



Study of Formation and Evolution double PBHs in the early universe

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NRNU Mephi, M21-115
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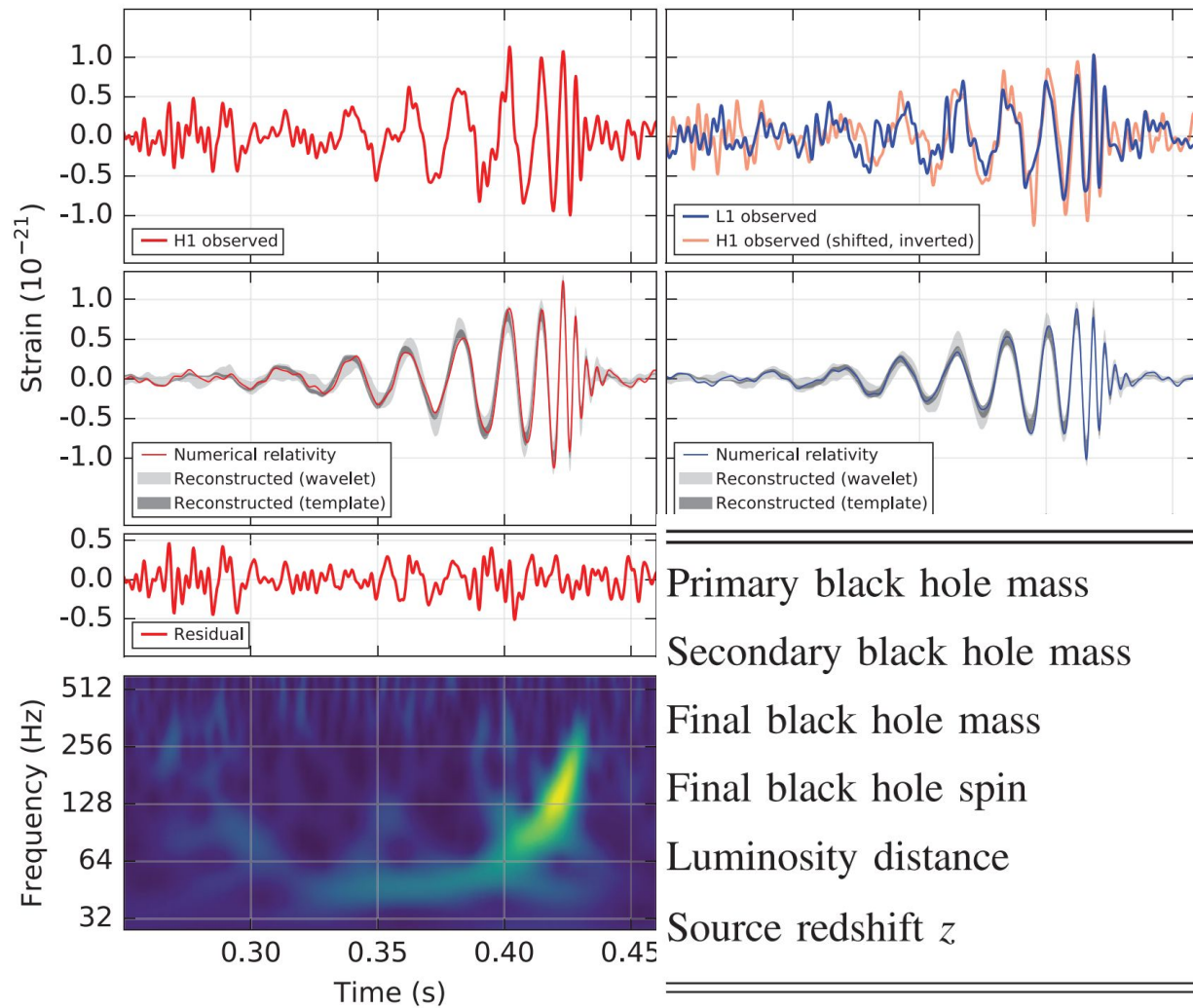
Hanford, Washington (H1)

Livingston, Louisiana (L1)

GW150914

[arXiv:1602.03837](https://arxiv.org/abs/1602.03837) [Observation of Gravitational Waves from a Binary Black Hole Merger LIGO](#)

[B. P. Abbott et al. \(LIGO Scientific Collaboration and Virgo Collaboration\) "Properties of the Binary Black Hole Merger GW150914" 2016](#)



Primary black hole mass

$$36^{+5}_{-4} M_{\odot}$$

Secondary black hole mass

$$29^{+4}_{-4} M_{\odot}$$

Final black hole mass

$$62^{+4}_{-4} M_{\odot}$$

Final black hole spin

$$0.67^{+0.05}_{-0.07}$$

Luminosity distance

$$410^{+160}_{-180} \text{ Mpc}$$

Source redshift z

$$0.09^{+0.03}_{-0.04}$$

Primordial black holes (PBH)

DID BLACK HOLES FORM IMMEDIATELY AFTER THE BIG BANG?

