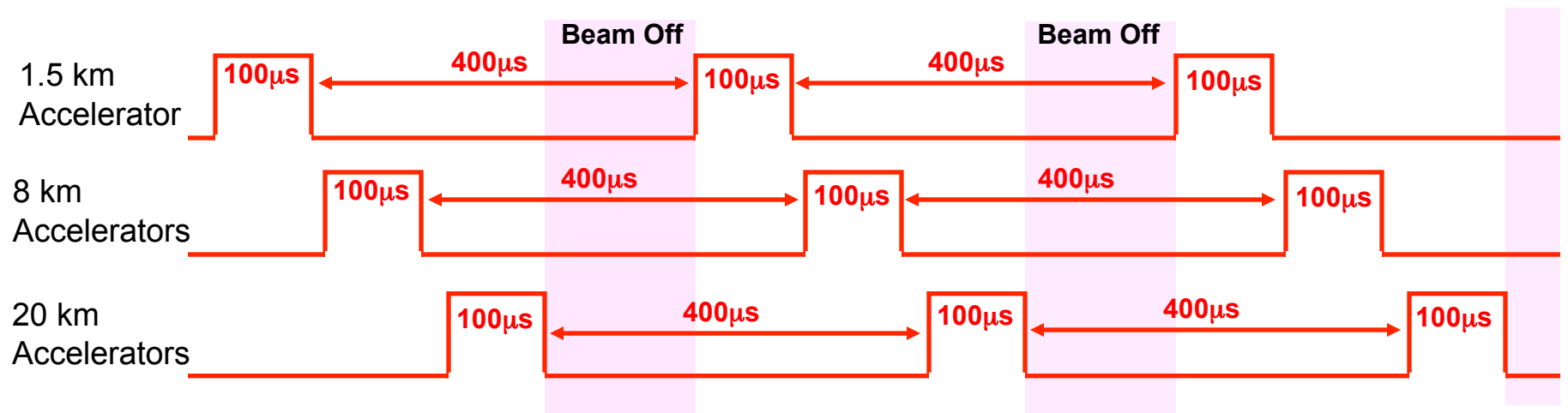
Not a background in scintillator!

Non-beam backgrounds, cont'd

- Atmospheric $\bar{\nu}_e$ IBD events:
 $\bar{\nu}_e + p \rightarrow e^+ + n$
- Diffuse supernova neutrinos



Measure all of these during the beam off periods:



Beam-related Background

- Intrinsic $\bar{\nu}_e$ in beam

From $\pi^- \rightarrow \mu^-$ events which failed to capture in the beam stop
 $\sim 4 \times 10^{-4}$ of ν_e rate (low)

- Beam ν_e in coincidence with random neutron capture signal
Estimated to be very small from Super-K rates

- ν_e -Oxygen CC scatters producing an electron+ n signal
Subsequent n from nuclear de-excitation should be very small.

All fall as $1/r^2$ from the 3 accelerators,
near accelerator provides a measurement

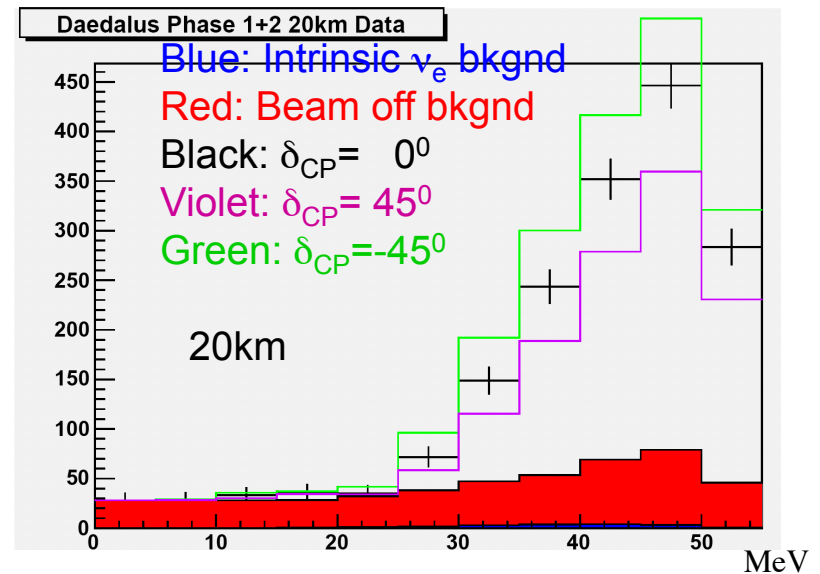
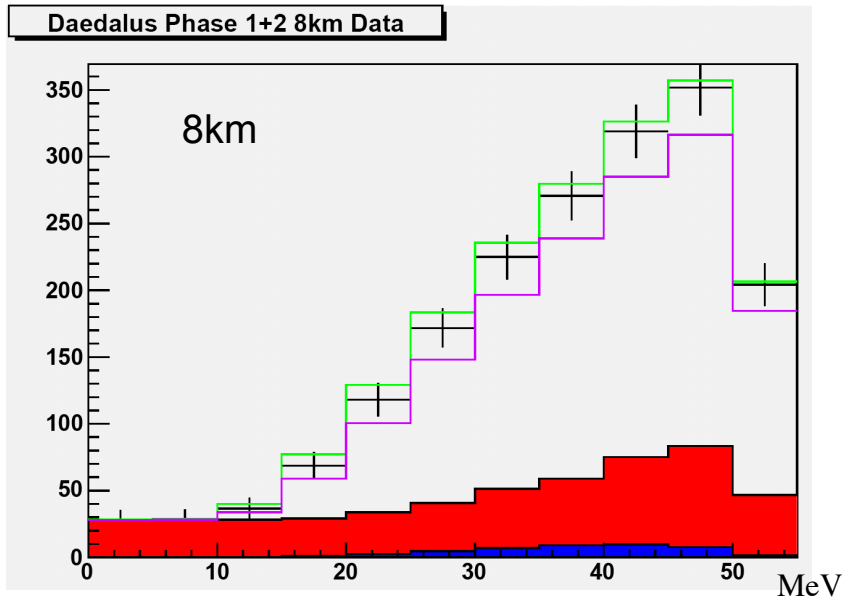
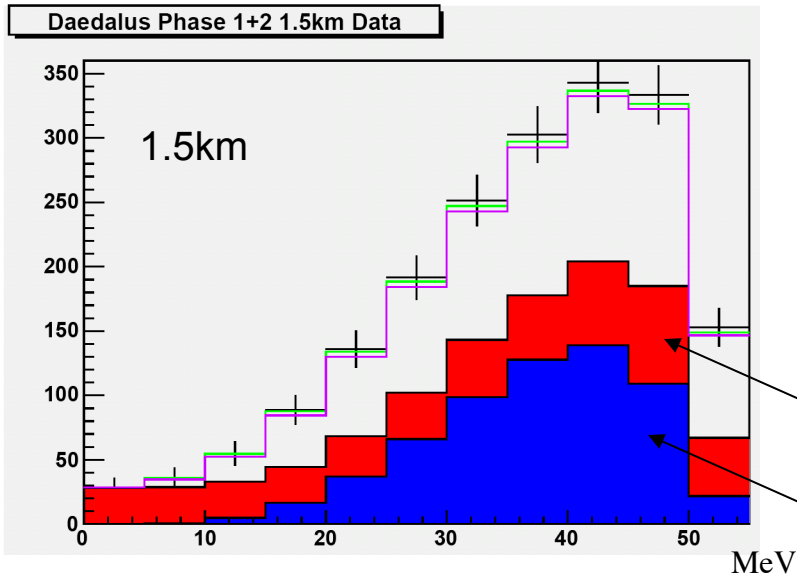
To discuss how well we can do, I need to pick a model...

