

Dark Matter and the Higgs Boson Discovery

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L. Roszkowski, APC, 14/3/2014

Outline

- ✧ Introduction – DM: evidence and general properties
- ✧ Theory frameworks for DM candidates
- ✧ SUSY neutralino as DM
- ✧ Implications of $m_h \sim 126$ GeV and direct limits on SUSY:
 - New DM candidate: ~ 1 TeV higgsino
- ✧ Prospects for detection
- ✧ Other recent developments and claims
- ✧ Summary



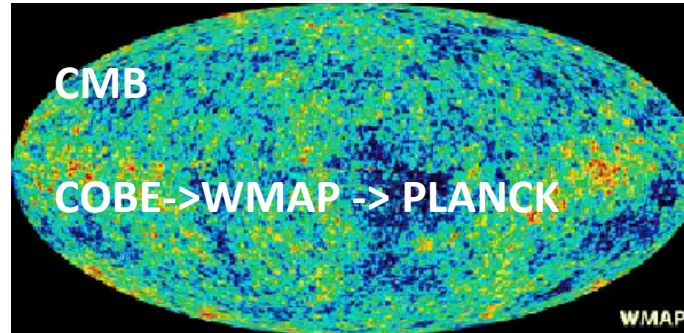
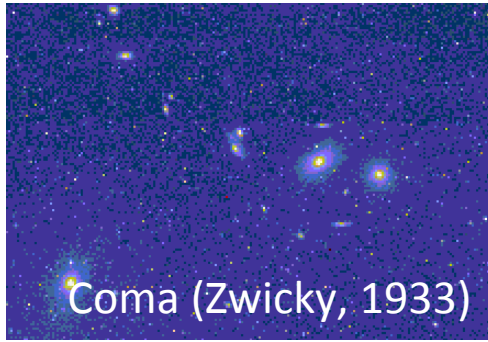
There is more out there
than meets the eye



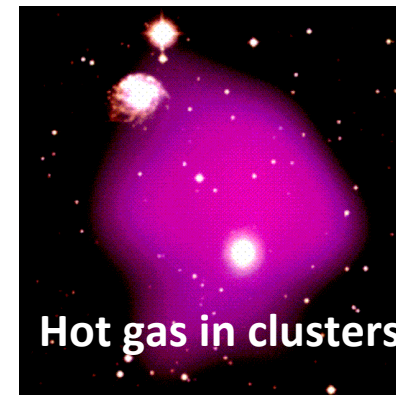
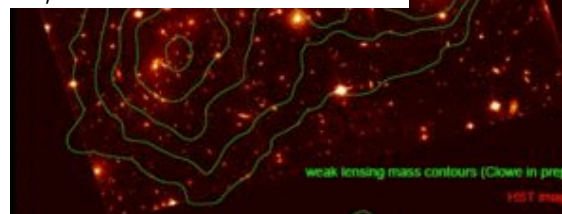
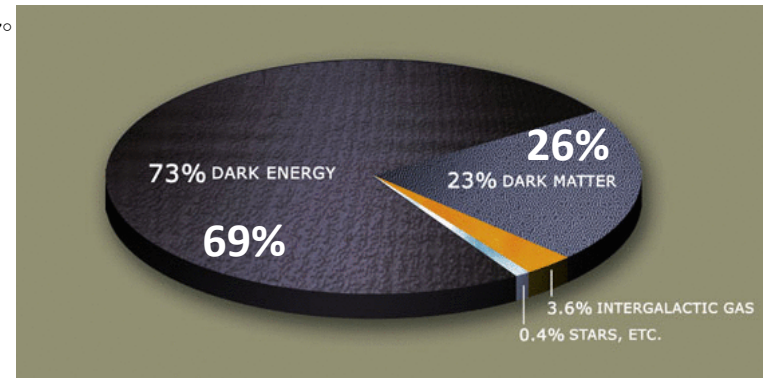
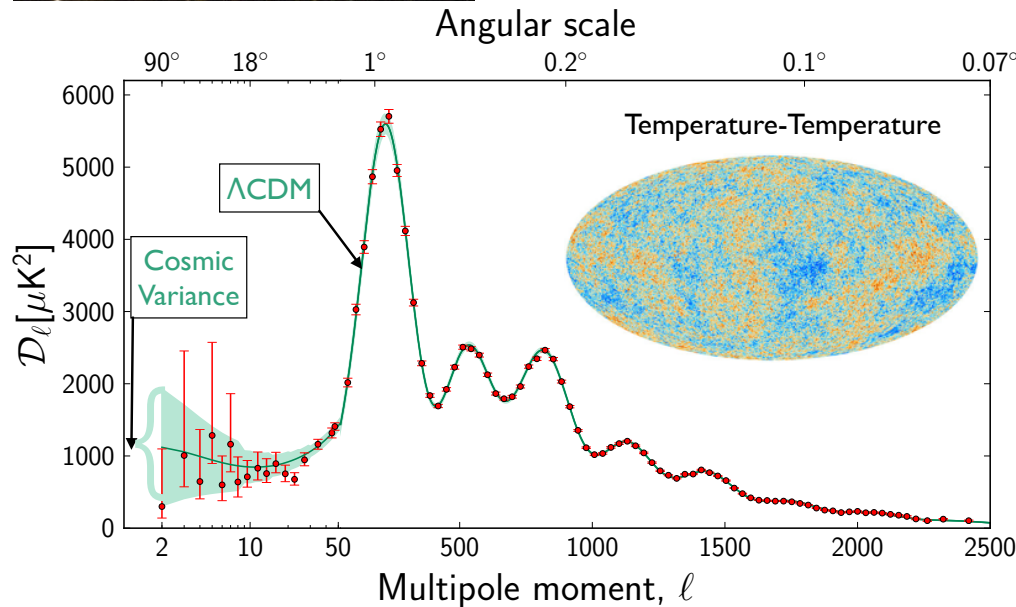
The WIMP Reigns

...but remains elusive

Footprints of Dark Matter



...felt but not seen



What is the DM?

- non-baryonic
- cold (CDM)
 - or possibly (?) warmish
- no electric nor (preferably) color interactions
- relic from the Big Bang
- element of some sensible particle theory



plausible choice \Rightarrow **WIMP**

(weakly interacting massive particle)

...a very broad class, not a single candidate

...How weak can weak be?

WIMP: most likely an unknown particle

A simple, persuasive argument:

- WIMPs decouple from thermal equilibrium
- freeze-out when $\Gamma \lesssim H$

WIMP relic abundance

$$\Omega h^2 \simeq \frac{1}{\left\langle \left(\frac{\sigma_{\text{ann}}}{10^{-38} \text{cm}^2} \right) \left(\frac{v/c}{0.1} \right) \right\rangle}$$

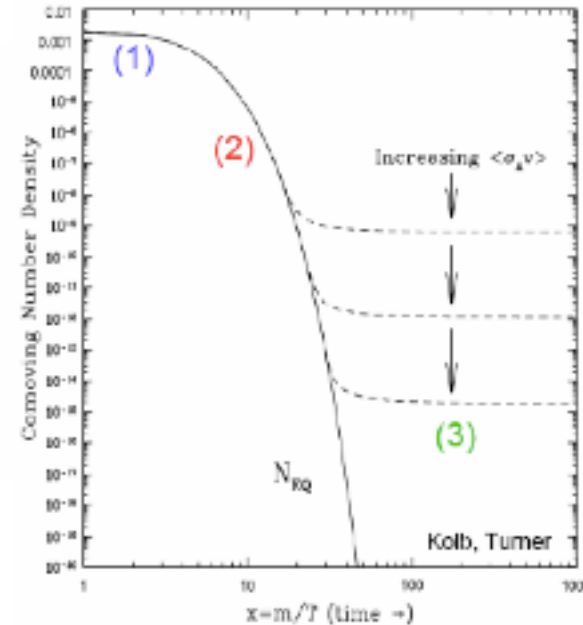
σ_{ann} – c.s. for WIMP pair-annihilation in the early Universe

v – their relative velocity, $\langle \dots \rangle$ – thermal average

$$\sigma_{\text{ann}} \sim \sigma_{\text{weak}} \sim 10^{-38} \text{cm}^2 = 10^{-2} \text{pb} \Rightarrow \Omega h^2 \sim 1$$

A hint? Possibly, but...

Not “WIMP Miracle” but weak int. – relic density coincidence



CDM: some theory frameworks

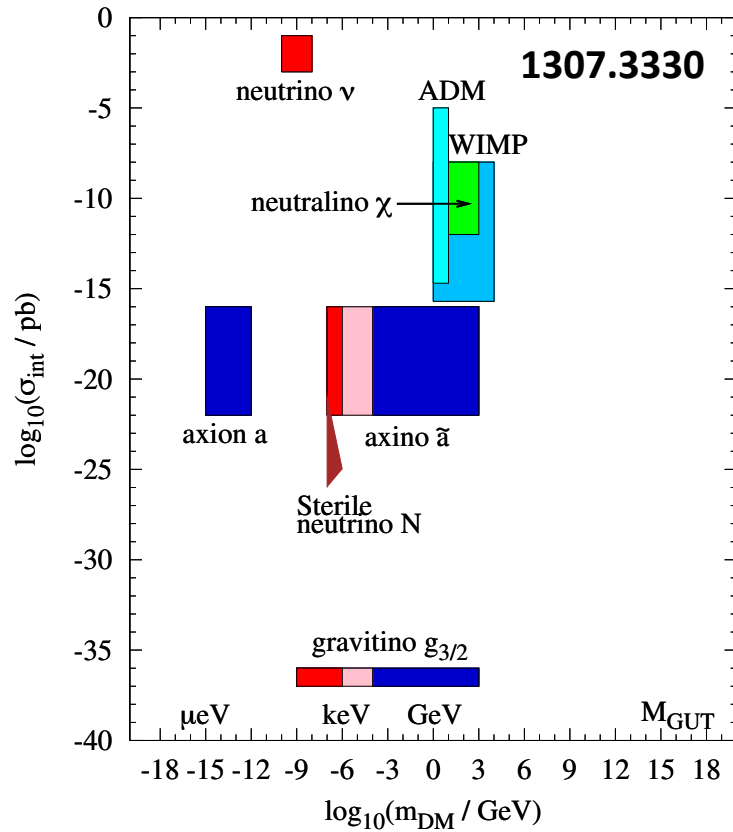
- ✧ **SUSY** <- by far most popular (and best motivated)
- ✧ **DM and various extensions of the SM (portals/hidden valleys,...)**
- ✧ **Asymmetric DM**
- ✧ **Self-interacting DM**
- ✧ **Universal extra dim's**
- ✧ **...**

SUSY frameworks:

- ❖ **Unified (=GUT-constrained) models (Constrained MSSM, ...)**
 - virtues: well-motivated, predictive, realistic
 - limitations: may miss some solutions
- ❖ **Phenomenological (supersymmetrized SM)**
- ❖ **Scenarios motivated by this or that...**



Well-motivated candidates for dark matter



● neutrino ν – hot DM

● neutralino χ

● “generic” WIMP

● axion a

● axino \tilde{a}

● gravitino \tilde{G}

S
U
S
Y

- vast ranges of interactions and masses
- different production mechanisms in the early Universe (thermal, non-thermal)
- need to go beyond the Standard Model
- **WIMP candidates testable at present/near future**
- axino, gravitino EWIMPs/superWIMPs not directly testable, but some hints from LHC

Where is the WIMP?

