# Extending the frontiers of Particle Physics

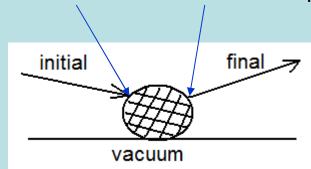
Lecture from the course « Introduction to cosmoparticle physics »

## Elementary Particles

- •Classical mechanics, Quantum mechanics an elementary particle is eternal
- •The problems with *Energy* and *Angular momentum* conservation in  $\beta$ -decay

(1930) – electron and neutrino are created in this process

=> annihilation and creation of particles



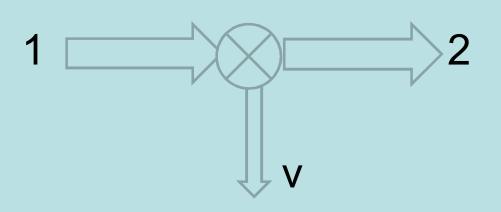
=> New formalism – Quantum field theory

Note: in 1922 A.Friedman discovered non-stationary solutions of Einstein equations for the Universe, proved in the end of 20-s by E.Hubble. Static Universe and its eternal constituents were changed by the modern concepts of particles and expanding Universe nearly simultaneously.

#### Particle Interactions

 Charge of an eternal particle converts in QFT in transition in which initial particle annihilates and the final particle is created with emission or absorption of v quantum of the field of interaction (in QED – photon)

$$e\phi = e|\Psi|^2 \phi = e\Psi_2^*\Psi_1 \phi$$



Elementary act of particle interaction, mediated by boson v, in which initial particle 1 converts into final particle 2

# The Standard Model (SM) of Particle Physics

Gauge principle (Quantum Electrodynamics (QED)):

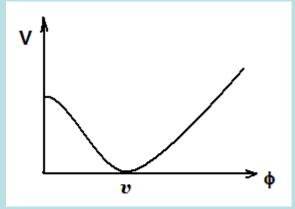
$$L = i\,\overline{\psi}\widehat{\partial}\,\psi + eA_{\mu}J_{\mu} - \frac{1}{4}F_{\mu\nu}F^{\mu\nu}, \qquad J_{\mu} = \overline{\psi}\gamma_{\mu}\psi, \quad F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$$
 
$$\psi \to \psi \exp(ie\chi(x)) \qquad \qquad \text{U(1)-symmetry transformation}$$
 
$$A_{\mu} \to A_{\mu} + \partial_{\mu}\chi$$

- •Generalization on weak and strong interactions:  $SU(3)_c xSU(2)_L xU(1)_Y = SM$
- Higgs mechanism of spontaneous symmetry breaking

$$m_f = \frac{fv}{\sqrt{2}}, \qquad m_W = \frac{fv}{2}, \qquad m_H = \sqrt{2\lambda}v$$

=>in case of *global* symmetry breaking, a Nambu-Goldstone boson *α* appears which interacts as

$$L_{
m int} \propto rac{1}{v} igl( \partial_{\mu} lpha igr) J_{\mu}$$



#### Problems of SM (1) $SU(3)_c \times SU(2)_L \times U(1)_Y$

- 1) Excessive number of parameters:  $g, g', v, \lambda, \{f_f\},...$
- 2) Problem of number of generations

$$\begin{pmatrix} v_e \\ e \end{pmatrix} \begin{pmatrix} v_\mu \\ \mu \end{pmatrix} \begin{pmatrix} v_\tau \\ \tau \end{pmatrix}$$

$$\begin{pmatrix} u \\ d \end{pmatrix} \begin{pmatrix} c \\ s \end{pmatrix} \begin{pmatrix} t \\ b \end{pmatrix}$$

3) The problem of neutrino mass.

$$m(v_e)$$
 < 1 eV <<  $m(e)$ =511 keV

4) CKM mixing matrix

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} 3 & \times & 3 \\ 3 & & \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

5) The problem of C-, P- and CP-violation observed (in weak interactions and K-mesons).