- Water Cerenkov
- “Homestake Accelerator Arrangement” -- 1.5, 8 and 20 km
- Gd doping of water so the neutron can be observed
- 300 kt (we are stats limited, so you can scale)

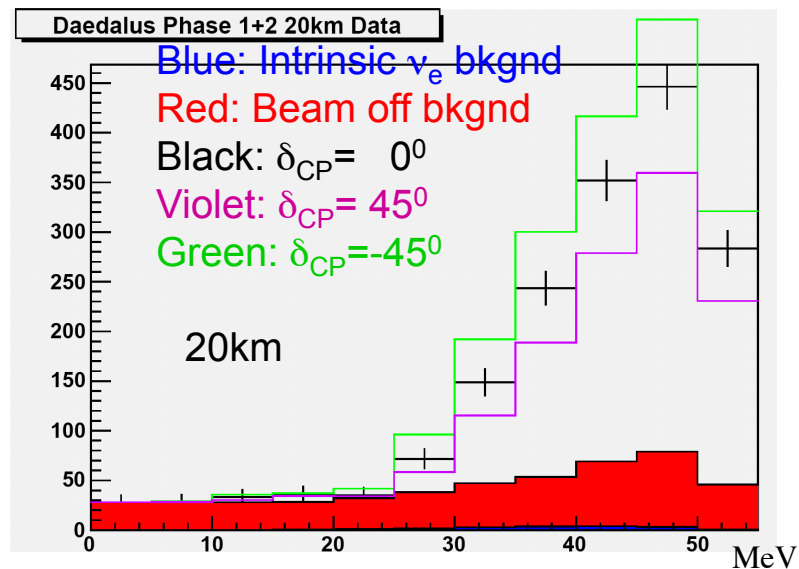
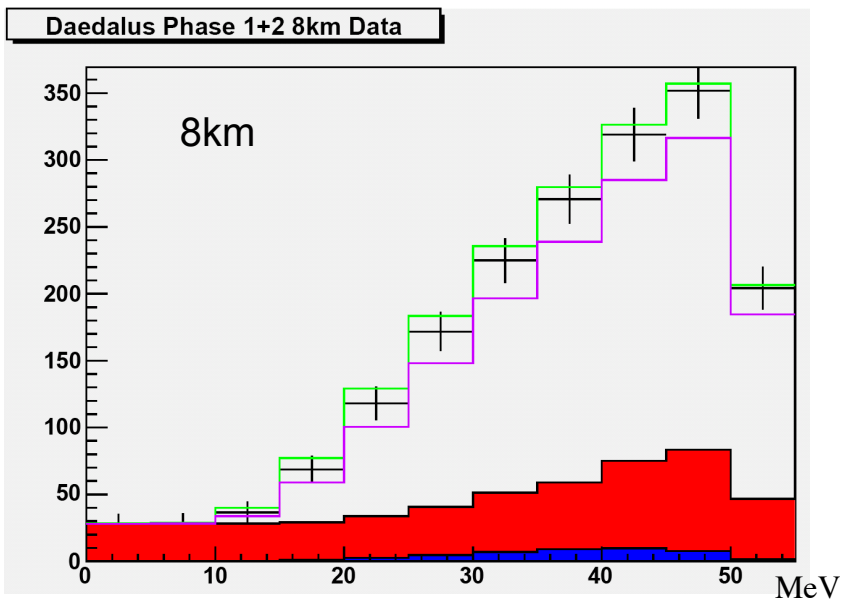
Daedalus Event Energy Distributions (Signal & Background)