- **Mass range: at least 20 orders of magnitude**
- **Interaction range: some 32 orders of magnitude**



© Ron Leishman * www.ClipartOf.com/1047187

Strategies for WIMP Detection

- **direct detection (DD)**: measure WIMPs scattering off a target

go underground to beat cosmic ray bgnd

- **indirect detection (ID)**:

- **HE neutrinos from the Sun (or Earth)**

WIMPs get trapped in Sun's core, start pair annihilating, only ν 's escape

- **antimatter (e^+ , \bar{p} , \bar{D}) from WIMP pair-annihilation in the MW halo**

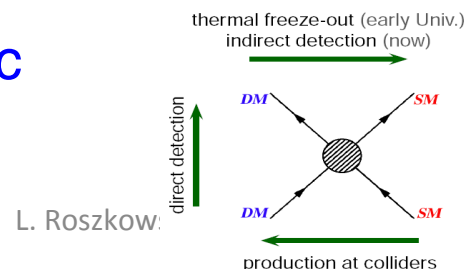
from within a few kpc

- **gamma rays from WIMP pair-annihilation in the Galactic center**

depending on DM distribution in the GC

- **other ideas: traces of WIMP annihilation in dwarf galaxies, in rich clusters, etc**

- **the LHC**



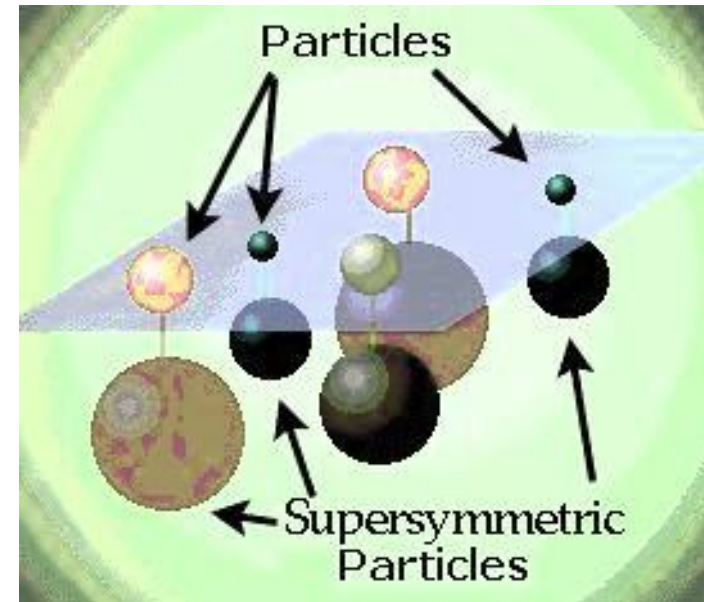
more speculative

Supersymmetry



Symmetry among particles

bosons \leftrightarrow fermions



Supersymmetric dark matter?

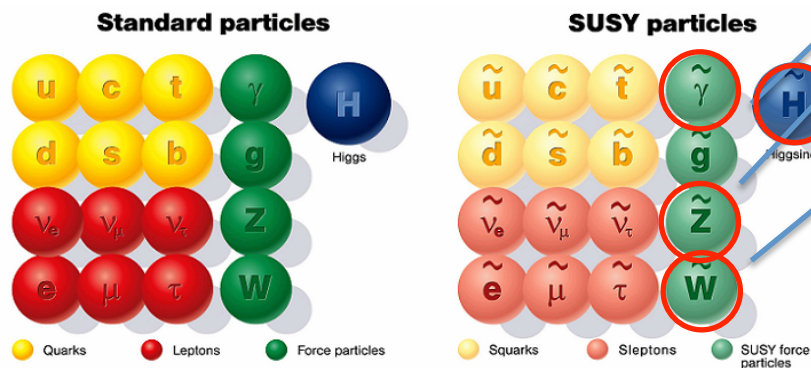
LSP – Lightest SUSY particle:

- **Weakly interacting**
Neutral (electric+color)
- **Massive**
- **Stable (R-parity)**

Possible candidates for LSP:

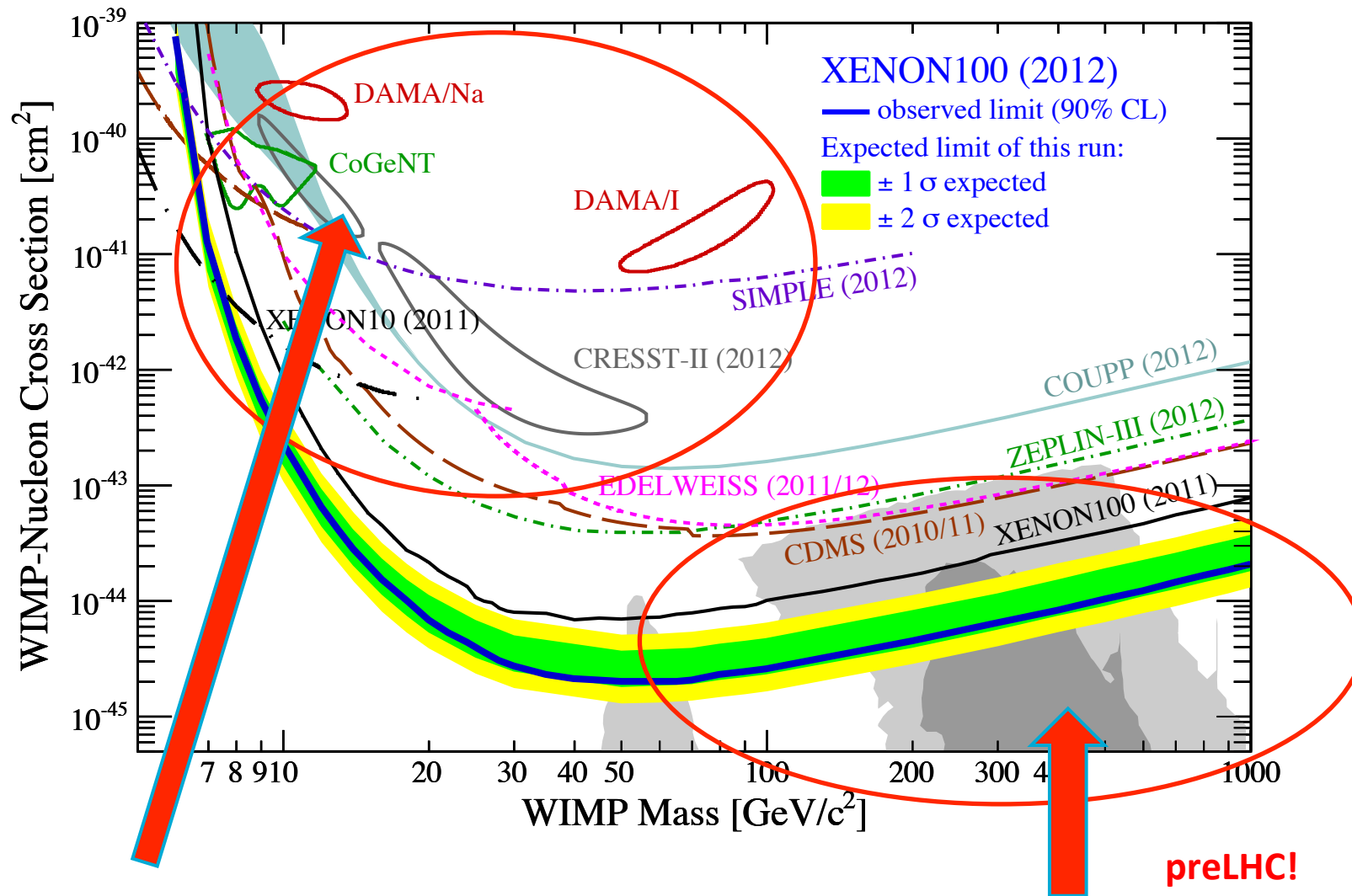
- **Part of ordinary SUSY spectrum:**
Neutralino: mass state of bino, wino, higgsinos

Sneutrino – not good (LEP, DM searches)

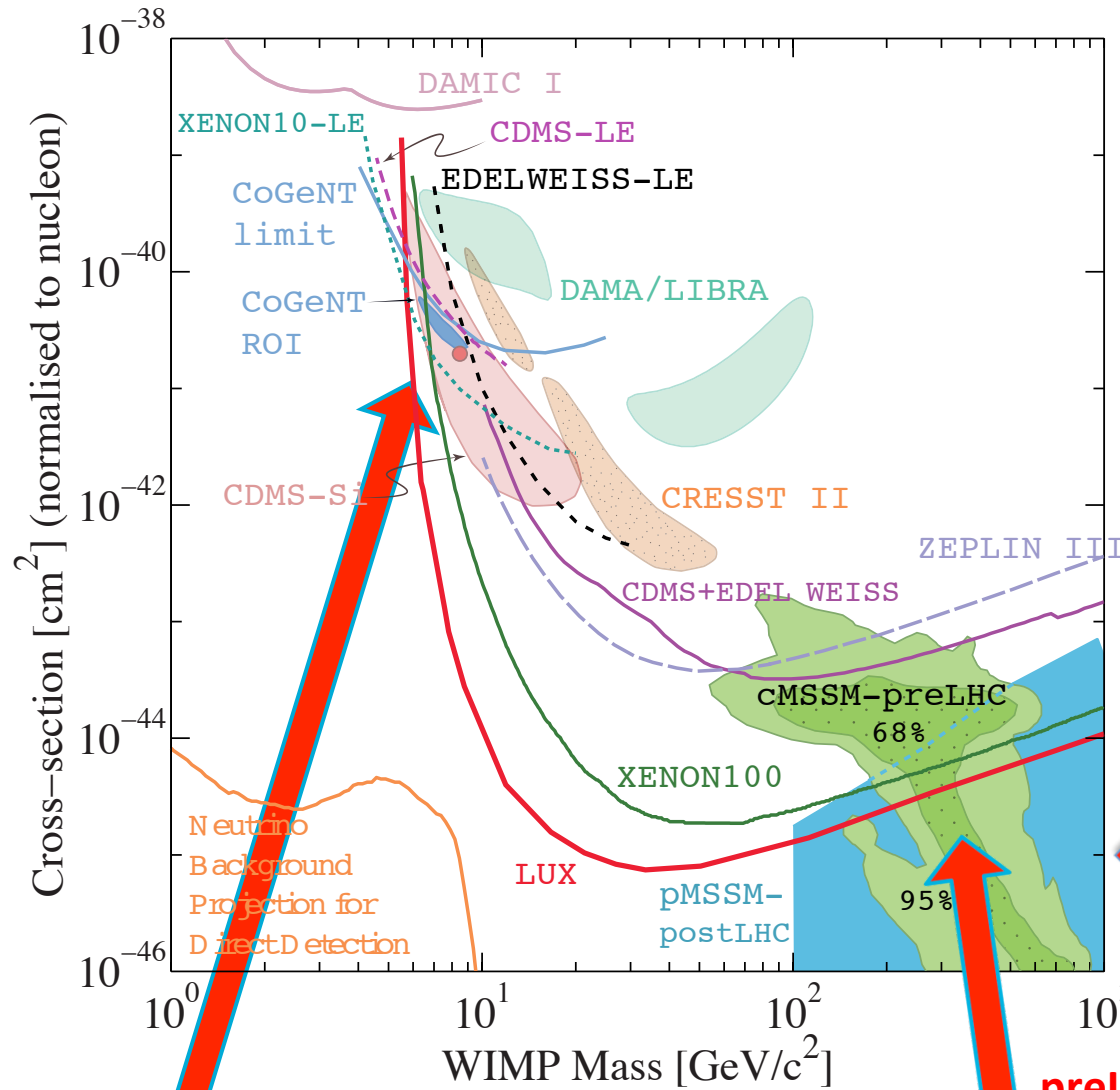


- Add gravity: **gravitino** LSP
- Add axion: **axino** LSP

Direct Detection AD 2011 - Before LHC



Direct Detection Nov. 2013



PDG update 2013
(1204.2373)

LHC:
theory region has
moved down and
right

in a very specific way

**Smoking gun
of SUSY?**

preLHC!

motivated by theory (SUSY)

Confusion region gone

Main news from the LHC so far...