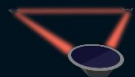


$$M \sim \frac{c^3 t}{G} \sim 10^{15} \left(\frac{t}{10^{-23} \text{ s}} \right) \text{ g}$$

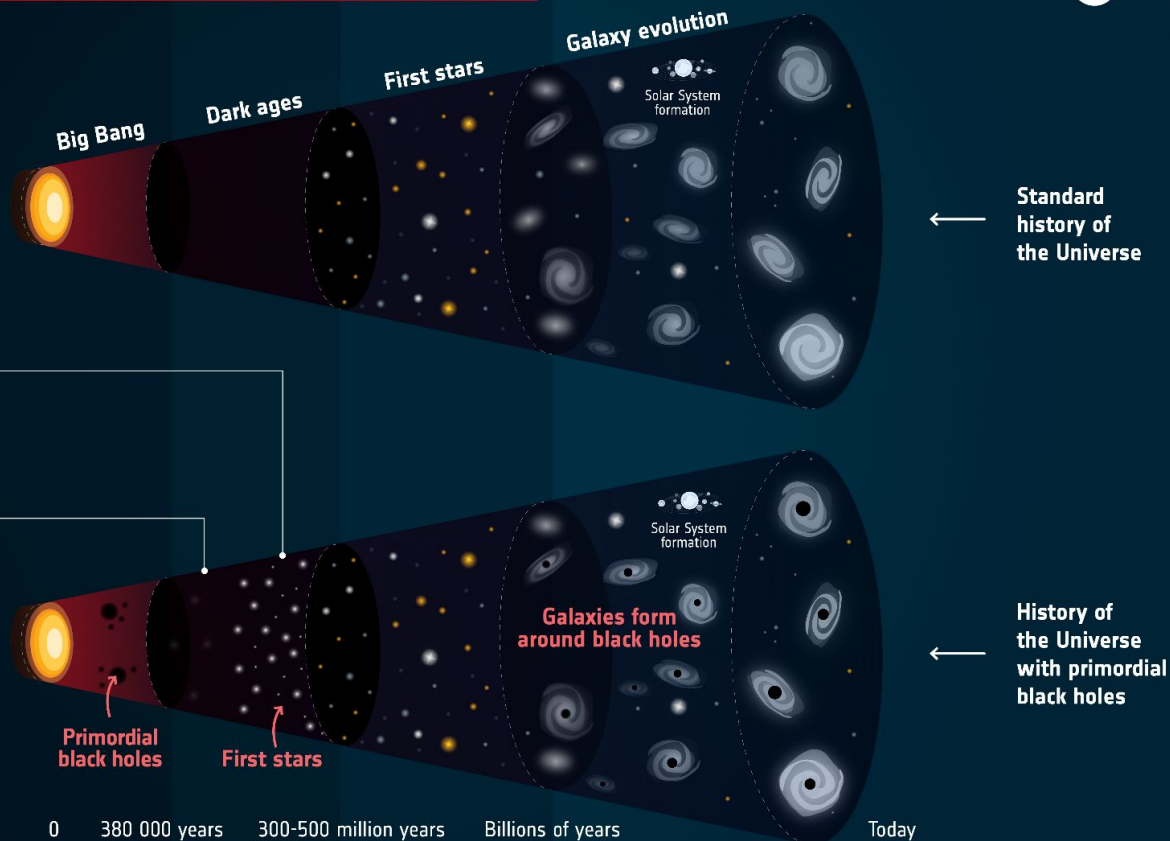
$t \sim 10^{-4} \text{ s}$
for $M = 30 M_{\odot}$