$$(\sin^2 2\theta_{13} = 0.04)$$

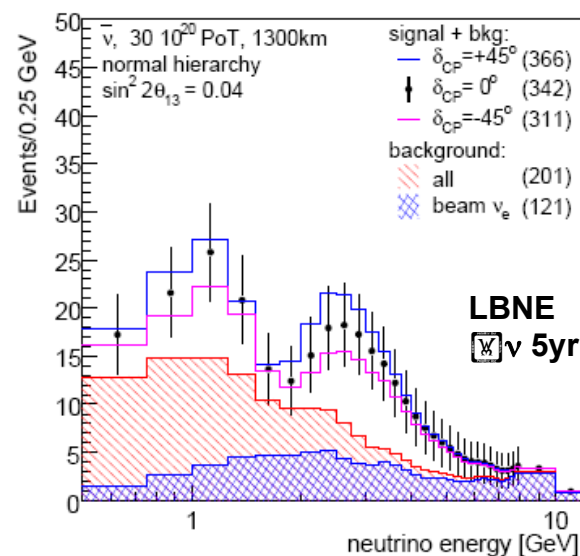
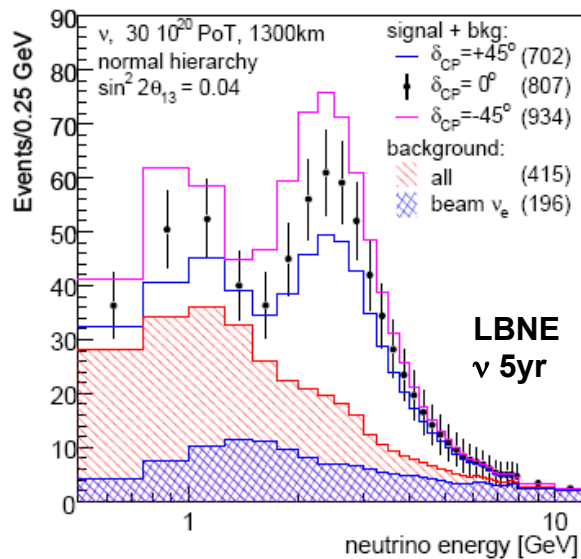
beam off
beam on