➤ SM-like Higgs particle at ~ 126 GeV

➤ No (convincing) deviations from the SM

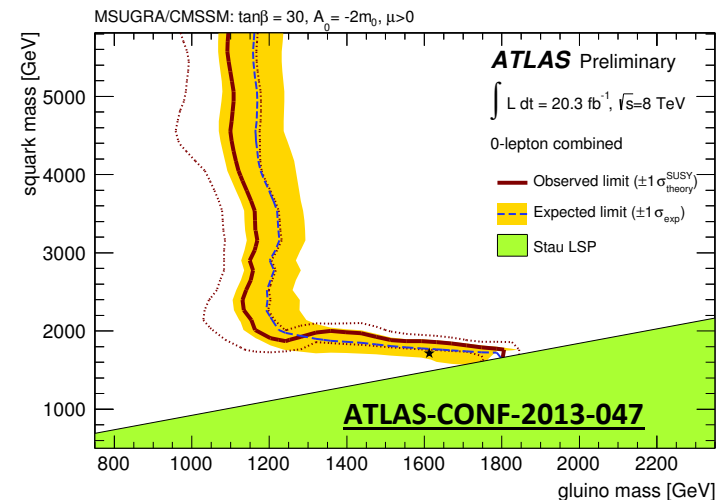
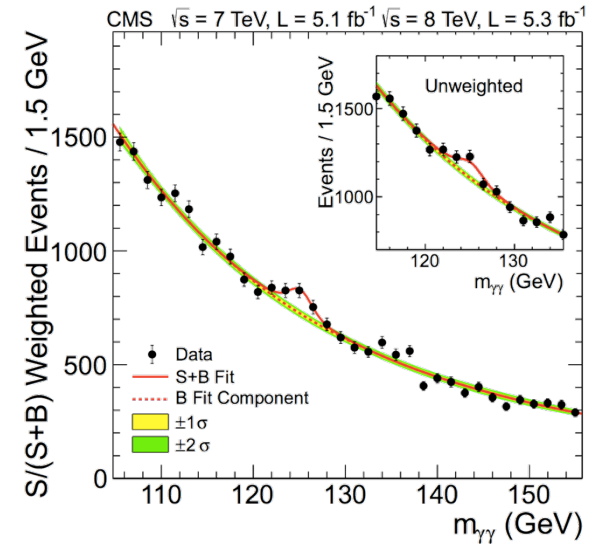
$$\text{BR}(B_s \rightarrow \mu^+ \mu^-)_{\text{LHCb}} = (2.9_{-1.0}^{+1.1}) \times 10^{-9}$$

$$\text{BR}(B_s \rightarrow \mu^+ \mu^-)_{\text{CMS}} = (3.0_{-0.9}^{+1.0}) \times 10^{-9}$$

$$\text{BR}(B_s \rightarrow \mu^+ \mu^-)_{\text{SM}} = (3.65 \pm 0.23) \times 10^{-9}$$

➤ Stringent lower limits on superpartner masses

SUSY masses pushed to 1 TeV+ scale...



...and from the media...

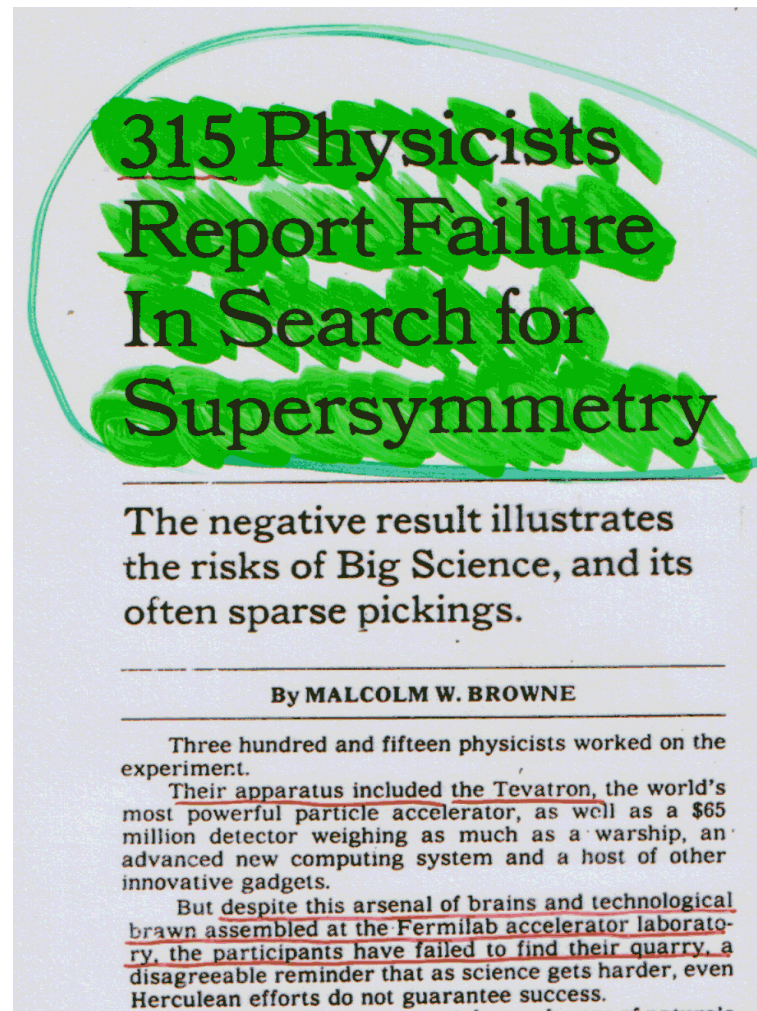
Is Supersymmetry Dead?

The grand scheme, a stepping-stone to string theory, is still high on physicists' wish lists. But if no solid evidence surfaces soon, it could begin to have a serious PR problem

**SCIENTIFIC
AMERICAN™**

April 2012

Nothing new...



CDF, ~2003

The 126 GeV SM-Like Higgs Boson

A blessing or a curse for SUSY?

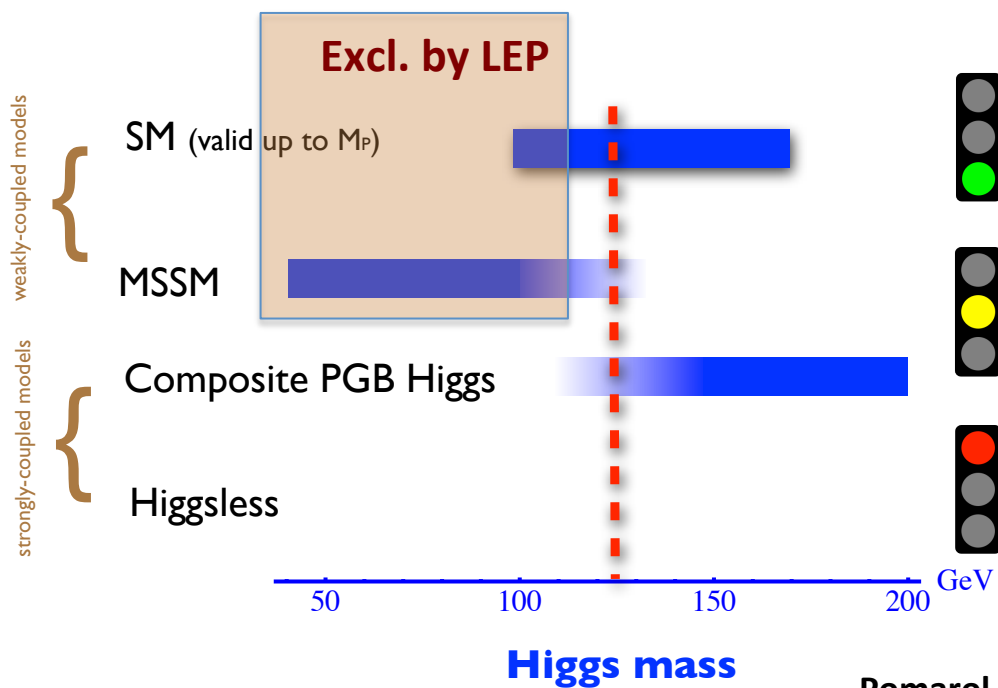
The 126 GeV Higgs Boson and SUSY

A blessing...

- Fundamental scalar --> SUSY
- Light and SM-like --> SUSY

Low energy SUSY prediction:
Higgs mass up to ~135 GeV

Constrained SUSY prediction:
SM-like Higgs with mass
up to ~130 GeV

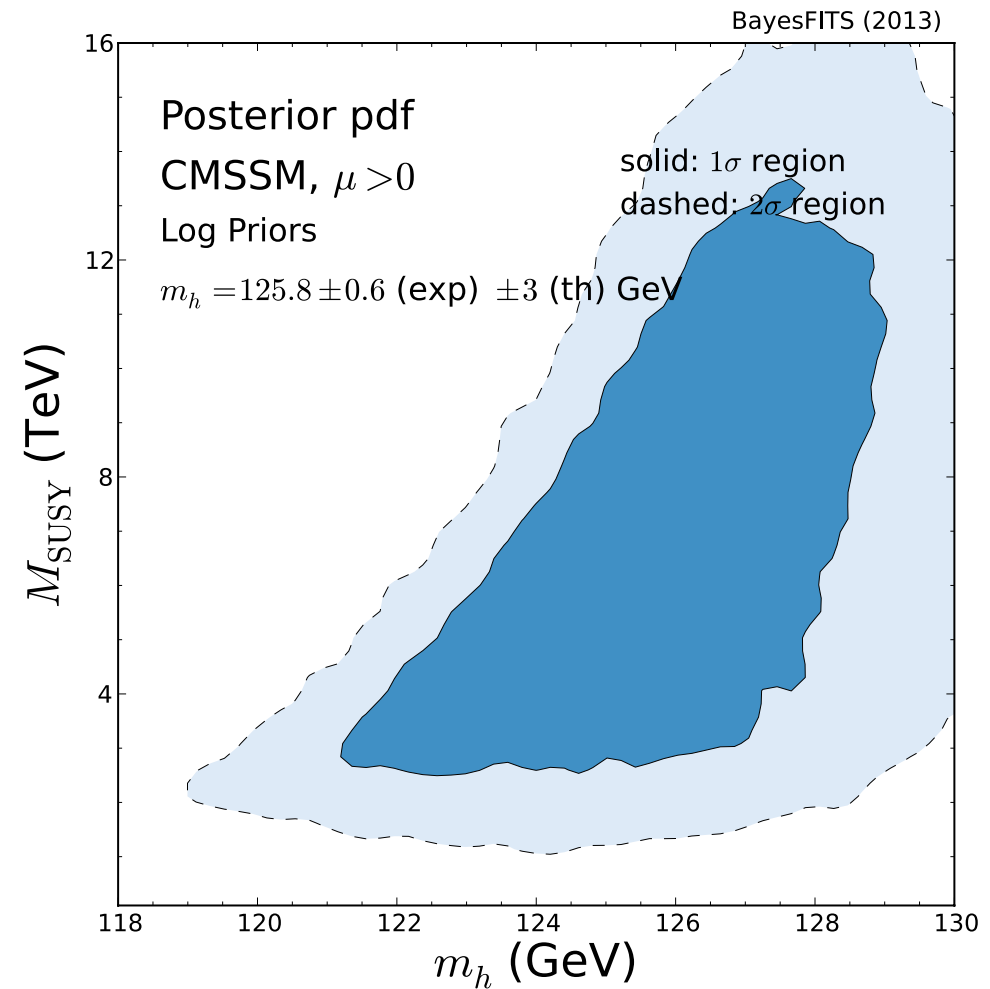


The 126 GeV Higgs Boson and SUSY

A curse...

Only $m_h \sim 126$ GeV and CMS lower bounds on SUSY applied.