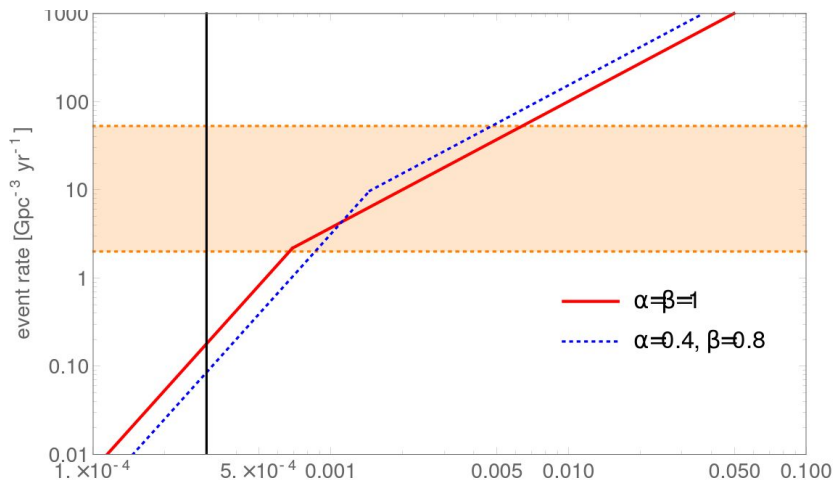


Webb might observe stars that were formed during the dark ages near **primordial black holes**



LISA might pick up gravitational waves from **merging black holes** in the early Universe





[Misao Sasaki "Primordial Black Hole Scenario for the Gravitational-Wave Event GW150914"](#)

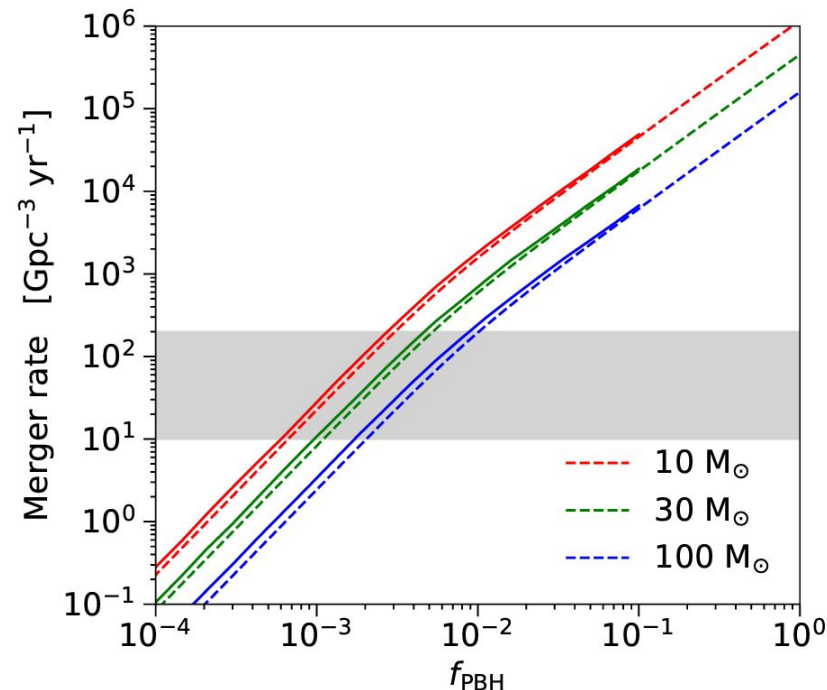
Merger rate

Number of GW signals per year and Gpc^3 .

Shaded bar - observed merger rate measured by LIGO

Black line - the upper limit on rate from the nondetection of the CMB spectral distortion

The *solid lines* on the right graph are the merger rate of PBHs in the *halo of dark matter particles*, the *dotted lines* are without a halo.



[Bradley J. Kavanagh "Black Holes' Dark Dress: On the merger rate of a subdominant population of primordial black holes"](#)

P

PyGra

Project ID: 43381994

Programm realisation

☆ Star0

20 Commits

4 Branches

1 Tag

4.5 MB Project Storage

1 Release



Merge branch 'pygra_dev' into 'main'

...