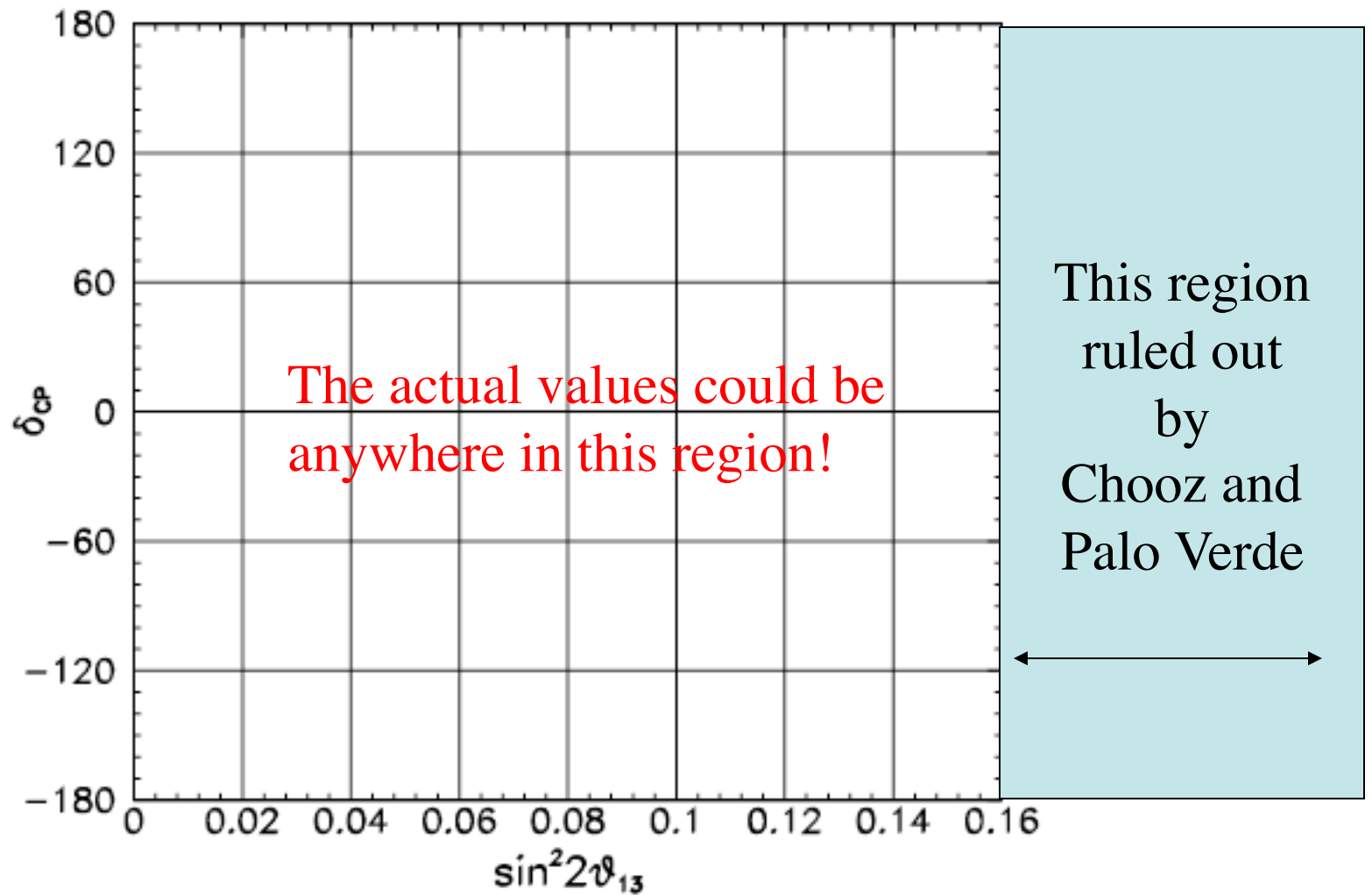
Compare signal to-background



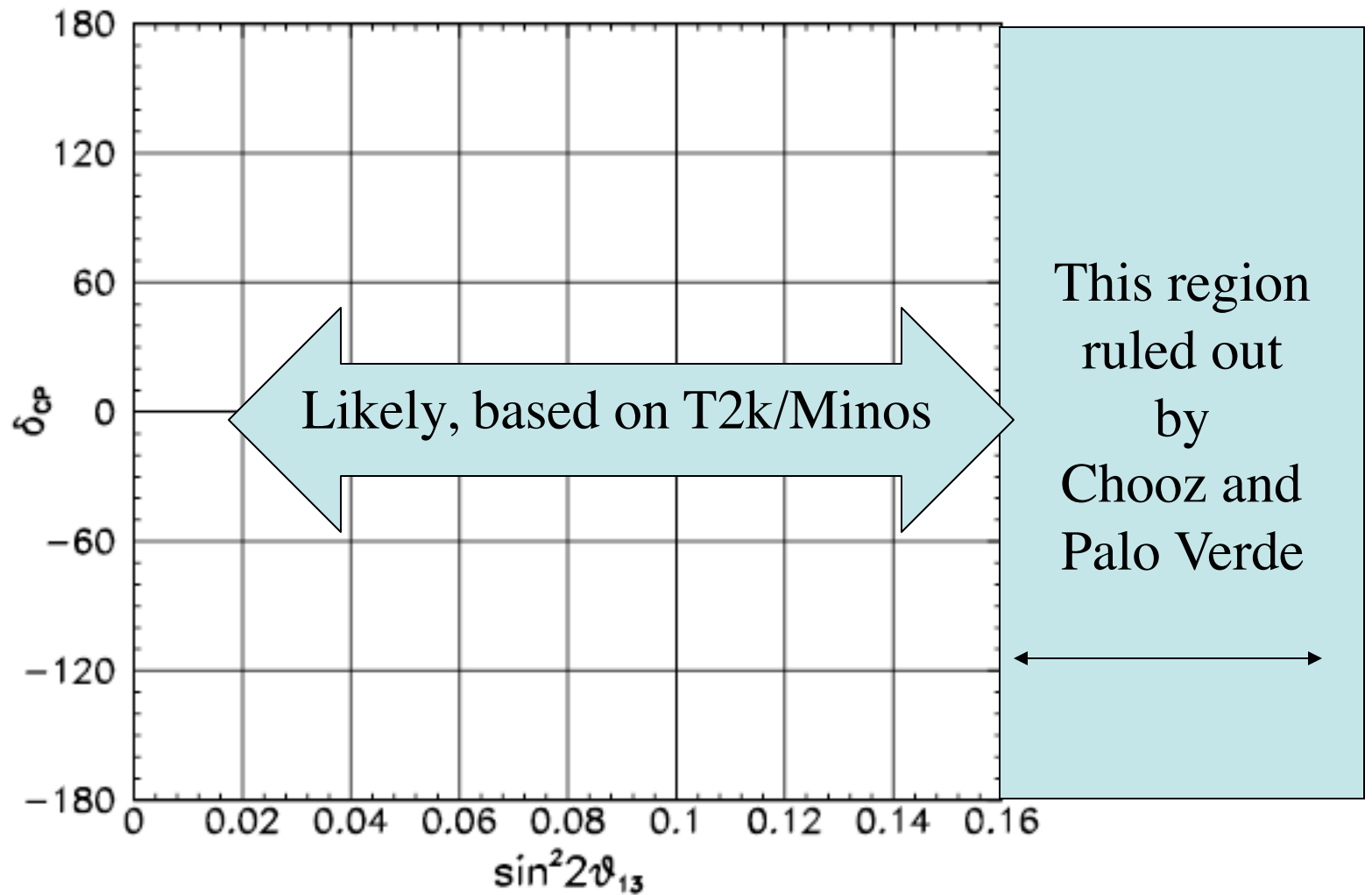
With LBNE...



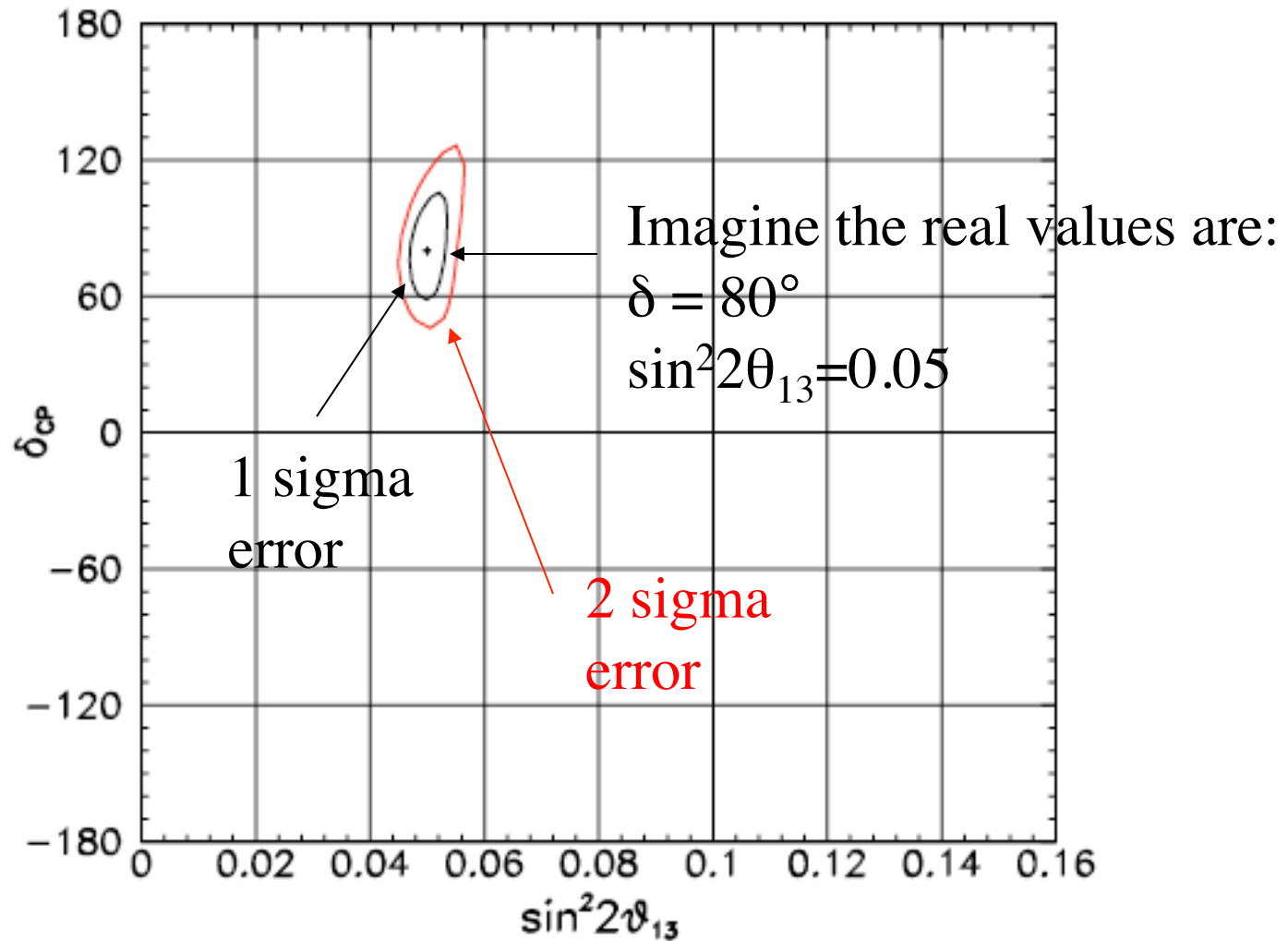
So what will we learn about δ vs θ_{13} from DAE δ ALUS??