**126 GeV Higgs->
Multi-TeV SUSY**



Constrained Minimal Supersymmetric Standard Model (CMSSM)

G. L. Kane, C. F. Kolda, L. Roszkowski and J. D. Wells, Phys. Rev. D 49 (1994) 6173

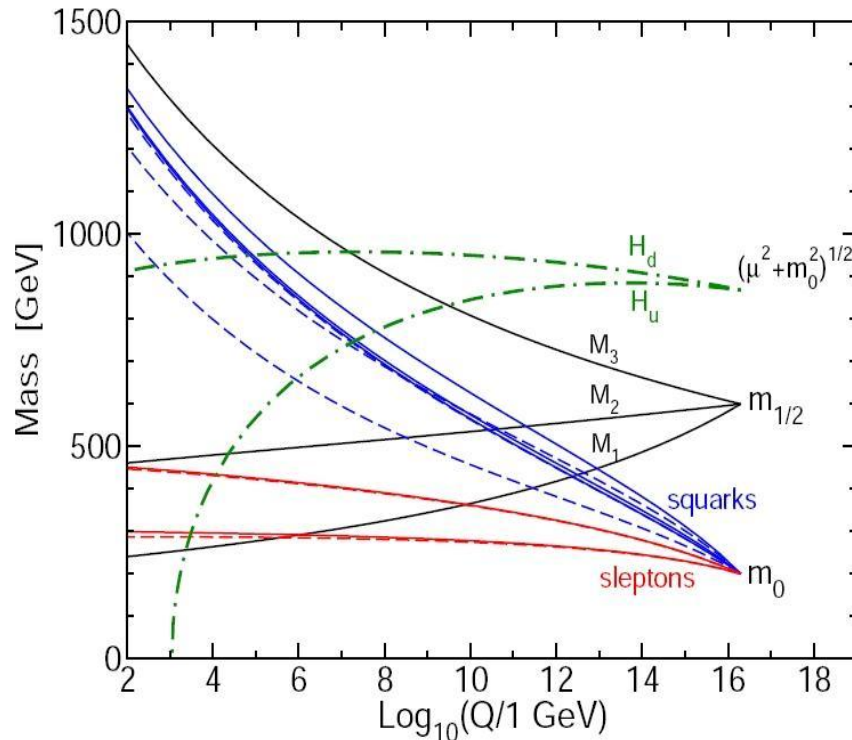


figure from hep-ph/9709356

At $M_{\text{GUT}} \simeq 2 \times 10^{16}$ GeV:

- gauginos $M_1 = M_2 = m_{\tilde{g}} = m_{1/2}$
- scalars $m_{\tilde{q}_i}^2 = m_{\tilde{l}_i}^2 = m_{H_b}^2 = m_{H_t}^2 = m_0^2$
- 3-linear soft terms $A_b = A_t = A_0$
- radiative EWSB
$$\mu^2 = \frac{m_{H_b}^2 - m_{H_t}^2 \tan^2 \beta}{\tan^2 \beta - 1} - \frac{m_Z^2}{2}$$
- five independent parameters: $m_{1/2}, m_0, A_0, \tan \beta, \text{sgn}(\mu)$
- well developed machinery to compute masses and couplings



In general supersymmetric SM too many free parameters

~126 GeV Higgs in the CMSSM

- Include **only** $m_h \sim 126$ GeV **and** lower limits from direct SUSY searches

$$\mathcal{L} \sim e^{-\frac{(m_h - 125.8 \text{ GeV})^2}{\sigma^2 + \tau^2}}$$

$$\sigma = 0.6 \text{ GeV}, \tau = 2 \text{ GeV}$$

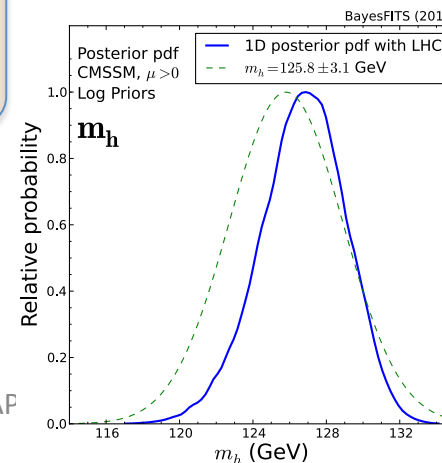
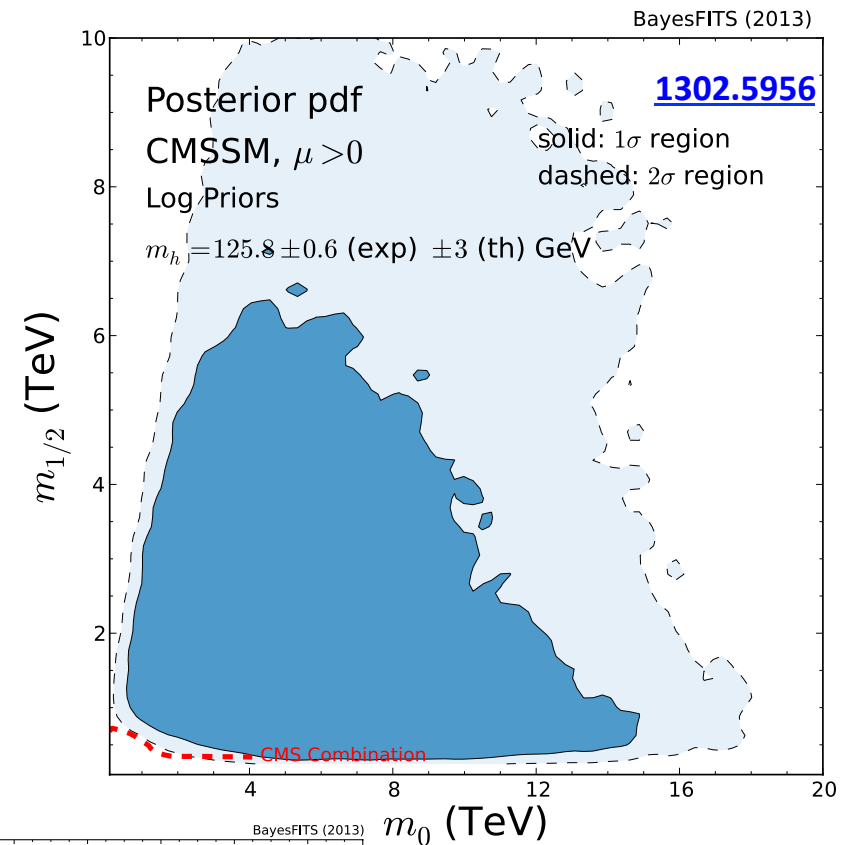
We use DR-bar approach (SoftSusy).
It gives larger m_h .

A curse...

~126 GeV Higgs mass implies multi-TeV scale for SUSY

Consistent with:

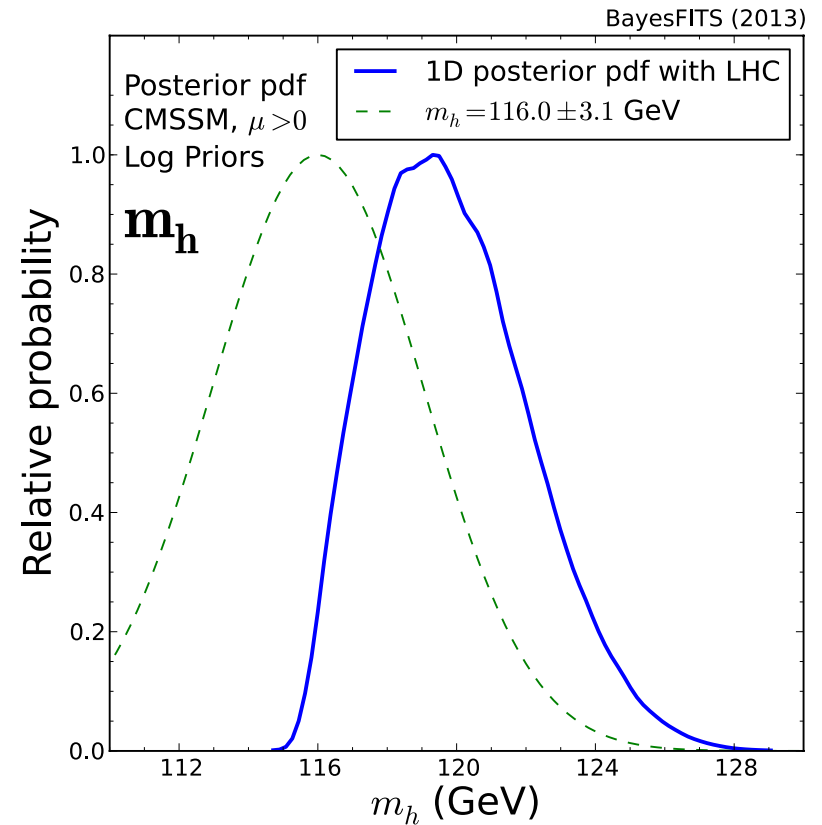
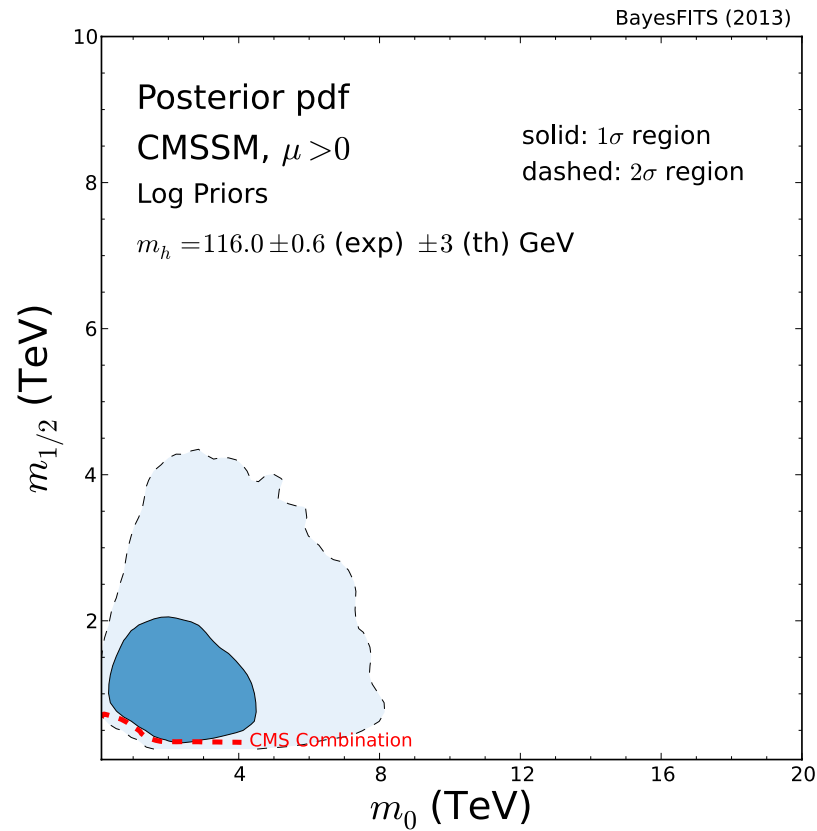
- SUSY direct search lower limits at LHC
- constraints from flavor



A weak upper bound on M_{SUSY}

...except at very small $\tan\beta > 1$ where it goes away

If m_h were, say, 116 GeV...