#### Problems of SM (2) $SU(3)_c \times SU(2)_L \times U(1)_Y$

6) Problem of CP-violation in QCD

$$\Delta L_{\rm QCD} = \frac{\alpha_s^2}{16\pi} \theta \cdot \varepsilon_{\alpha\beta\mu\nu} G^{a\alpha\beta} G^{a\mu\nu}$$
 Strong CP-violation

$$d_n \approx e\theta m_u / m_n^2 \sim \theta \cdot 10^{-16} \, e \cdot \text{cm}$$
 From experiment one has  $|\theta| < 10^{-9}$ 

7) Divergence of Higgs boson mass (due to radiation corrections), and, as consequence, instability of Higgs potential against radiation corrections

$$-\frac{H}{\bar{f}} - \frac{H}{-due \text{ to }} - \frac{\Lambda^2}{g_f \varphi f f}.$$

$$-\frac{1}{\bar{f}} - \frac{H}{-due \text{ to }} - \frac{\Lambda^2}{2}$$

$$-\frac{1}{\bar{f}} - \frac{1}{\bar{f}} - \frac{1$$

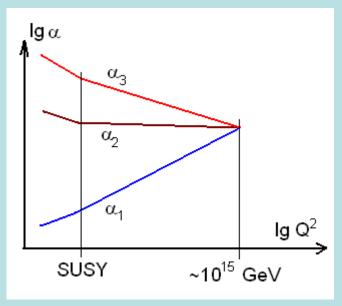
## Other arguments for extension of SM SU(3)<sub>c</sub>xSU(2)<sub>L</sub>xU(1)<sub>Y</sub>

#### Unifying all interactions:

for aesthetical reasons

for reasons of similarity in their descriptions:  $L\sim gA_{\mu}J_{\mu}$ 

convergence of gauge constants (SUSY changes the lines and gives it naturally)



## Trends in extension of SM (1)

"See-saw" mechanism of generation of neutrino mass

$$\Delta L_m = m_D \left( \overline{v}_L v_R + \overline{v}_R v_L \right) + m_R \left( \overline{v}_R^c v_R + \overline{v}_R v_R^c \right) + m_L \left( \overline{v}_L^c v_L + \overline{v}_L v_L^c \right)$$

$$\Delta L_m = m_1 \left( \overline{v}_1^c v_1 + \overline{v}_1 v_1^c \right) + m_2 \left( \overline{v}_2^c v_2 + \overline{v}_2 v_2^c \right)$$

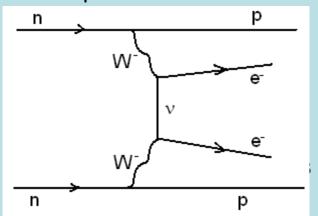
$$m_1 \approx \frac{m_D^2}{m_R} << m_D \sim m_l, \qquad m_2 \approx m_R$$

$$To avoid problems with SM parameters$$

Majorana mass term violates lepton number, which will lead to processes with as  $\Delta L=2$ .

Consequence: neutrinoless double β-decay

Spontaneous breaking of symmetry of lepton charge implies the existence of Nambu-Goldstone boson – Majoron.



#### Trends in extension of SM (2)

#### "Mirror" world

Kobzarev, Okun, Pomeranchuk (1966)

$$(O) \equiv \begin{pmatrix} v \\ e \\ p \\ n \end{pmatrix}_{O}$$

$$(M) \equiv \begin{pmatrix} v \\ e \\ p \\ n \end{pmatrix}_{M}$$

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The equivalence between left- and right-handed coordinate systems is restored, if reflection in mirror is accompanied by change of ordinary particles by their mirror partners. Mirror partners are strictly symmetric to ordinary particles. Therefore they can not have ordinary electromagnetic and strong interactions (doubling of atomic levels, or pion states). Successive analysis have shown that (O) and (M) also can not share W and Z boson mediated weak interaction.

## Trends in extension of SM (3)

SU(5) GUT model

$$SU(5)\supset SU(3)_c xSU(2)_L xU(1)_Y$$

- •There are 24 gauge bosons, which contain 12 known ones (g,Z,W,γ) and 12 new ones (X,Y)
- All known fermions (15, excluding  $v_R$ ) are distributed into SU(5) 5-plets and 10plets.
- Λ<sub>GUT</sub>~10<sup>15</sup> GeV

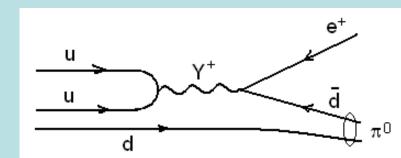
Example of 5-plet: 
$$\left( \overline{d}_{red} \ \ \overline{d}_{yellow} \ \overline{d}_{blue} \ \ e^- \ \ v_e \right)_{\!\! L}$$

Proton decays

$$au_p^{\text{theor}} \sim \frac{\Lambda_{\text{GUT}}^5}{m_p^4} \sim 10^{31} \, \text{years}$$
 vs  $au_{p \to e\pi}^{\text{exp}} > 10^{33} \, \text{years}$ 

$$\tau_{p \to e\pi}^{\text{exp}} > 10^{33} \,\text{years}$$

It strongly restricts the model.