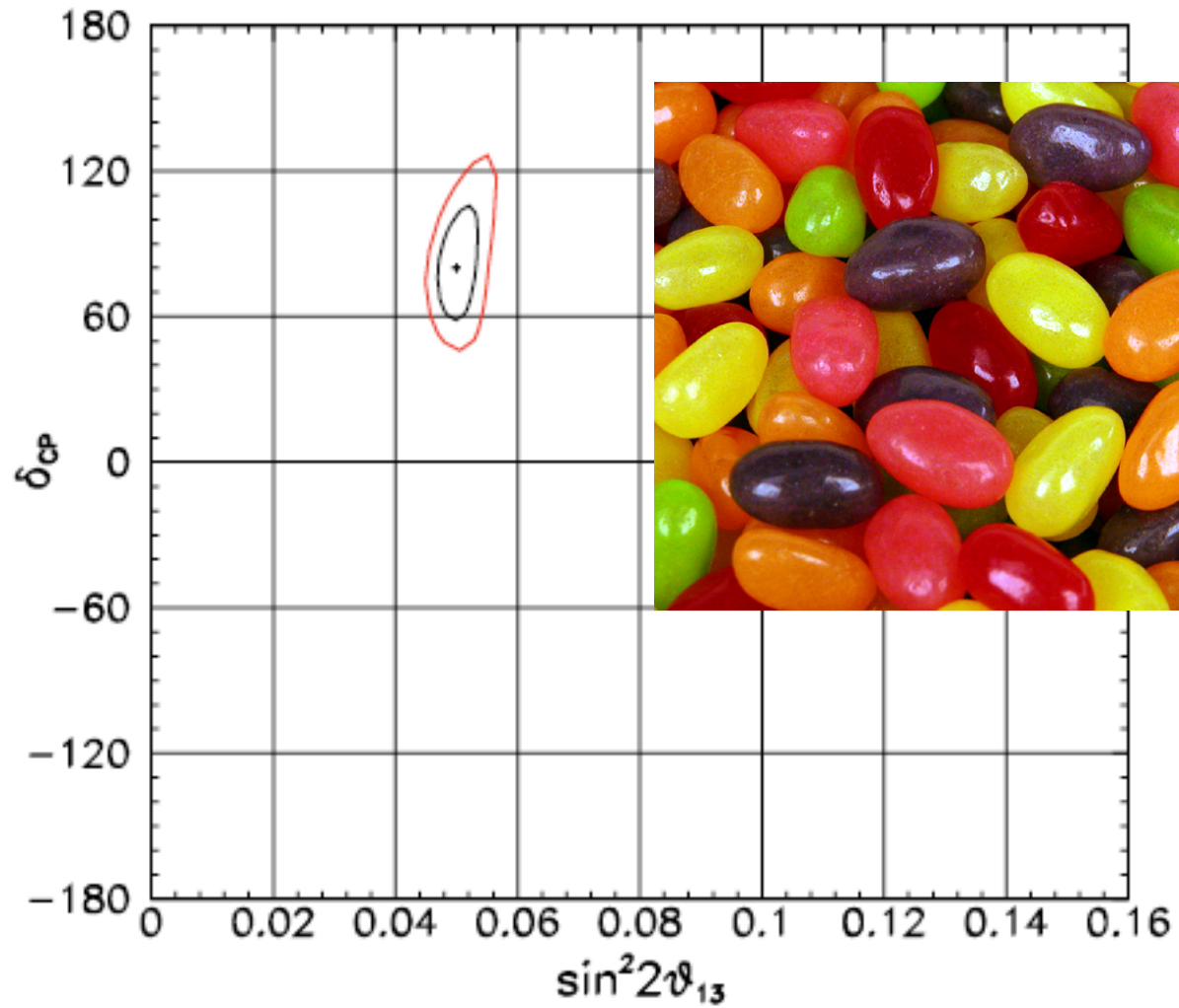
So what will we learn about δ vs θ_{13} from DAE δ ALUS??