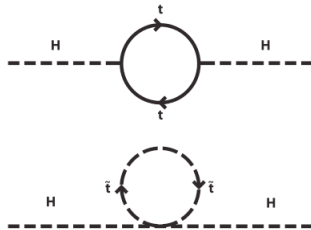


...significant tension with LHC bounds

Higgs boson mass and SUSY

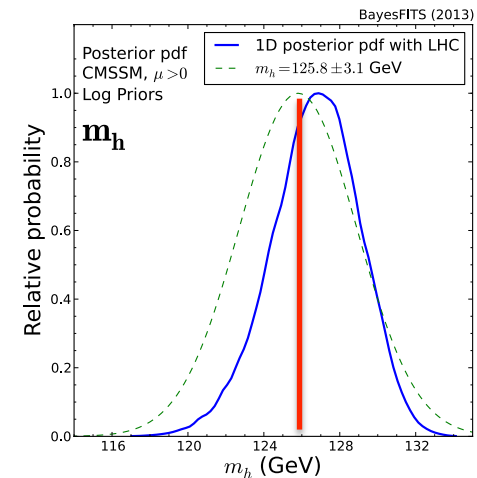
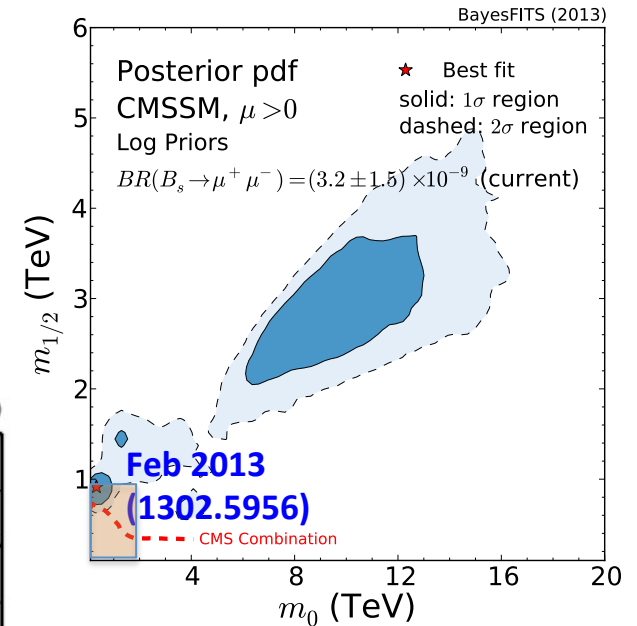
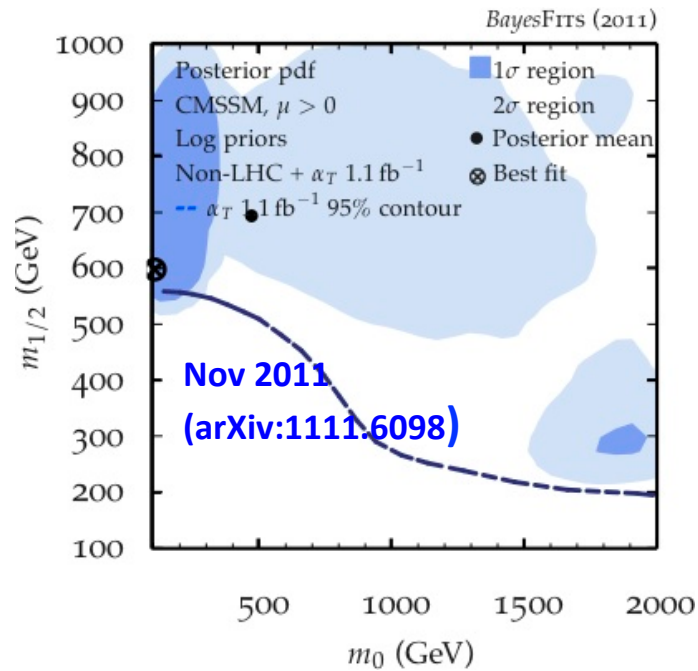
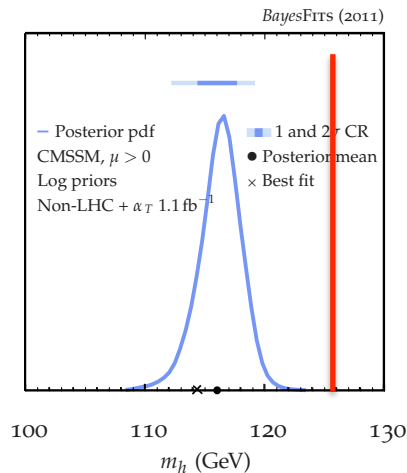
➤ 1 loop correction

$$\Delta m_h^2 = \frac{3m_t^4}{4\pi^2 v^2} \left[\ln \left(\frac{M_{\text{SUSY}}^2}{m_t^2} \right) + \frac{X_t^2}{M_{\text{SUSY}}^2} \left(1 - \frac{X_t^2}{12M_{\text{SUSY}}^2} \right) \right]$$



$$M_{\text{SUSY}} \equiv \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$$

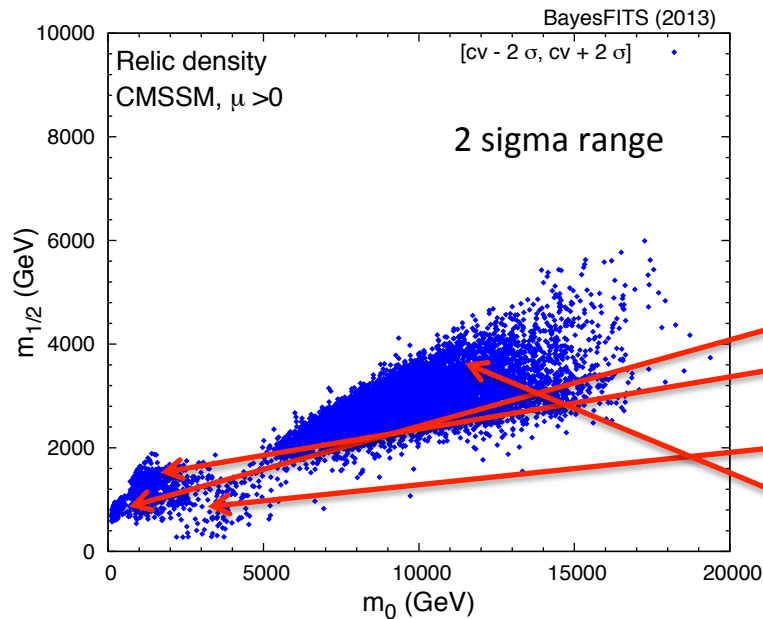
$$X_t = A_t - \mu \cot \beta$$



Higgs mass plus dark matter density

- **Unified SUSY: relic density is typically 1-2 orders of magnitude too large**
- **Effect of LHC direct SUSY limits much weaker**

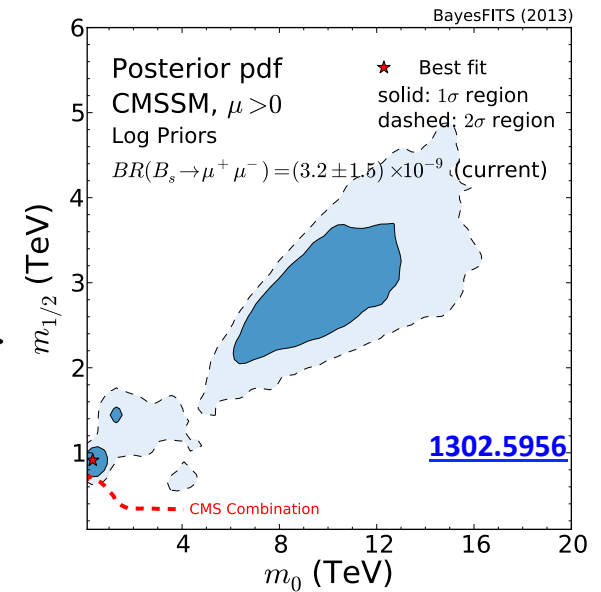
All constraints included -->



Remaining mechanisms of reducing it to correct range:

- ✧ neutralino-stau coannihilation
- ✧ pseudoscalar Higgs A resonance
- ✧ focus point/hyperbolic branch region
- ✧ **~1 TeV higgsino LSP at large MSUSY**
- ✧ and (very rare) LSP-stop coannihilation

Akula+Nath,
1207.1839
 $\Omega h^2 \propto m_A^4 / \tan^2 \beta$



Scan with **all** other relevant constraints imposed

CMSSM: these are the only DM-favored regions



~1 TeV higgsino DM

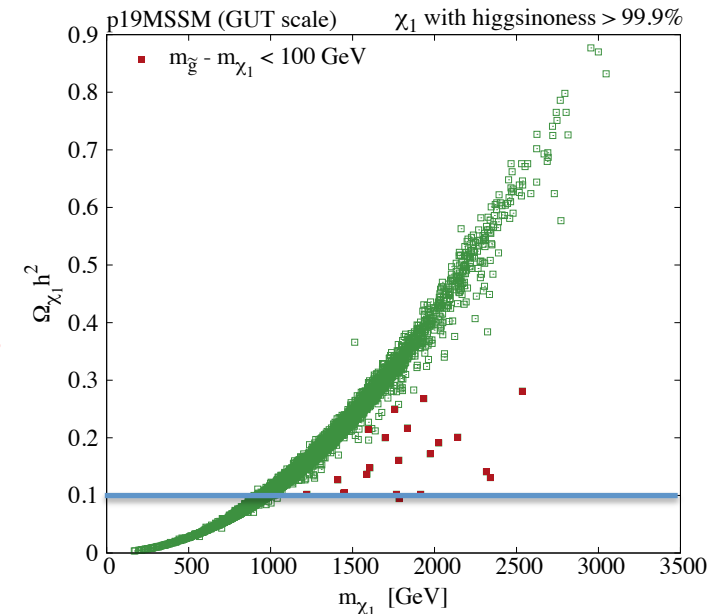
- ✧ Generic, present in many SUSY models
(both GUT-based and not)

Condition: gauginos heavy enough

When $m_{\tilde{B}} \gtrsim 1 \text{ TeV}$:
easiest to achieve $\Omega_{\chi} h^2 \simeq 0.1$
when $m_{\tilde{H}} \simeq 1 \text{ TeV}$

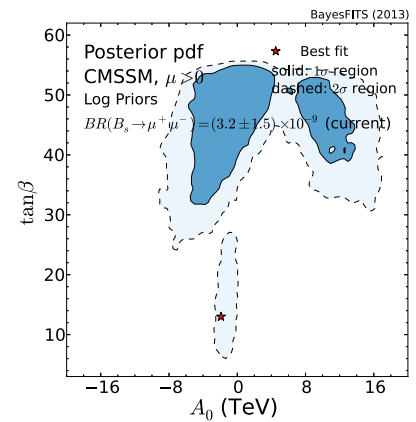
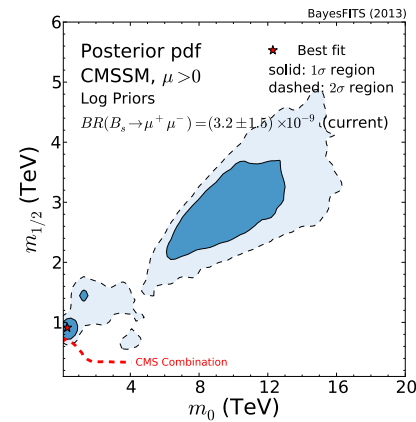
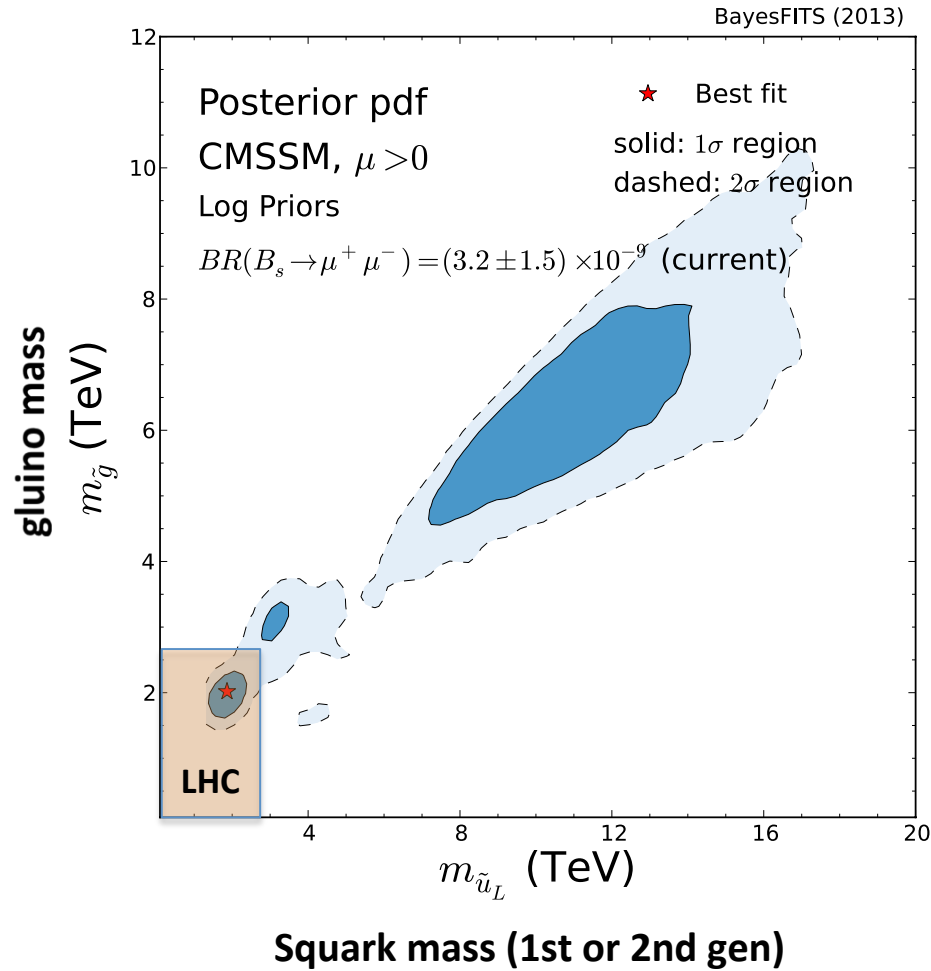
- ✧ Implied by ~126 GeV Higgs mass
and relic density
- ✧ Smoking gun of SUSY!?