Gregory Vorobyev authored 2 days ago

37b770c 

main

pygra

Find file



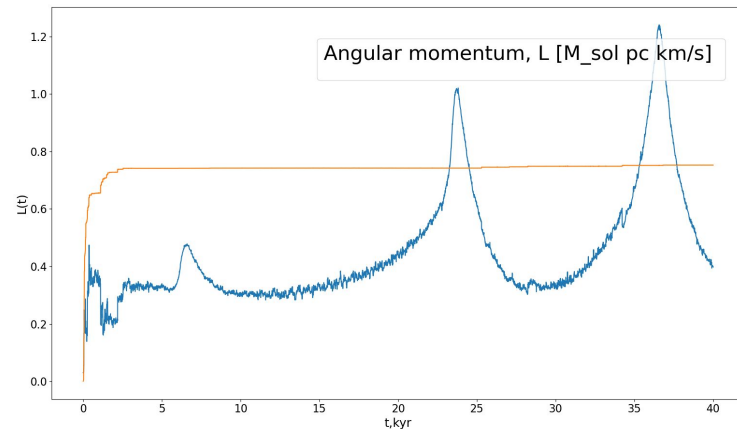
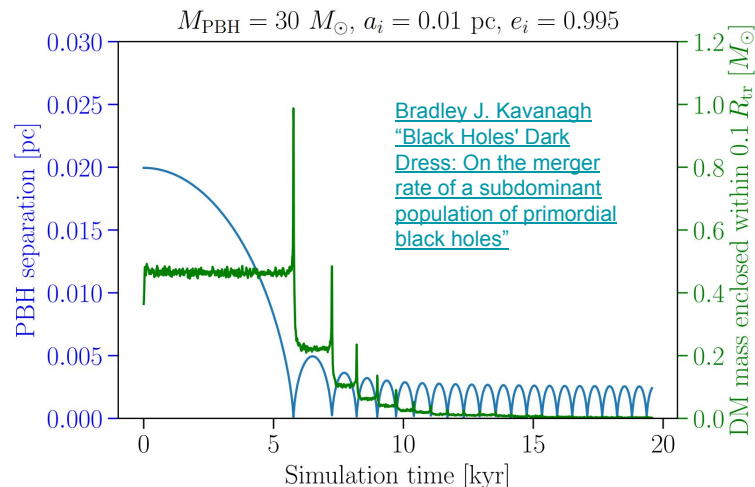
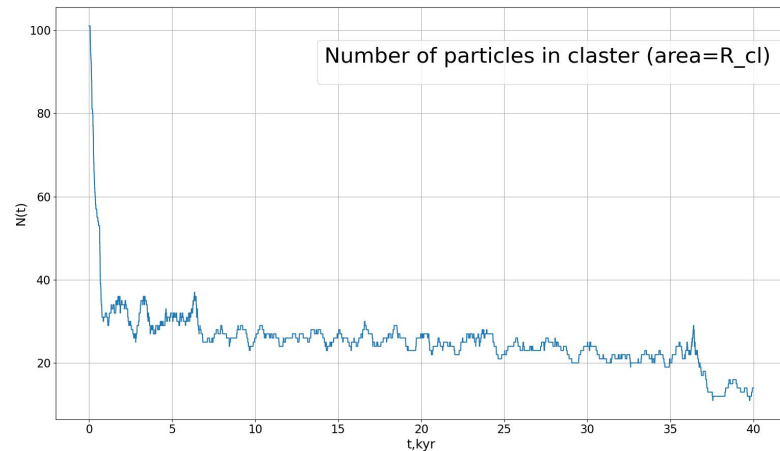
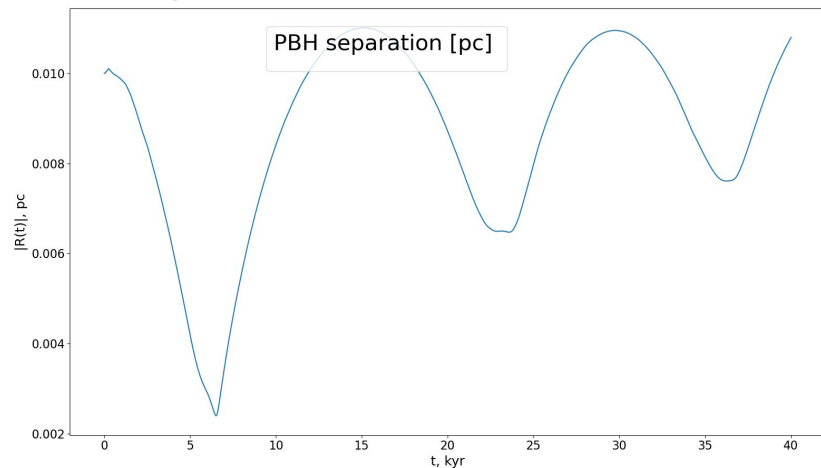
Clone

README



[G.Vorobyev
gitlab.com/MirumeYato/pygra](https://gitlab.com/MirumeYato/pygra)

Merger of binaries surrounded by clusters



Conclusion

In this presentation, we discussed issues of modern cosmology, the successes of gravitational-wave astronomy and the concept of primary black holes (PBHs). The gravitational wave signal GW150914 and hypotheses about the nature of the black holes that gave rise to it are considered. One of the hypotheses noted the possibility of the existence of PBH clusters with central massive BHs at the epoch of the RD-MD transition, the merging of which will also affect the rate of mergers observed by the experiments of the LIGO, Virgo, and KaGra collaborations.

A Python script for N-body simulations was developed. Clusters simulations was provided and next step will be estimating of merger rate for cluster coalescence.

As result with this work we will understand more about nature of binaries systems and maybe give more gravitational wave samples for collaborations like LIGO, Virgo and KaGra.

Thank you for your attention