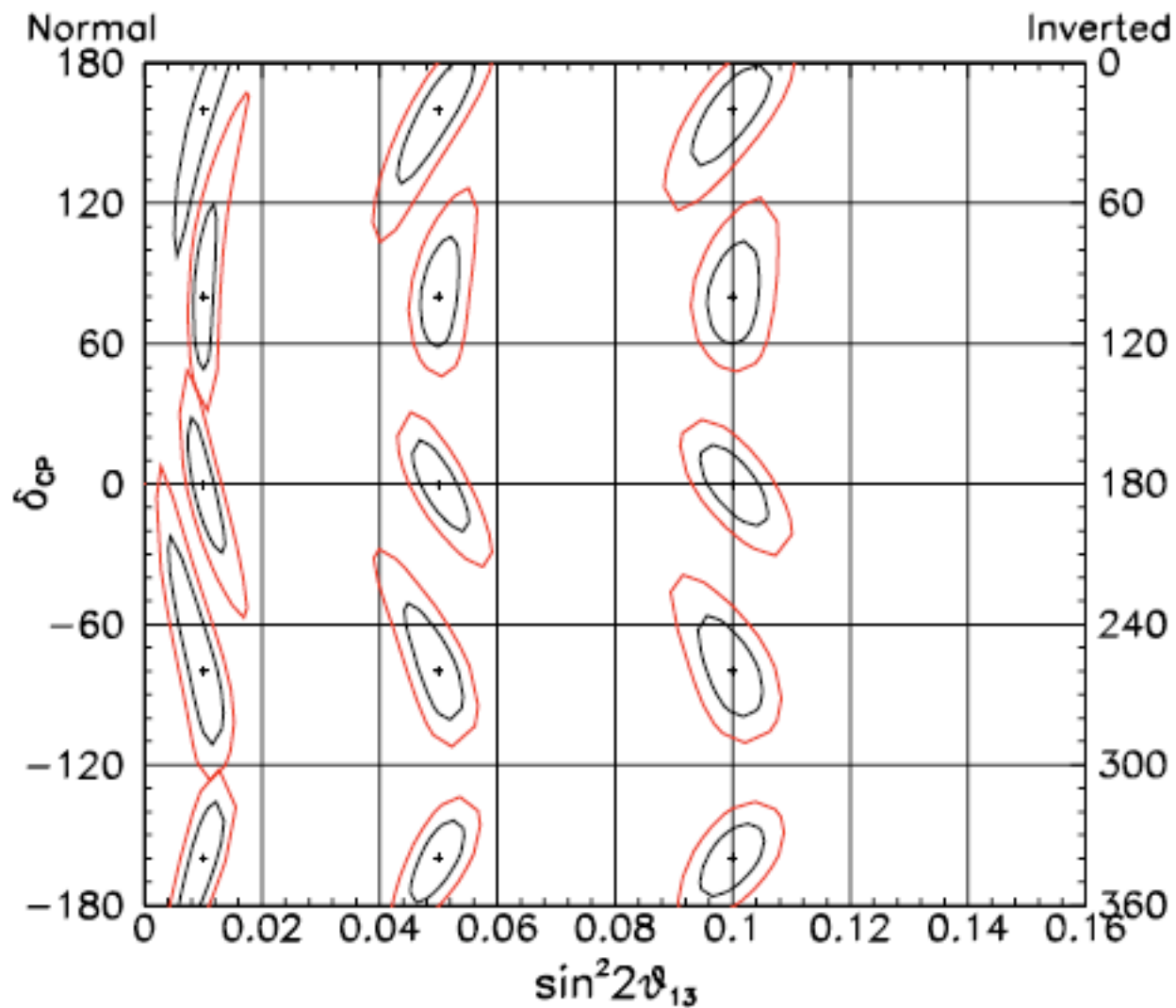
If we succeeded in observing a signal,
what would this plot look like?



You get a “jelly bean”

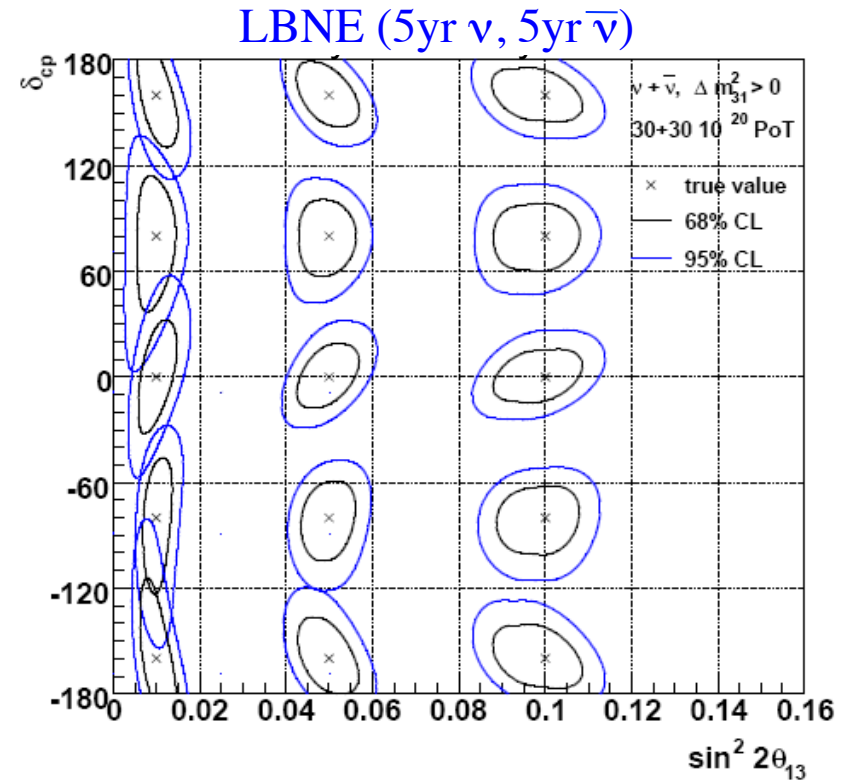
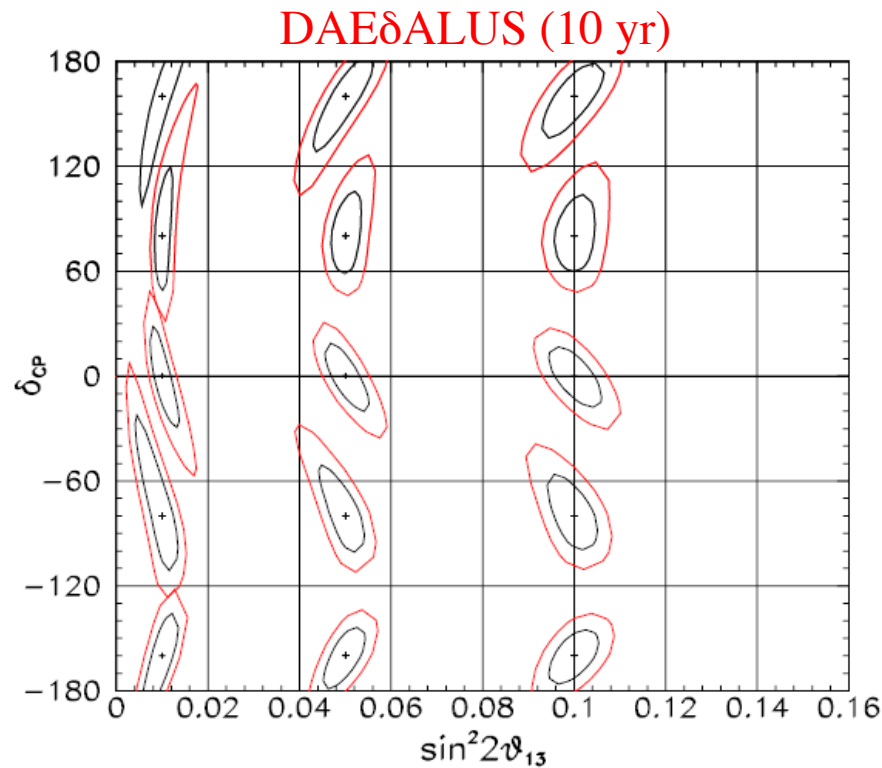