Unified SUSY:
First pointed out
in NUHM in [0903.1279](#)



Can such multi-TeV ranges of SUSY parameters be experimentally tested?

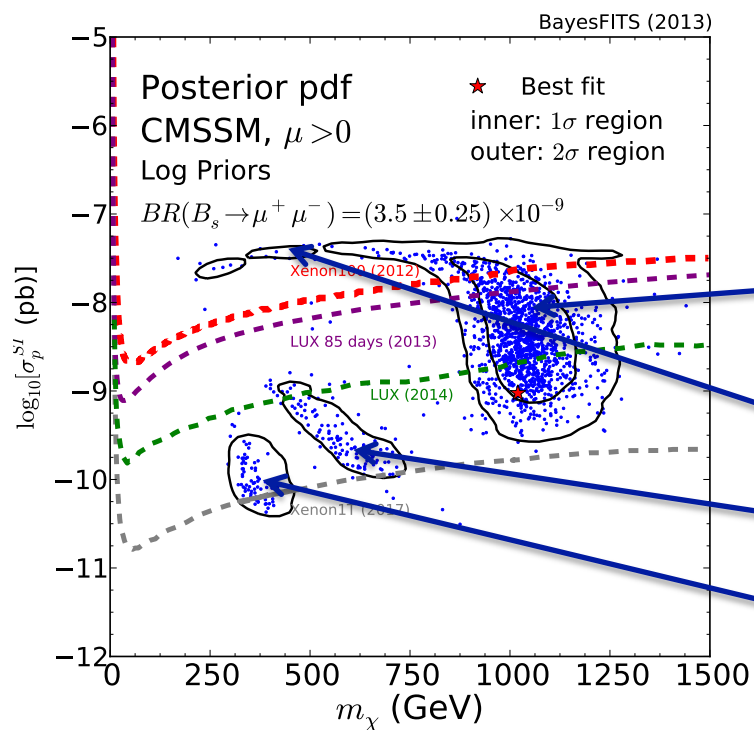
Are we done with the LHC?



LHC14 reach:
Glino: ~2.7 GeV
Squarks: ~3 TeV

**LHC – only stau coannihilation
 will be +/- covered
 ...signal “not guaranteed”**

CMSSM and 1-tonne DM detectors



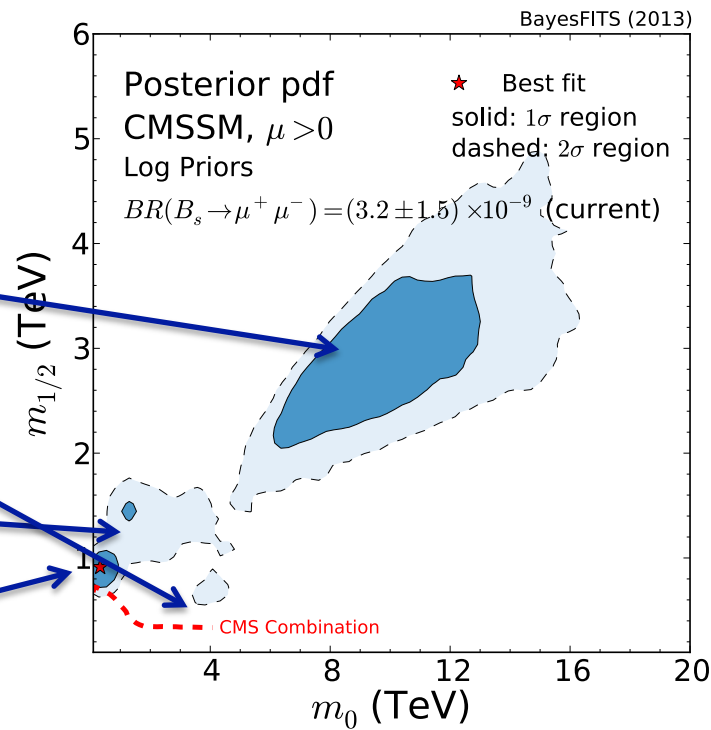
$\mu > 0$

~1 TeV
higgsino LSP

FP/HB

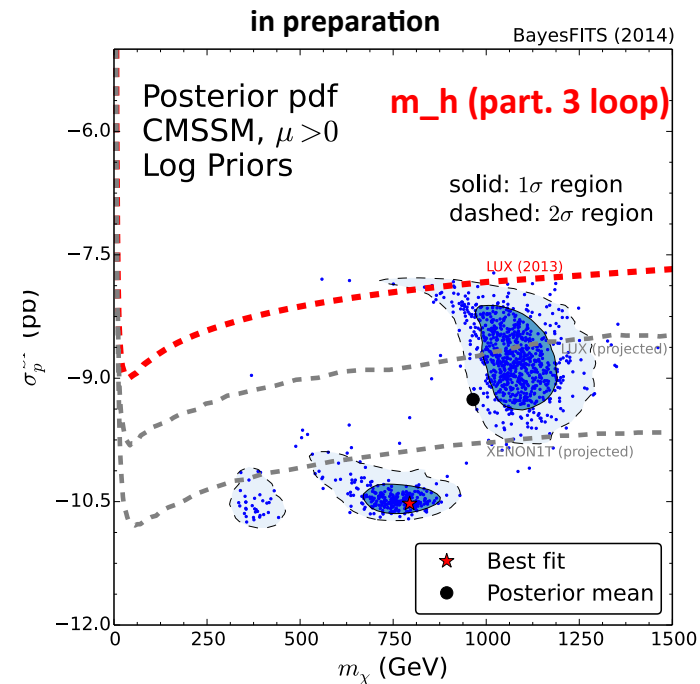
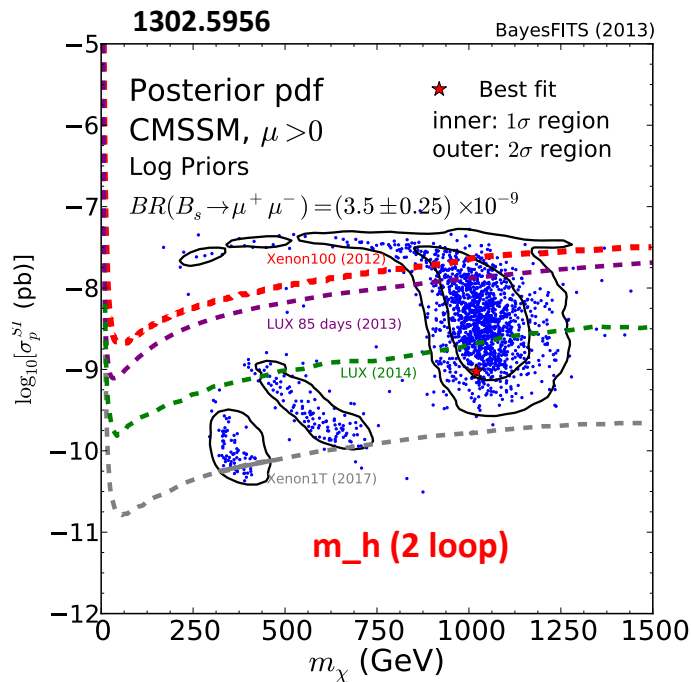
A-funnel

Stau coan'n



Focus point region ruled out by LUX (tension with X100)

~1TeV higgsino DM: exiting prospects for LUX, X100 and 1t detectors



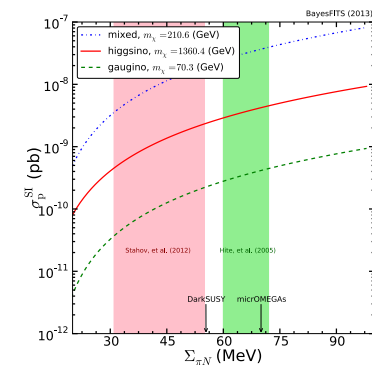
- LUX limit: FP region practically excluded
- Theory sigma_p down by ~1 order of mag

$$\sigma_{\pi N} = 34 \pm 2 \text{ MeV}$$

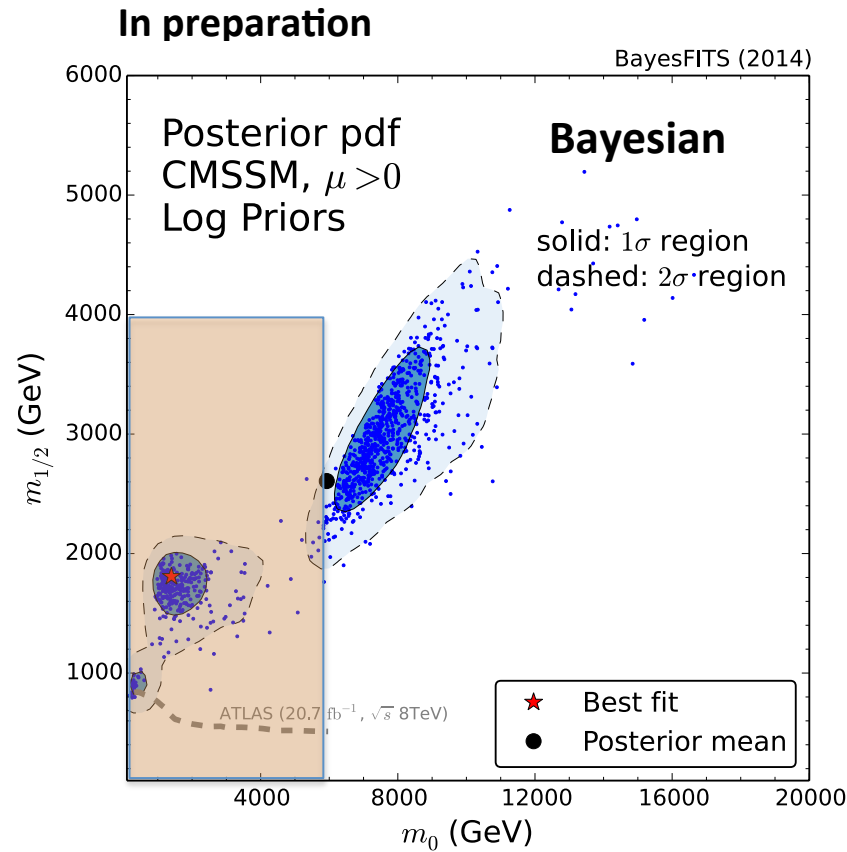
Recent (micrOmegas3.1):