Note:  $\Lambda B = \Lambda L$ 

## Trends in extension of SM (4)

#### Symmetry of families and Familons

To resolve **the problem of quark-lepton families** a new global symmetry  $SU(3)_H$  can be introduced. It's spontaneous breaking leads to 8 new scalar fields Nambu-Goldstone bosons - *familions*.

Familions interact with quarks and leptons at energy scale  $v_a$  as

$$L_{aff} = ic_{aff} \frac{m_f}{v_a} a \overline{f} \gamma_5 f$$

Their interactions lead to various transitions between fermions of different families.

In the model of singlet familion there is only 1 such Nambu-Goldstone boson.

## Trends in extension of SM (5)

#### Model of **Peccei-Quinn**

To resolve the problem of strong CP-violation a new global symmetry  $U(1)_{PQ}$  with new scalar field is introduced.

Scalar field is associated with spontaneous U(1)<sub>PQ</sub> violation at energy scale  $v_a$ , and provides mechanism of dynamical suppression of  $\theta$  (due to its Nambu-Goldstone boson – axion).

Axion interacts with quarks as

$$L_{aff} = ic_{aff} \frac{m_f}{v_a} a \overline{f} \gamma_5 f$$

and have potential

$$V_{\rm eff} \sim m_u \Lambda_{\rm QCD}^3 \left( 1 - \cos(\theta + a / v_a) \right)$$

minimizing at  $a = \langle a \rangle = -\theta v_a$ 

$$\Delta L_{\rm QCD} = \frac{\alpha_s^2}{16\pi} \left( \theta + \frac{a}{v_a} \right) \cdot \varepsilon_{\alpha\beta\mu\nu} G^{a\alpha\beta} G^{a\mu\nu} \to 0$$

## Trends in extension of SM (5)

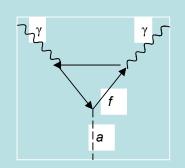
#### Model of **Peccei-Quinn: properties of axion**

The axion has the mass, which is related with its energy scale  $10^{10} C_{2}V$ 

$$m_a \approx 0.6 \,\mathrm{meV} \frac{10^{10} \,\mathrm{GeV}}{v_a}$$

It decays into two photons at

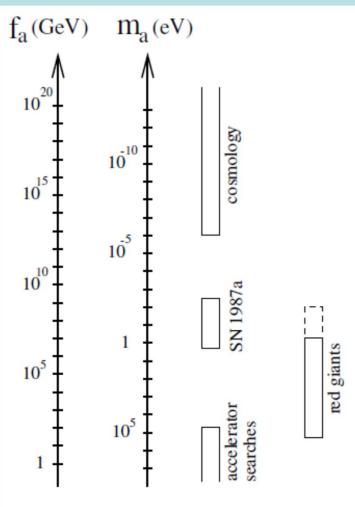
$$\tau(a \to \gamma \gamma) = \frac{64\pi^3}{c_{a\gamma\gamma}^2 \alpha^2} \frac{v_a^2}{m_a^3} \approx \frac{2 \cdot 10^{17} \text{ years}}{c_{a\gamma\gamma}^2} \left(\frac{\text{eV}}{m_a}\right)^5 \approx$$
$$\approx \frac{2 \cdot 10^{33} \text{ years}}{c_{a\gamma\gamma}^2} \left(\frac{v_a}{10^{10} \text{ GeV}}\right)^5.$$



According existing constraints its mass and energy scale must be

$$0.5 \cdot 10^{-5} \text{ eV} < m_a < 0.5 \cdot 10^{-2} \text{ eV}$$
  
 $2 \cdot 10^9 \text{ GeV} < v_a < 10^{12} \text{ GeV}$ 

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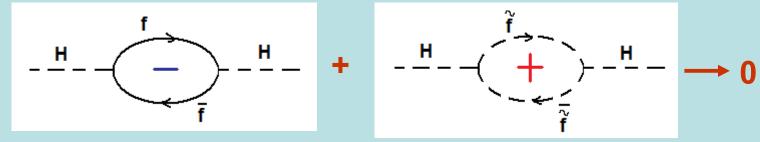
## Trends in extension of SM (6)

#### **Supersymmetry** (SUSY)

All particles with spin s get supersymmetric partner – particle with spin s' = |s-1/2| and all other identical quantum numbers (except the mass in case of broken SUSY).

$$e_{L,R}(s=0) \leftrightarrow \widetilde{e}_{L,R}(s=0)$$
  
 $\gamma(s=1) \leftrightarrow \widetilde{\gamma}(s=1/2)$  Photino (bino) is Majorana particle  
or  $B(s=1) \leftrightarrow \widetilde{B}(s=1/2)$ 

1) SUSY helps to resolve the problem of large radiation correction in Higgs sector.