How well do we do on a “jelly bean” plot?



We can clearly observe CP violation!

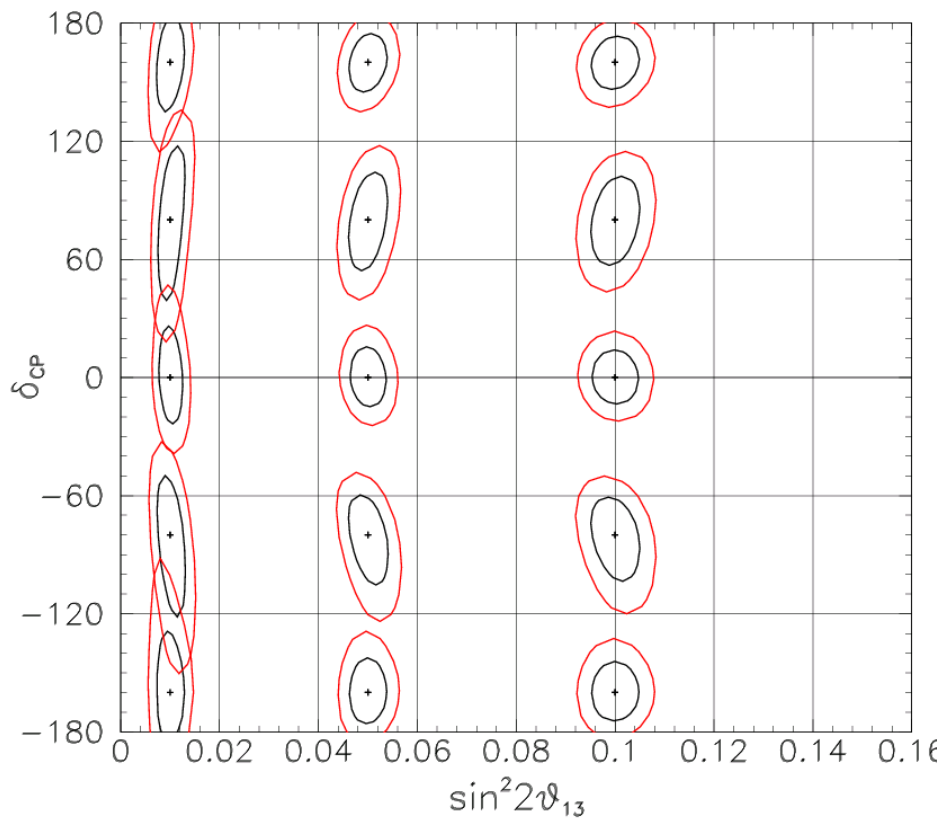
By construction our capability is equal to LBNE,
 With same sized detector
 But our measurement has completely different issues!