$$\sigma_s = 42 \pm 5 \text{ MeV}$$

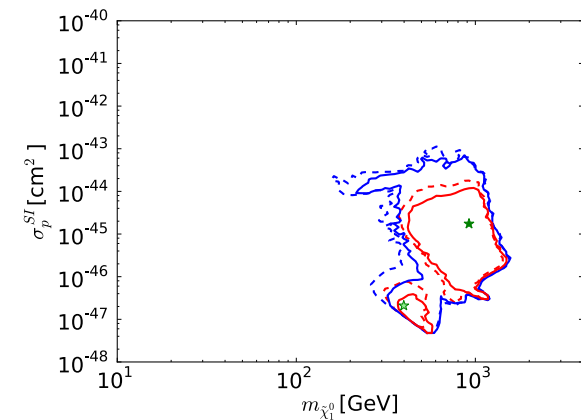
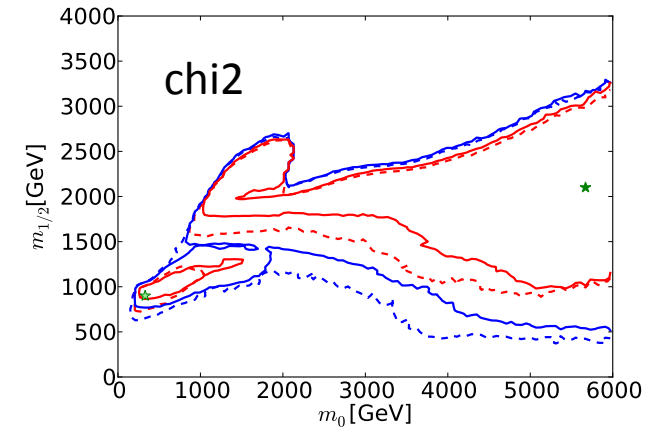
- Effect of 3 loop corr's to m_h: not very significant



Bayesian vs chi-square analysis (updated to include 3loop Higgs mass corrs)



Buchmueller et al [1312.5250](#)

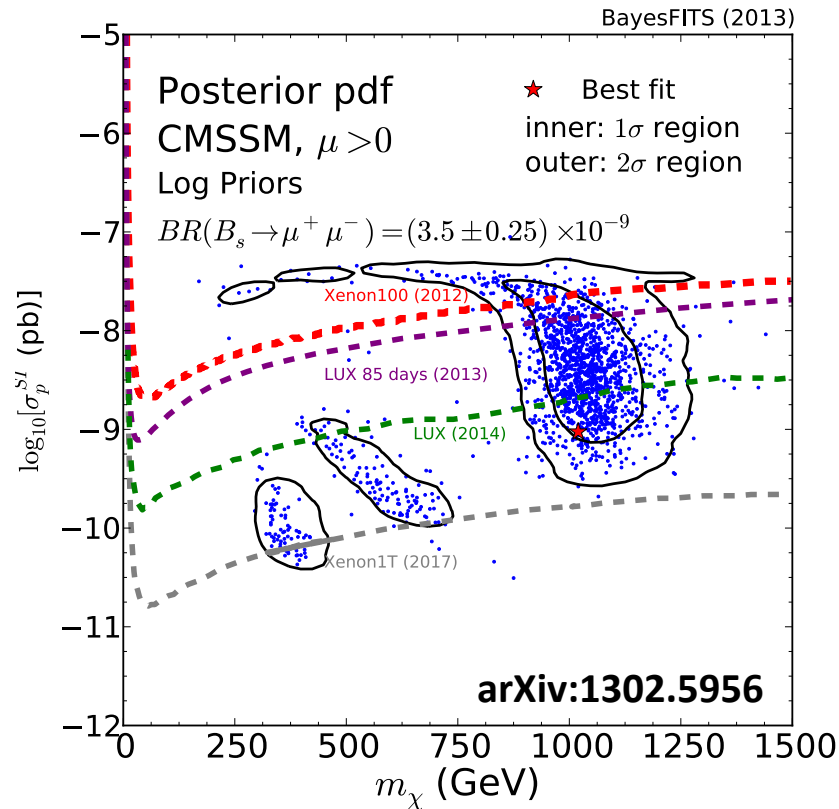


Reasonably good agreement in overlapping region

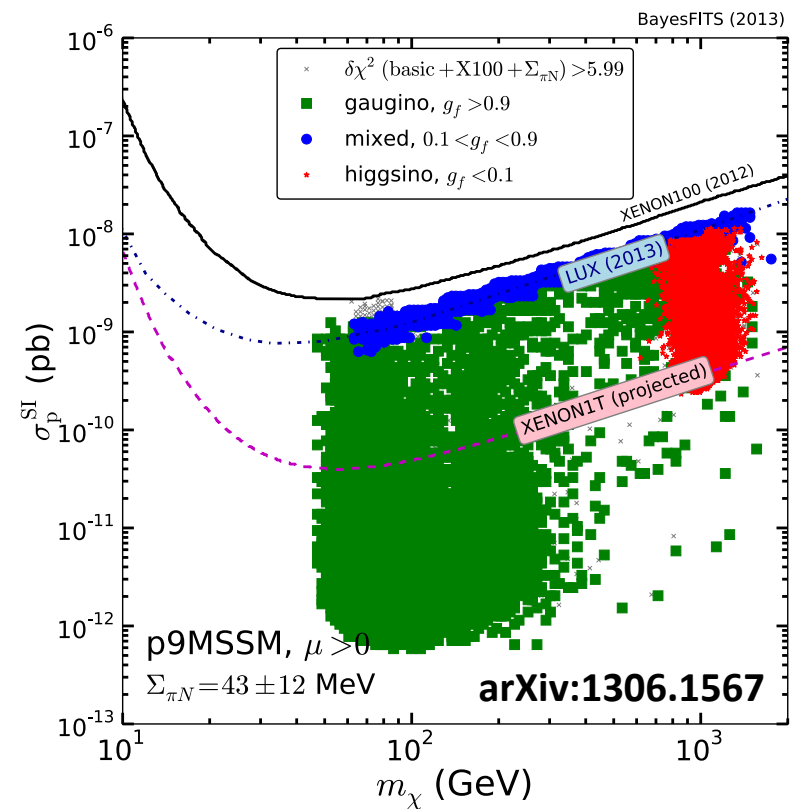
~1 TeV higgsino-like WIMP: implied by ~126 GeV Higgs -> large $m_{1/2}$ and m_0

Unified vs pheno SUSY

Unified SUSY (Constrained MSSM)



General SUSY (p9MSSM)



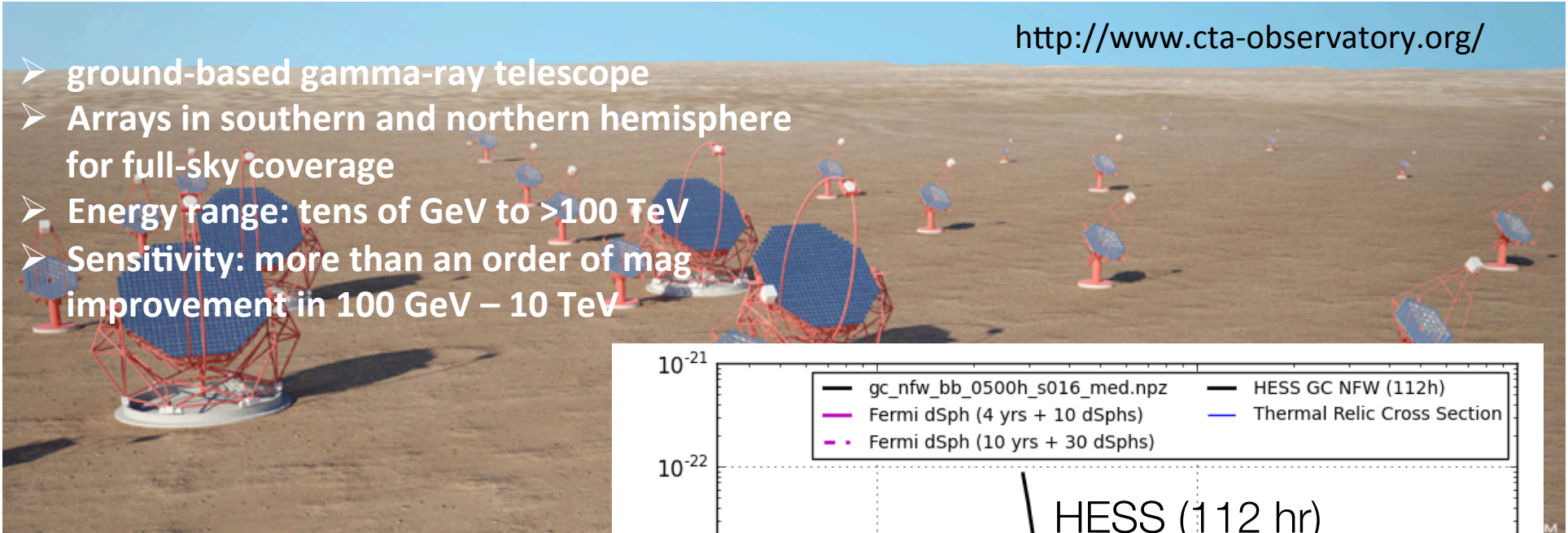
MSSM:

- much bigger ranges allowed
- ~1 TeV higgsino DM: prospects for detection similar to unified SUSY
- New LUX limit -> starts to exclude mixed (bino-higgsino) neutralino