- 2) SUSY improves convergence of the gauge constants (in GUT).
- 3) SUSY may serve as a basic symmetry for construction of quantum gravity.

**SUSY** must be, evidently, broken. A "hidden" sector is introduced which induces soft SUSY breaking in our sector by means of 1) gravity (**SUGRA**), 2) special gauge fields (**GMSB**). <sup>14</sup>

#### Trends in extension of SM (6)

#### **Supersymmetry** (SUSY)

To avoid fast decay of proton in GUT, a multiplicative quantum number -R-parity is introduced. R=+1 is assigned for all ordinary particles, and R=-1 - for SUSY partners.

Consequence: the lightest SUSY particles (LSP) is stable.

**Neutralino** is often suggested for a role of LSP. Neutralino is a superposition of four SUSY partners of bino, neutral wino, two neutral higgsino.

$$\chi_{1,2,3,4}^{0} = N_{1,2,3,4}^{(B)} \widetilde{B} + N_{1,2,3,4}^{(W)} \widetilde{W}^{3} + N_{1,2,3,4}^{(H_{1})} \widetilde{H}_{1}^{0} + N_{1,2,3,4}^{(H_{2})} \widetilde{H}_{2}^{0}$$

Aslo, gravitino or axino are sometimes suggested for a role of LSP.

#### Trends in extension of SM (7)

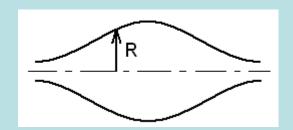
Kaluza-Klein (KK) models

$$g_{MN} = \begin{pmatrix} g_{\mu\nu} & g_{\mu n} \\ g_{m\nu} & g_{mn} \end{pmatrix}$$

$$p_{NN=0,...D}$$

$$m, n=4,...D$$
Describes other interactions

If extra d dimensions are compactified with R=const, then they are manifested in form of the KK-mass states for all particles, if R $\neq$ const – in form of interaction.



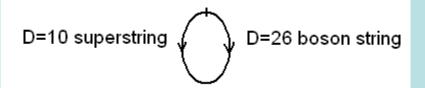
## Trends in extension of SM (9)

#### Superstring theory, M-theory

Each species of particles is a certain vibration mode of a string – fundamental element of zero thickness and of Planckian length. Strings may be closed and open.

Most popular string model is a so called *heterotic string* (closed) – combines the

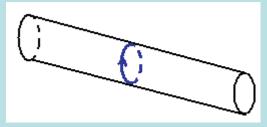
models of superstring and of bosonic string.



Fixed gauge groups:  $E_8xE_8$ ', SO(32)

Each contains 248 bosons and 248 fermion fields

Possible consequence: existence of homotopically stable particle



Superstring U supergravity ⊂ M-theory (?)

D=10

**D**≤11

#### Conclusions

- The Standard model contains internal inconsistencies and is not complete. The first experimental evidence of its incompleteness gives nonzero mass of neutrino.
- The problems of the Standard Model imply its extensions for their solution.
- There are only few experimental probes for these extensions, most of which involve physics of very high energy.
- One should extend the set of such probes by astrophysical and cosmological tests.
- We'll see in the successive lectures that the modern cosmology, in its turn, involves physics beyond the Standard model.
- Mutual relationship of problems of frontiers of physics at extreme microscopic and macroscopic scales is the subject of cosmoparticle physics

#### Basic ideas of cosmoparticle physics

- Physics beyond the Standard model can be studied in combination of indirect physical, astrophysical and cosmological effects
- New symmetries imply new conserved charges. Strictly conserved charge implies stability of the lightest particle, possessing it.
- New stable particles should be present in the Universe. Breaking of new symmetries implies cosmological phase transitions. Cosmological and astrophysical constraints are supplementary to direct experimental search and probe the fundamental structure of particle theory at the scale V
- Combination of physical, cosmological and astrophysical effects provide an over-determined system of equations for parameters of particle theory