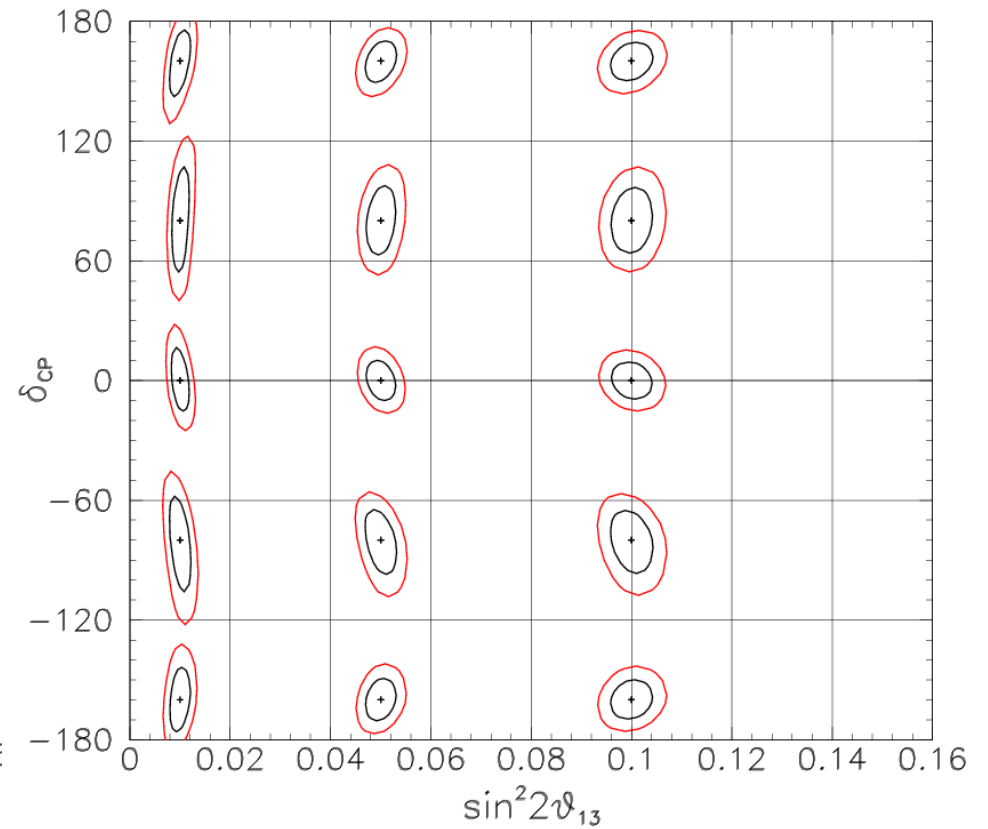
Normal Hierarchy

What the Combined Experiments can do!

5yr Combined Running

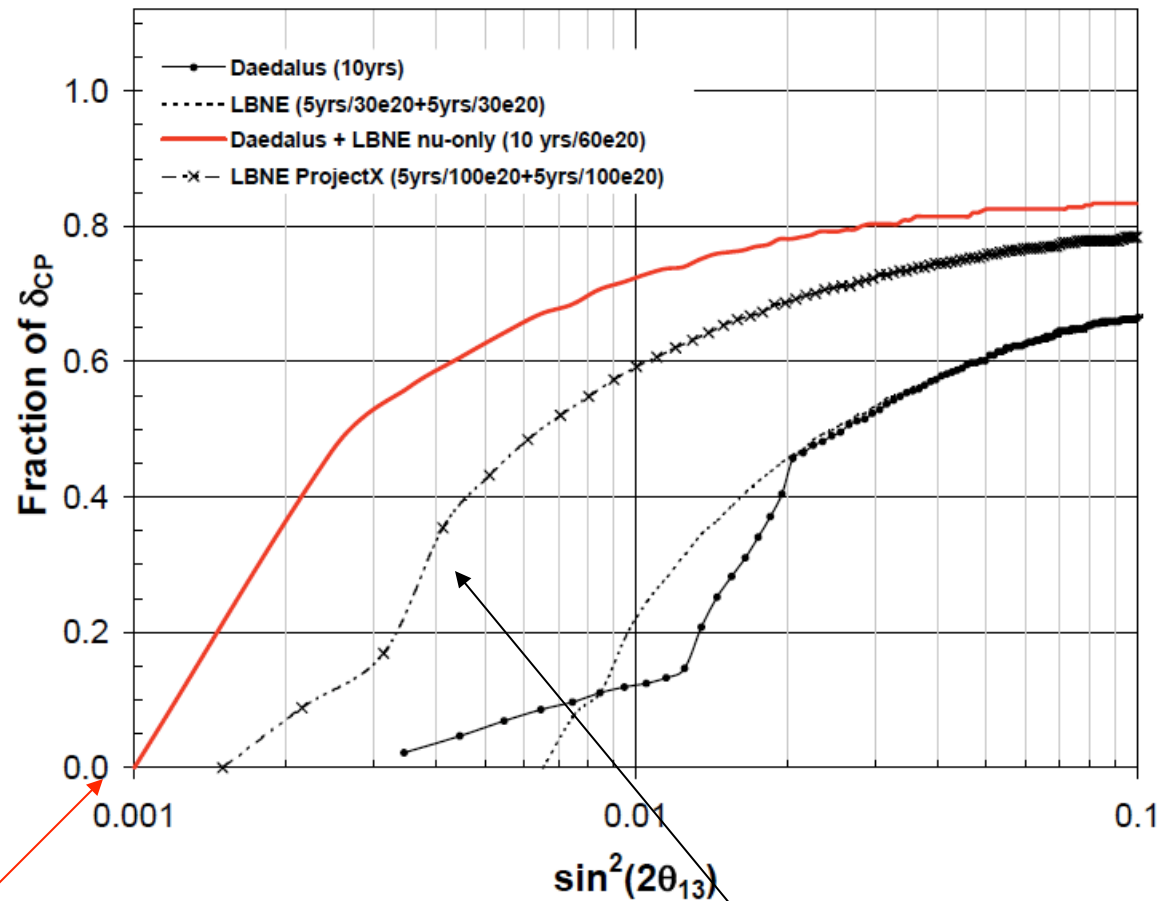


10yr Combined Running



The fraction of “ δ -space” where a measurement will be $>3\sigma$

Exclusion of $\delta_{CP} = 0^0$ or 180^0 at 3σ
(300kt Water Cherenkov for 10 year runs)



Better than a Project X (“superbeam”) experiment!

I have used a US-based example,
but this can be done anywhere you have

1. A detector with a lot of free protons (~ 100 s of ktons).
A scintillator detector should work too.
→ This is enough to match the conventional beam designs
2. A conventional beam at a reasonable distance.
→ This allows you to probe beyond the sensitivity of superbeams!

There is a big wide world of opportunities
for this design!

θ_{13}

Reactor

DAEdALUS

δ_{cp}

Long
Base-line
appearance

Hierarchy

Conclusions:

We have learned a lot in the last 10 years!
But exciting questions remain open in the field of oscillations.

We are developing a strong set of experiments to
go after these mysteries

Asking, in different ways...

ν_e or not ν_e ?

... I hope DAE δ ALUS will be one of them!