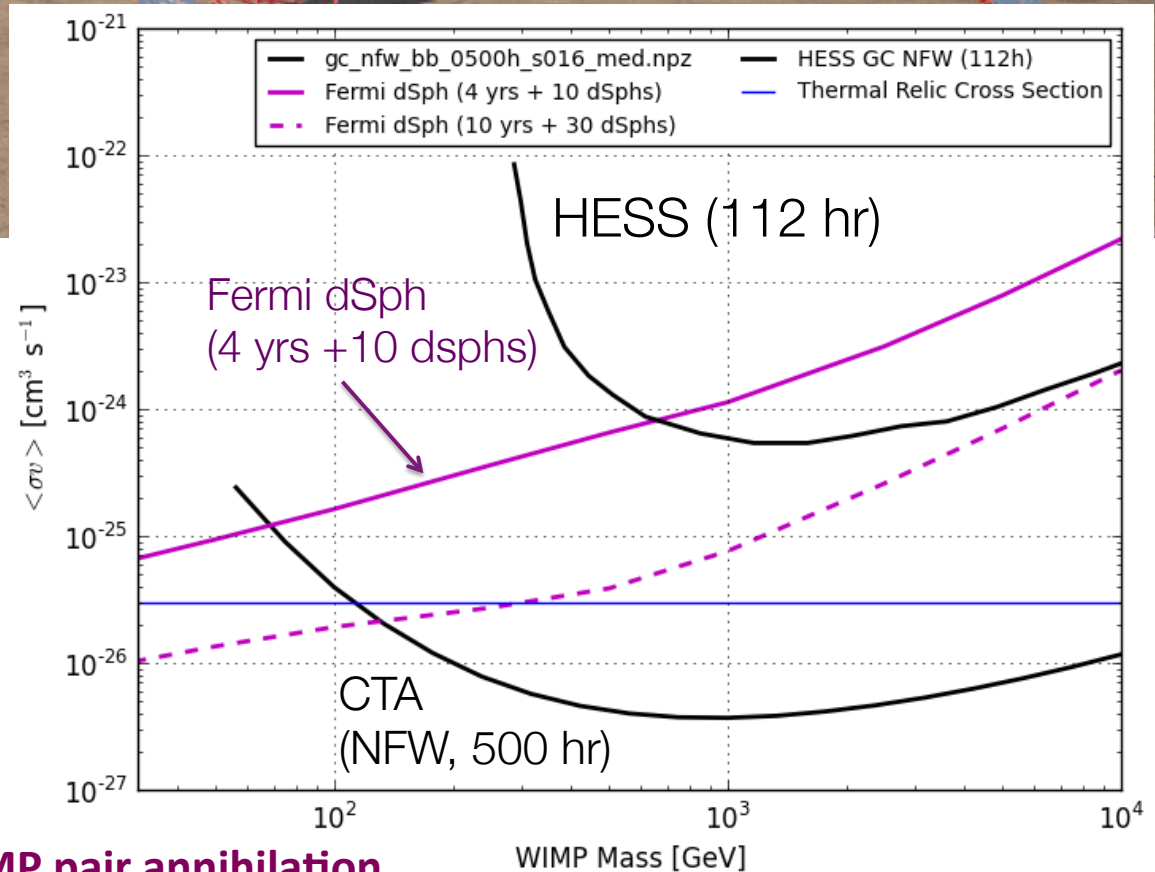
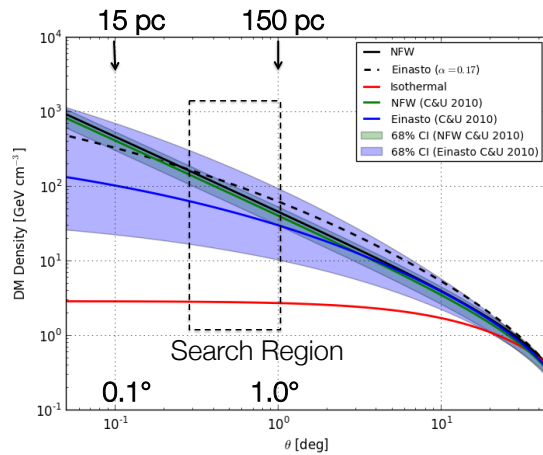
CTA – New guy in DM hunt race

<http://www.cta-observatory.org/>

- ground-based gamma-ray telescope
- Arrays in southern and northern hemisphere for full-sky coverage
- Energy range: tens of GeV to >100 TeV
- Sensitivity: more than an order of mag improvement in 100 GeV – 10 TeV



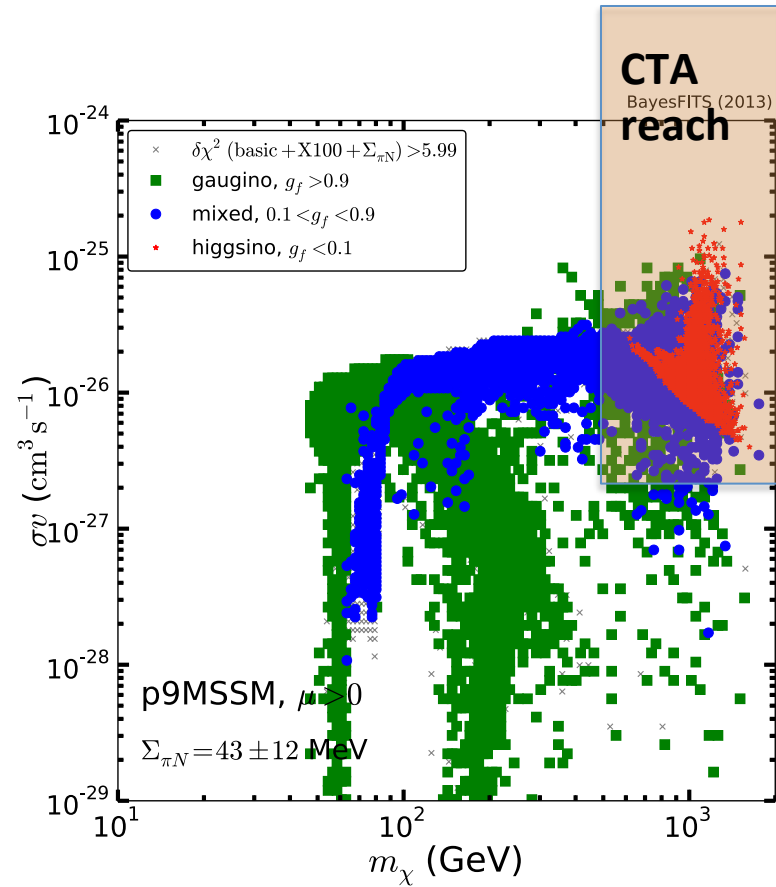
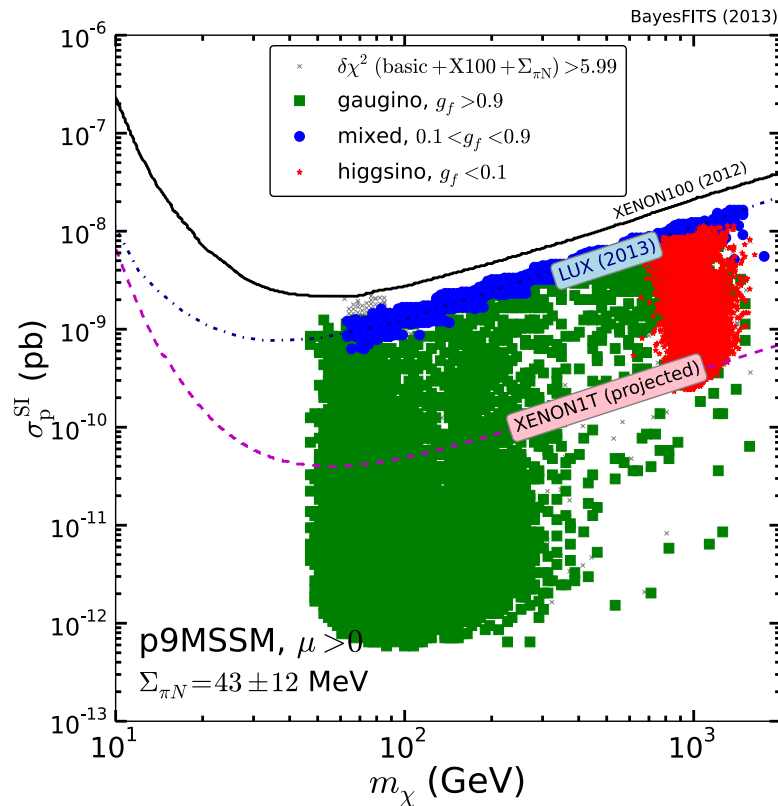
Galactic Center DM Halo



diffuse gamma radiation from WIMP pair annihilation

CTA and SUSY DM

Direct Detection



General SUSY (p9MSSM)

arXiv:1306.1567

MSSM:

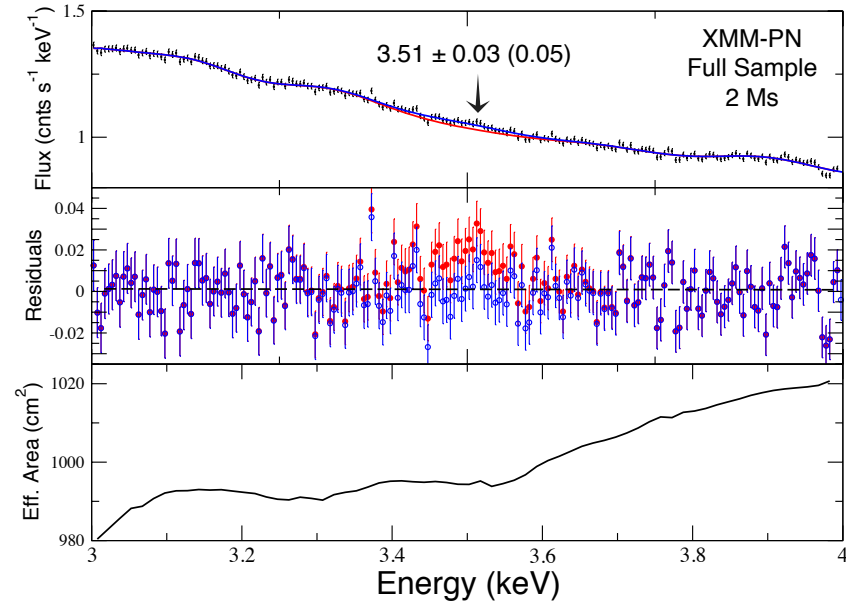
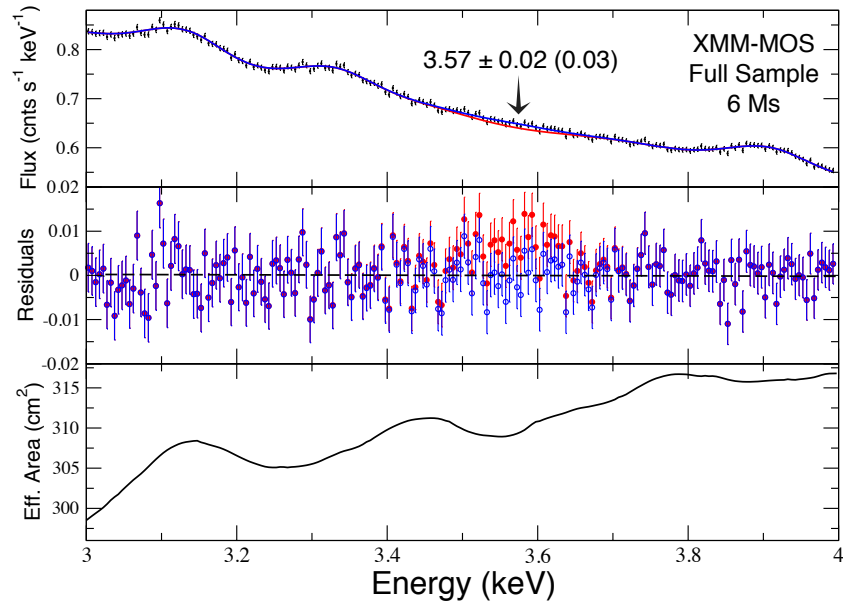
- CTA to probe large WIMP masses
- ~1 TeV higgsino DM: to be completely covered by DD and CTA

X-Ray Signal of DM?

➤ **3.5 keV line is claimed to be seen in clusters of galaxies and in M31**

Bulbul, et al., 1402.2301
Boyarsky, et al., 1402.4119

(XMM data)



Combined data significance 4.4sigma

Lots of theoretical speculations:

- Sterile neutrino decaying into an active one + photon
- Sterile ν \rightarrow axino
- Sterile ν \rightarrow axion-like particle
- ...

Gazing into a crystal ball..





The background of the slide is a deep space image. It features a dense field of stars in various colors, including blue, purple, and white. A prominent, bright yellow star is located near the center. To the left, there is a large, diffuse purple nebula. The overall scene is dark, with the stars and nebula providing the primary light source.

**We need a genuine
WIMP signal...**

**... from more than one
DM search experiment**

To take home:

- **DM: jury is still out, discovery claims come and go...**
- **SUSY neutralino remains most popular choice**
- **Higgs of 126 GeV \rightarrow \sim 1TeV (higgsino) DM – generic prediction of simple, generic SUSY:**
 - To be probed by 1-tonne DM detectors
 - Big bite by LUX already in 2014
 - Independent probe by CTA
 - Far beyond direct LHC reach

Smoking gun of SUSY!?

SUSY may be too heavy for the LHC

DM searches may hopefully come to the rescue

- **3.5 keV X-ray line as DM signal? Too early to tell...**

